

# Sparticle mass reconstruction in supersymmetry with long-lived staus

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Outline	mSUGRA parameters	Basic strategy and results 00000000 0000000 000	Conclusion

# List of publications

- Neutralino reconstruction in supersymmetry with long-lived staus, SB and Biswarup Mukhopadhyaya. Published in Phys. Rev. D79:115009, (2009), arXiv:0902.4349 [hep-ph]
- Chargino reconstruction in supersymmetry with long-lived staus, SB and Biswarup Mukhopadhyaya. Published in Phys. Rev. D81:015003, (2010), arXiv:0910.3446 [hep-ph]
- Reconstruction of the left-chiral tau-sneutrino in supersymmetry with a right-sneutrino as the lightest supersymmetric particle, SB. Accepted in Phys. Rev. D, arXiv:1002.4395 [hep-ph].

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# Outline

- Introduction
- Choice of mSUGRA parameters
- Basic strategy and results
  - Neutralino reconstruction
  - Chargino reconstruction
  - Left-chiral tau sneutrino reconstruction

# Conclusion

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## Introduction

- Standard Model of particle physics is incomplete...
  - ► The Higgs mass is unstable under the radiative correction
  - It does not explain the observed non-vanishing neutrino masses and mixing
  - It also does not explain the observed DM density and baryon asymmetry of the Universe.
- There are ample reasons to look for physics beyond the Standard Model.

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- ► Physics beyond the Standard Model⇒ new particles.
- ► Supersymmetry: Bosons ↔ Fermions → not an exception.
- Measuring mass and spin of these new particles are of paramount importance.

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# General features of SUSY at the LHC

- ► SUSY production is dominated by gluinos and squarks.
- the gluinos and squarks cascade down, generally in several steps, giving rise to the canonical SUSY signal,

multi-jets and/or leptons and missing transverse energy ( $\not E_T$ ) due to two invisible LSPs

(in a *R*-parity, defined as  $R = (-)^{3B+L+2S}$ , conserving scenario.)

Mass reconstruction is quite a difficult task at the LHC.

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► Mass reconstruction is quite a difficult task at the LHC.

# But this possibility is not unique.

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## Alternative SUSY signal at the LHC

► One can have a charged particle as the NLSP and the decay of it into the LSP is suppressed ⇒ NLSP becomes stable on the length scale of the detector.

 $\Rightarrow$  The essence of SUSY signal lies not in  $\not\!\!\!E_T$ , but in charged track due to massive particle, seen in the muon chamber.

► Additional kinematic information ⇒ opens up the possibility of mass reconstruction

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## Right-handed sneutrino LSP scenario

- The scenario we have studied as an example is a right-handed sneutrino LSP and the NLSP is the superpartner of tau.
- In MSSM one can achieve this by the addition of a right-handed neutrino superfield for each family in the MSSM spectrum, assuming neutrino to be of Dirac type.
- The superpotential in this case is given by

 $W_{MSSM} = y_l \hat{L} \hat{H}_d \hat{E}^c + y_d \hat{Q} \hat{H}_d \hat{D}^c + y_u \hat{Q} \hat{H}_u \hat{U}^c + \mu \hat{H}_d \hat{H}_u + y_\nu \hat{H}_u \hat{L} \hat{\nu}_R^c$ 

 These sneutrinos will have all their interactions proportional to y<sub>v</sub>

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It is possible to accommodate such a scenario in a high-scale framework of SUSY breaking:

$$\begin{array}{c} \underline{\mathsf{mSUGRA}}\\ m_0, \ M_{1/2}, \ A_0, \ sign(\mu), \ \tan\beta \\ \downarrow \end{array}$$

All the low energy parameters

The RGE of the right-chiral sneutrino mass parameter at the one-loop level is given by-

$$rac{dM_{ ilde{
u}_R}^2}{dt}\simeq rac{2}{16\pi^2}y_
u^2\,\, A_
u^2$$

► The small Dirac masses of the neutrinos  $m_{\nu} = y_{\nu} \langle H_u^0 \rangle = y_{\nu} v \sin \beta \Rightarrow y_{\nu} \sim 10^{-13}$ such a small  $y_{\nu} \Rightarrow M_{\tilde{\nu}_R}$  to remain frozen at  $m_0$ 

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▶ The soft SUSY breaking terms contributing to neutrino mass

$$-\mathcal{L}_{soft} \sim M_{\tilde{L}}^2 \tilde{L}^{\dagger} \tilde{L} + M_{\tilde{\nu}_R}^2 |\tilde{\nu}_R|^2 + (y_{\nu} A_{\nu} H_u. \tilde{L} \tilde{\nu}_R^c + h.c.)$$

The sneutrino mass-squared matrix is of the form

$$m_{\tilde{\nu}}^2 = \begin{pmatrix} M_{\tilde{L}}^2 + \frac{1}{2}m_Z^2\cos 2\beta & y_\nu v(A_\nu \sin \beta - \mu \cos \beta) \\ y_\nu v(A_\nu \sin \beta - \mu \cos \beta) & M_{\tilde{\nu}_R}^2 \end{pmatrix}$$

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► The lighter sneutrino mass eigenstate

$$\tilde{\nu}_1 = -\tilde{\nu}_L \sin\theta + \tilde{\nu}_R \cos\theta$$

#### where the mixing angle $\theta$ is given as

$$an 2 heta = rac{2y_
u v \sineta | \coteta \mu - A_
u|}{m^2_{ ilde
u_L} - m^2_{ ilde
u_R}}$$

which for each family is dominated by the right-chiral state.

Thus one naturally has RH-sneutrino LSP for a wide range of values of the gaugino mass parameter.

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## **mSUGRA** parameters

- ► The low energy mass spectrum is determined from the mSUGRA parameter set (m<sub>0</sub>, M<sub>1/2</sub>, A<sub>0</sub>, sign(μ) and tan β) specified at some high scale value using the Spectrum Generator: ISASUGRA v7.78
- ► We have worked in the region of m<sub>0</sub>-M<sub>1/2</sub> plane where a τ̃ LSP occur in a usual mSUGRA scenario without the right handed sneutrino.
- ► The value of A<sub>0</sub> is fixed at 100 GeV without any loss of generality and sign(µ) has been taken to be +ve.

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Input	BP-1	BP-2	BP-3	BP-4	BP-5	BP-6
	$m_0 = 100$	$m_0 = 100$	$m_0 = 100$	$m_0 = 100$	$m_0 = 100$	$m_0 = 100$
mSUGRA	$m_{1/2} = 600$	$m_{1/2} = 500$	$m_{1/2} = 400$	$m_{1/2} = 350$	$m_{1/2} = 325$	$m_{1/2} = 325$
	$\tan \beta = 30$	aneta=30	aneta=30	aneta=30	aneta=30	aneta=25
m <sub>ṽiR</sub>	100	100	100	100	100	100
$m_{\tilde{e_I}}, \tilde{m_{\tilde{\mu}_I}}$	418	355	292	262	247	247
$m_{\tilde{e}_R}$ , $m_{\tilde{\mu}_R}$	250	214	183	169	162	162
$m_{\tilde{\nu}_{eL}}, m_{\tilde{\nu}_{\mu L}}$	408	343	279	247	232	232
$m_{\tilde{\nu}_{\tau L}}$	395	333	270	239	224	226
$m_{\tilde{\tau}_1}$	189	158	127	112	106	124
$m_{\chi_{1}^{0}}$	248	204	161	140	129	129
$m_{\chi^0_2}^{\chi^1_1}$	469	386	303	261	241	240
$m_{\chi_1^{\pm}}$	470	387	303	262	241	241
m <sub>ĝ</sub>	1362	1151	937	829	774	774
m <sub>ĩ1</sub>	969	816	772	582	634	543
$m_{h^0}$	115	114	112	111	111	111

Table: Proposed benchmark points (BP) for study of stau-NLSP scenario in the SUGRA with right-sneutrino LSP. All the values except  $\tan \beta$  are given in GeV.

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# **Collider simulation**

- ► We have used the event generator Pythia v6.4.16 for our simulation.
- ► *pp* collision has been simulated with a cm energy  $E_{cm} = 14 \ TeV$ .
- CTEQ5L pdf used in the simulation.
- QCD renormalisation and factorisation scales set at the  $\mu_{\mathbf{F}} = \mu_{\mathbf{R}} = \mathbf{m}_{\mathbf{average}}^{\mathbf{final}}$ .
- Hadronisation has been done, using the fragmentation functions inbuilt in Pythia.
- The effects of ISR and FSR and finite detector resolution have been taken into account.



#### Neutralino reconstruction

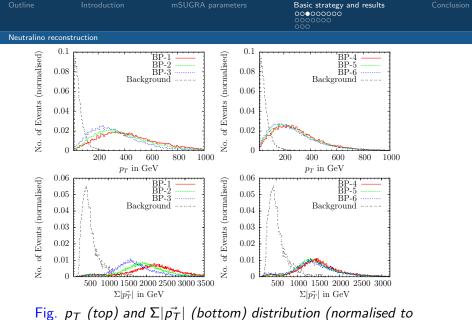
- Neutralino are produced in cascade decay
  - $egin{array}{ll} ilde{q} o q \chi^0_i o q ilde{ au}^* au \ ilde{g} o ilde{q} ilde{q} o ilde{q} q \chi^0_i o ilde{q} q ilde{ au}^* au \end{array}$
  - $\Rightarrow 2\tau_j + 2\tilde{\tau} (\text{charged-track}) + \not E_T + X$
- ► We have not consider any isolated leptons ⇒Minimize the contribution to *F<sub>T</sub>* of neutrinos other than from taus.

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#### **Event selection criteria**

Events have been selected with the following basic criteria  $p_{\tau}^{lep,stau} > 10 \text{ GeV}$  $p_{\tau}^{hardest-jet} > 75 \text{ GeV}$  $p_{\tau}^{other-jets} > 50 \text{ GeV}$  $E_T > 40 \text{ GeV}$  $|\eta| < 2.5$  for leptons, jets & stau  $\Delta R_{II} > 0.2, \ \Delta R_{Ii} > 0.4$  $\Delta R_{\tilde{\tau}l} > 0.2, \ \Delta R_{\tilde{\tau}i} > 0.4$  $\Delta R_{ii} > 0.7$ where,  $\Delta R = \sqrt{(\Delta \eta)^2 + (\Delta \phi)^2}$ 



unity) for the signal and the background for all benchmark points.

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# To reduce the SM background

- $p_T > 100 \text{ GeV}$  for each charged track
- $\Sigma |p_T| > 1 TeV$

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#### $\tau$ -reconstruction

- ► We have selected hadronic jets with E<sub>T</sub> > 50 GeV, assuming the identification efficiency of a true tau-jet to be 50%, and a non-tau jet rejection factor of 20.
- If x<sub>i</sub> is the fraction of the parent τ-energy carried by each product jet, then one can write

$$P_{h_i} = x_i P_{\tau_i} \quad (i = 1, 2)$$

(in the collinear approx.  $E_{ au} >> m_{ au}$ )

$$ec{\mathcal{F}_{T}} = ec{p}_{T}^{
u_{1}} + ec{p}_{T}^{
u_{2}} = (rac{1}{x_{1}} - 1) \ ec{p}_{h1} + (rac{1}{x_{2}} - 1) \ ec{p}_{h2}$$

D. L. Rainwater, D. Zeppenfeld and K. Hagiwara, Phys. Rev. D 59, 014037 (1999)

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• Reconstruction of the neutralinos in the channel  $\chi_1^0 (\chi_2^0) \rightarrow \tilde{\tau} \tau$  requires...

The information about  $\tilde{\tau}$  4-momenta  $\leftarrow$  knowledge of  $m_{\tilde{\tau}}$ 

First we identify a correct ττ̃-pair by using a seed value for m<sub>τ̃</sub> (=100 GeV) and demand

 $|M^{pair1}_{ au au}-M^{pair2}_{ au au}|$  is min. and <50~GeV

Demand for the correct pairs

$$\begin{array}{c} M_{\tau_1\tilde{\tau_1}} = M_{\tau_2\tilde{\tau_2}} \\ \Downarrow \end{array}$$

$$\sqrt{m_{\tilde{\tau}}^2 + |\vec{p}_{\tilde{\tau}_1}|^2} . E_{\tau_1} - \sqrt{m_{\tilde{\tau}}^2 + |\vec{p}_{\tilde{\tau}_2}|^2} . E_{\tau_2} = \vec{p}_{\tilde{\tau}_1} . \vec{p}_{\tau_1} - \vec{p}_{\tilde{\tau}_2} . \vec{p}_{\tau_2}$$

• This method is effective when both pairs of  $\tau \tilde{\tau}$  come from  $\chi_1^0 \chi_1^0$  or  $\chi_2^0 \chi_2^0$ .

#### Neutralino reconstruction

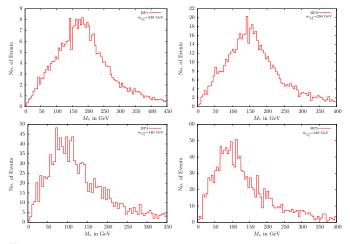


Fig. The  $\tilde{\tau}$  mass peak obtained from eventwise reconstruction of  $m_{\tilde{\tau}}$ .

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#### Neutralino reconstruction

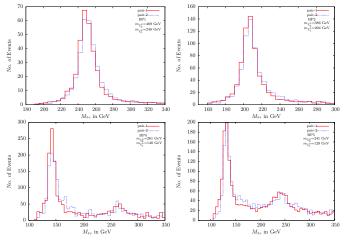


Fig.  $M_{\tilde{\tau}\tau}$  distribution for BP1, BP2, BP4 and BP5.

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CUTS		SIGNAL			BACKGROUND
	BP1	BP2	BP4	BP5	
Basic Cuts	1765	4143	20889	28864	4180
$+P_T$ Cut	1588	3631	15526	20282	214
$+\Sigma  P_T $ Cut	1442	3076	9538	11266	63
$   +  M_{ ilde{ au} au}^{pair1} - M_{ ilde{ au} au}^{pair2} $ Cut	408	887	2004	2244	6

Table: Number of signal and background events for  $2\tau_j + 2\tilde{\tau} + E_T + X$  final state with an integrated luminosity of 300 fb<sup>-1</sup>.

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# **Chargino reconstruction**

- $\chi_1^{\pm}$  is produced in association with  $\chi_1^0/\chi_2^0$  and hard jets.
- The  $\chi_1^{\pm}$  subsequently decays into a  $\tilde{\tau}^{(*)} \nu_{\tau}^{(-)}$  pair, while the  $\chi_1^0$  (or  $\chi_2^0$ ) decays into a  $\tilde{\tau} \tau$  pair.
- ► The final state consists of  $\tau_j + 2\tilde{\tau}(opposite sign charged tracks) + \not E_T + X$
- Since the decay of χ<sup>±</sup><sub>1</sub> involves an invisible particle (ν<sub>τ</sub>), a transverse mass distribution, rather than invariant mass distribution, will give us information of m<sub>χ<sup>±</sup></sub>.

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- The  $\vec{E}_T$  is comprises of  $\vec{E}_T = \vec{P}_{\nu_1}^T + \vec{P}_{\nu_2}^T$
- ► The transverse momenta of the neutrino (P<sup>T</sup><sub>ν2</sub>), out of a tau decay, is first reconstructed in the collinear approximation

$$P_{\tau_j} = x P_{\tau}$$

- ► Following the decay  $\chi_1^0$  (or,  $\chi_2^0$ )  $\rightarrow \tilde{\tau}^{\pm} \tau^{\mp}$  one can write  $m_{\chi_i^0}^2 = (P_{\tilde{\tau}} + P_{\tau})^2 = (P_{\tilde{\tau}} + P_{\tau_j}/x)^2$  (*i* = 1,2)
- One can solve this to obtain

$$\vec{P}_{\nu_2}^{T} = \vec{P}_{\tau}^{T} - \vec{P}_{\tau_j}^{T} = \frac{1-x}{x} \cdot \vec{P}_{\tau_j}^{T}$$



► The transverse momenta of the neutrino, out of \(\chi\_1^\pm-decay\) can be extracted from the knowledge of \(\vec{\vec{E}}\_T\) of that particular event

$$\vec{P}_{\nu_1}^T = \vec{E}_T - \vec{P}_{\nu_2}^T$$

From the transverse mass distribution of the  $\tilde{\tau}$ - $\nu_{\tau}$  pair, defined by,

$$M^T_{ ilde{ au}
u_ au} = \sqrt{(E^T_{ ilde{ au}} + E^T_{
u_ au})^2 - (ec{P}^T_{ ilde{ au}} + ec{P}^T_{
u_ au})^2}$$

one can obtain the value of  $m_{\chi^{\pm}_1}$ 



# SUSY backgrounds

- SUSY processes in this scenario itself are serious than SM backgrounds
- ► The dominant contributions come from

$$\chi_i^0 \chi_j^0 + X \to (\tilde{\tau}\tau)(\tilde{\tau}\tau) + X$$

one of the  $\tau$  not being identified

$$\widetilde{
u}_{ au_L}\chi_i^0 + X 
ightarrow ( ilde{ au}W)( ilde{ au} au) + X \ \downarrow \ ext{ with the }W ext{ not being identified }$$



Cut X: look at the invariant mass distribution of the τ (having same charge as that of the τ) with each jet in the final state and veto that event if

$$m_{\chi^0_i} - 20 < M_{ ilde{ au}j} < m_{\chi^0_i} + 20$$

Cut Y: veto event with W being identified in its hadronic decay if

$$M_W - 20 < M_{jj} < M_W + 20$$
 and  $\Delta R_{ ilde{ au} j_r} < 0.8$ 

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#### Chargino reconstruction

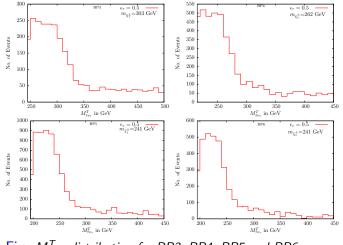


Fig.  $M_{\tilde{\tau}\nu_{\pi}}^{T}$  distribution for BP3, BP4, BP5 and BP6.

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BP4	Signal	SM	SUSY backgrounds	
	$(\chi_{1/2}^0 - \chi_1^{\pm})$	backgrounds	$\chi_i^0 - \chi_j^0$	$\chi^0_{1/2} - \tilde{\nu}_{\tau L}$
Basic cuts	18194	65588	44580	10613
With $p_T + \Sigma  p_T $ Cut	10697	202	31256	5713
Cut X+cut Y	4875	202	9872	1480
$M_{\tilde{\tau}\nu_{\tau}}^{T} > \frac{3}{4}m_{\chi_{2}^{0}}$	2345	1	1754	231
$ M_{edge} - M_{\tilde{\tau}\nu_{\tau}}^{T}  \le 20$	1076	1	381	58

Table: Number of signal and background event at an integrated luminosity of  $300 \text{fb}^{-1}$  for BP4.

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## Left-chiral tau-sneutrino reconstruction

► The  $\tilde{\nu}_{\tau_L}$  is produced predominantly via the decay of lightest chargino or second lightest neutralino .

$$\begin{array}{l} \chi_1^{\pm} \to \tilde{\nu}_{\tau_L}^{(*)} \tau^{\pm} \\ \chi_2^0 \to \tilde{\nu}_{\tau_L} \bar{\nu}_{\tau} \end{array}$$

► It has a sizeable decay branching fraction into a W<sup>7</sup>-pair (ranging from 34% to 80%).

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- ►  $\tilde{\nu}_{\tau_L}$  is produced mainly in association with  $\chi^0_{1,2}$  in cascade decay of squarks and gluinos-  $\chi^0_2 - \chi^0_{1,2} \rightarrow (\tilde{\nu}_{\tau_L}\nu_{\tau})(\tilde{\tau}\tau) \rightarrow (W\tilde{\tau}\nu_{\tau})(\tilde{\tau}\tau)$  $\chi^{\pm}_1 - \chi^0_{1,2} \rightarrow (\tilde{\nu}_{\tau_L}\tau)(\tilde{\tau}\tau) \rightarrow (W\tilde{\tau}\tau)(\tilde{\tau}\tau)$
- ► W has been reconstructed in its hadronic decay mode.



- We have adopted two different method to identify correct  $\tilde{\tau}W$ -pair
  - ► Using opposite sign charged trackthe W has been combined with the stau having same charge as that of the T
  - Using chargino-neutralino mass information-W has been combined with the τ̃ for which either of the following criteria is satisfied:

 $|M_{ au au_j} - m_{\chi_i^0}| > 20 \, GeV \, \, {
m or} \, \, |M_{ au au_jW} - m_{\chi_i^\pm}| < 20 \, GeV$ 

#### Chargino reconstruction

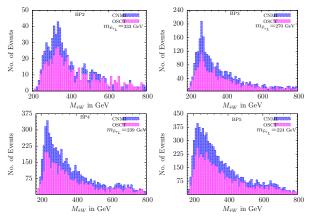


Fig.  $M_{\tilde{\tau}W}$  distribution for BP2, BP3, BP4 and BP5.

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# Conclusion

- The mSUGRA scenario, with conserved *R*-parity, can eminently lead to a situation in which the right-chiral sneutrino is the LSP and stau is long-lived.
- The right-chiral sneutrino becomes a viable dark-matter candidate, due to its very small neutrino Yukawa coupling.
- ► The stau being long-lived, the signal of SUSY consists of charged tracks in the muon chamber.

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- This feature can be used to reconstruct the mass of the SUSY particle in such a scenario.
- The method of reconstruction we have suggested is not limited to scenario with right-handed sneutrino LSP alone. It can be applied to all those scenarois in which stau is long-lived. (*e.g.* GMSB,..)
- The existence of quasi-stable massive charged particles is an interesting possibility to look for at the LHC...

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- The existence of quasi-stable massive charged particles is an interesting possibility to look for at the LHC...

It not only offers a distinct SUSY signal, but also opens a new vista in the reconstruction of the superparticle masses.

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# Thank you!

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