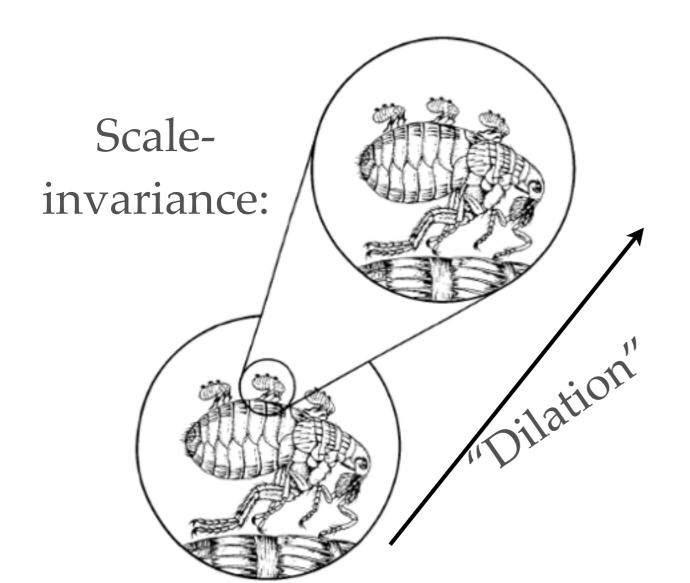
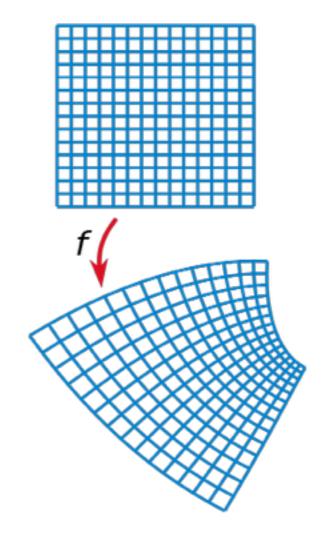
Universal Dynamics from the Conformal Bootstrap

Liam Fitzpatrick Stanford University

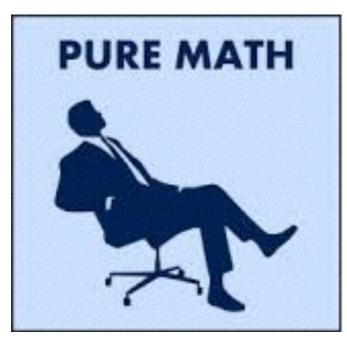
in collaboration with Kaplan, Poland, Simmons-Duffin, and Walters

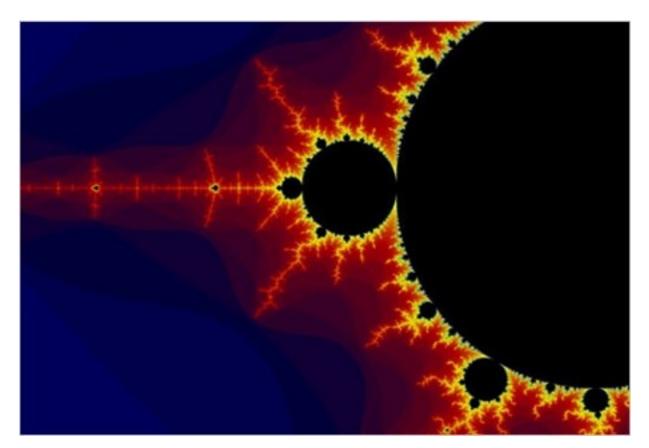
Conformal Symmetry Conformal = coordinate transformations that preserve angles in special relativity Includes scale transformations, rotations, etc.

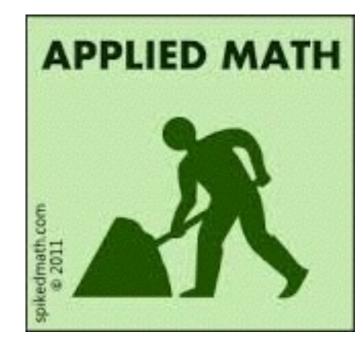




Why Study CFTs?





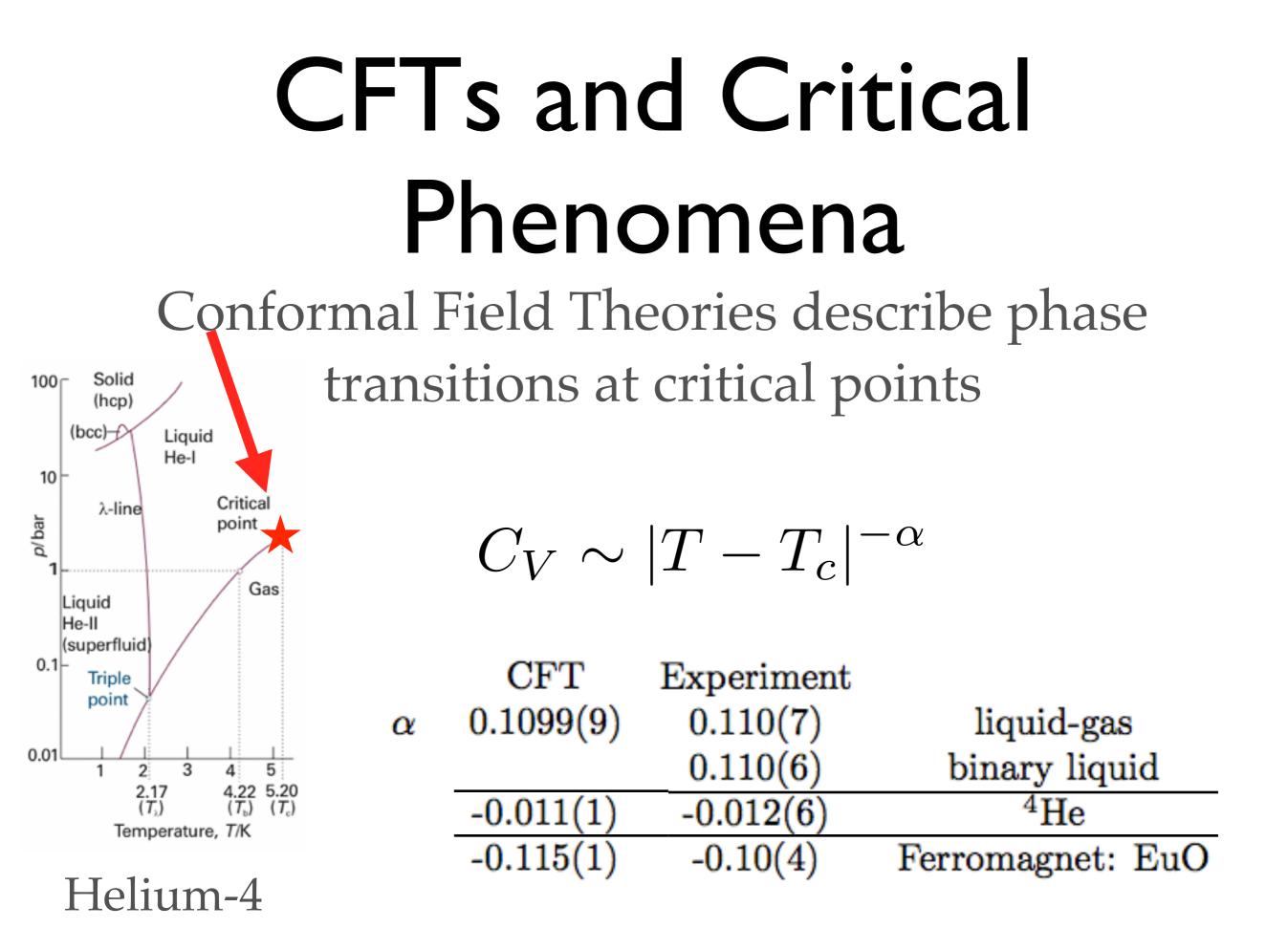


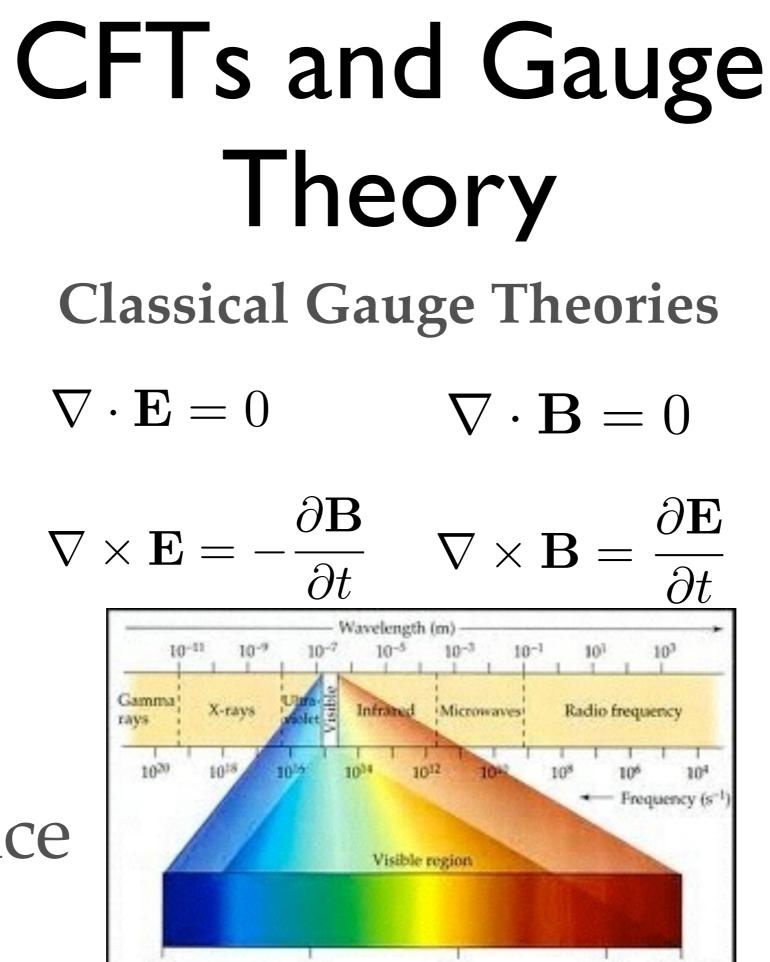


Why Study CFTs?

Conformal Field Theories (CFTs) are central objects in physics

- Phase Transitions and Critical Phenomena
- Classical Gauge Theories in 4d
- Endpoints and Intermediate Regimes of RG Flow
- and much more...





Scaleinvariance

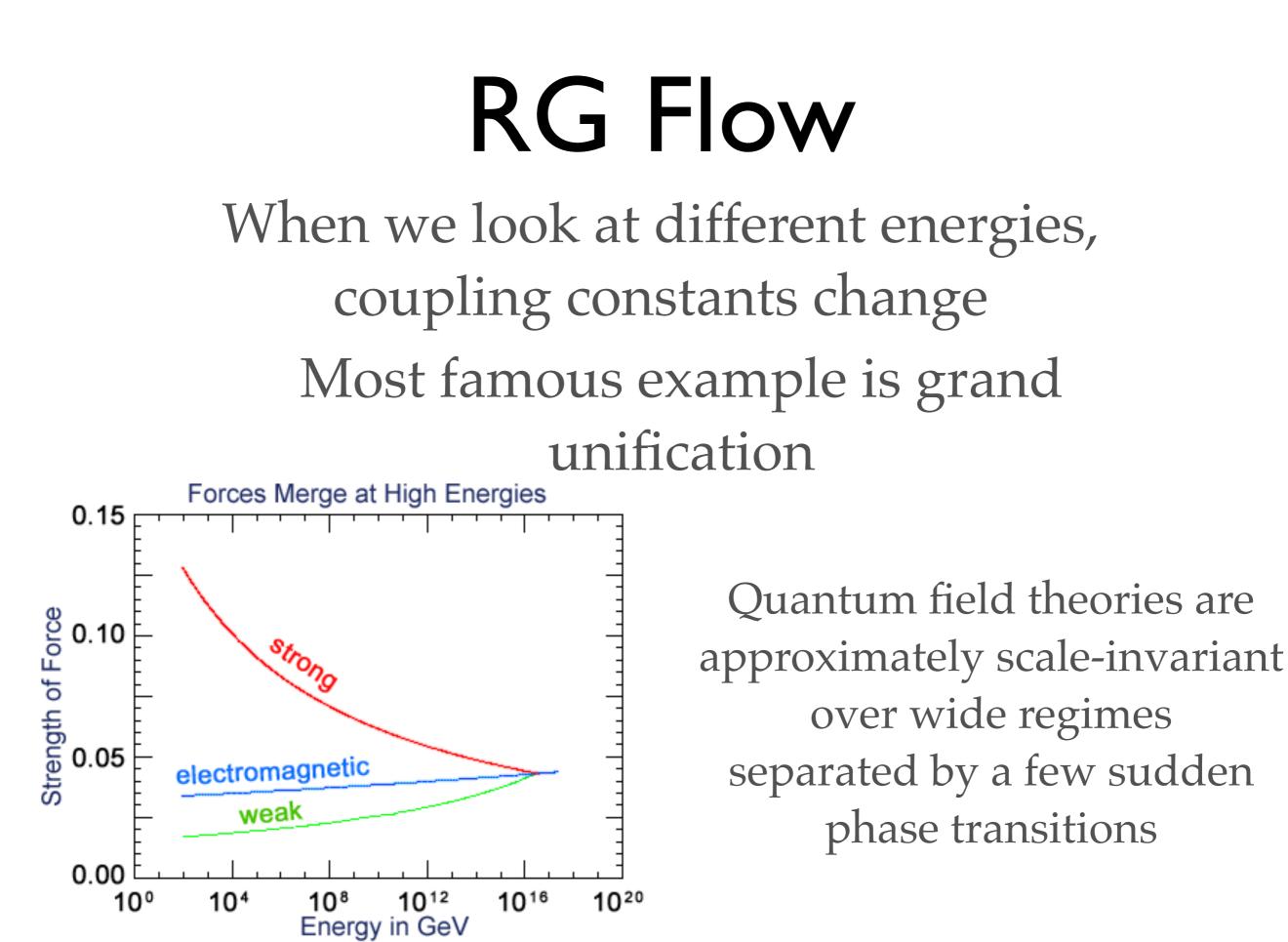
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750 mm

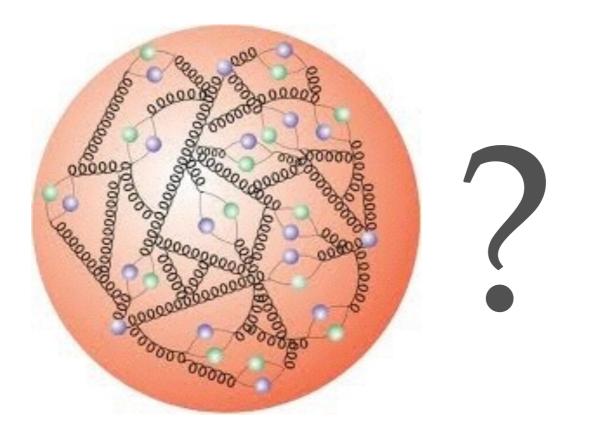


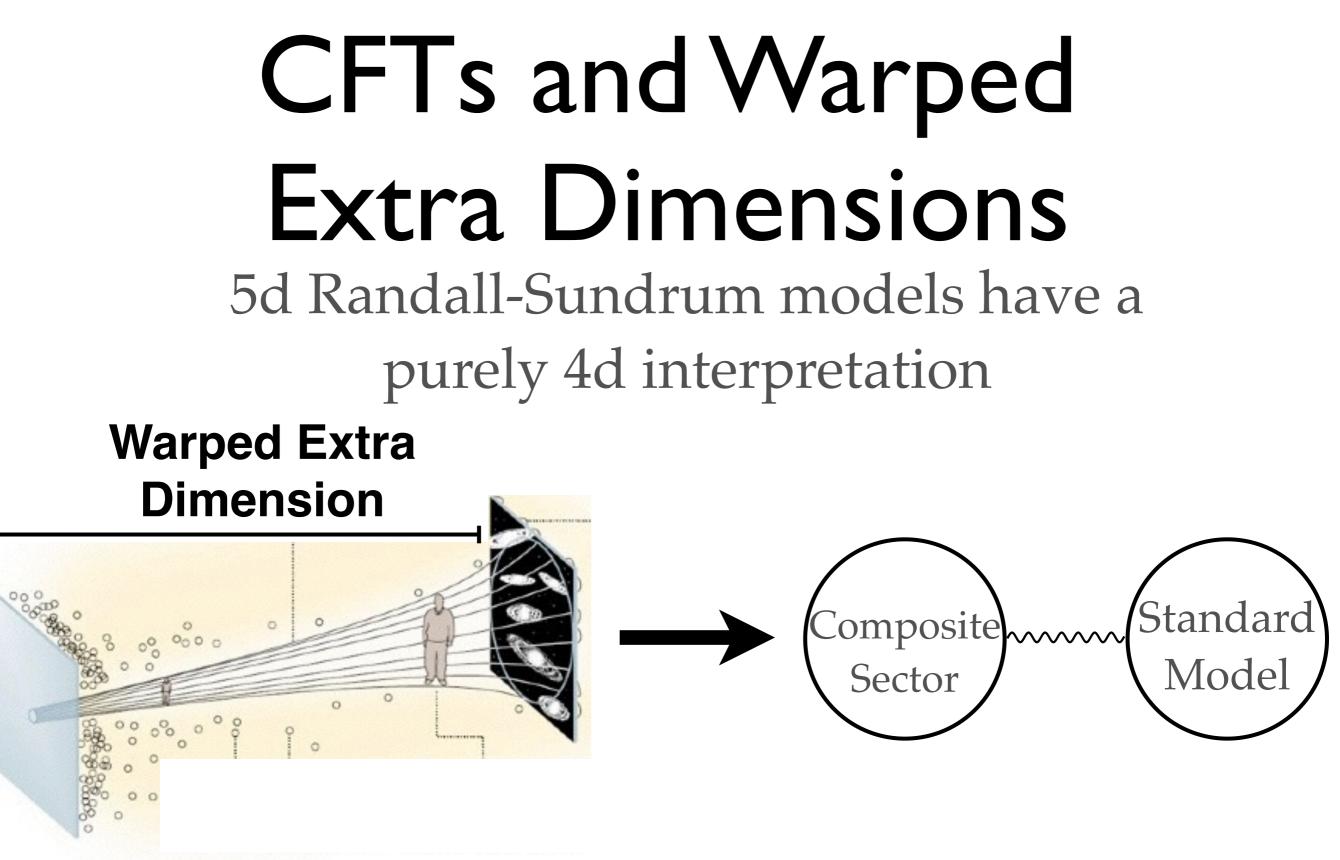
CFTs and Strong Coupling

Strongly coupled fixed points

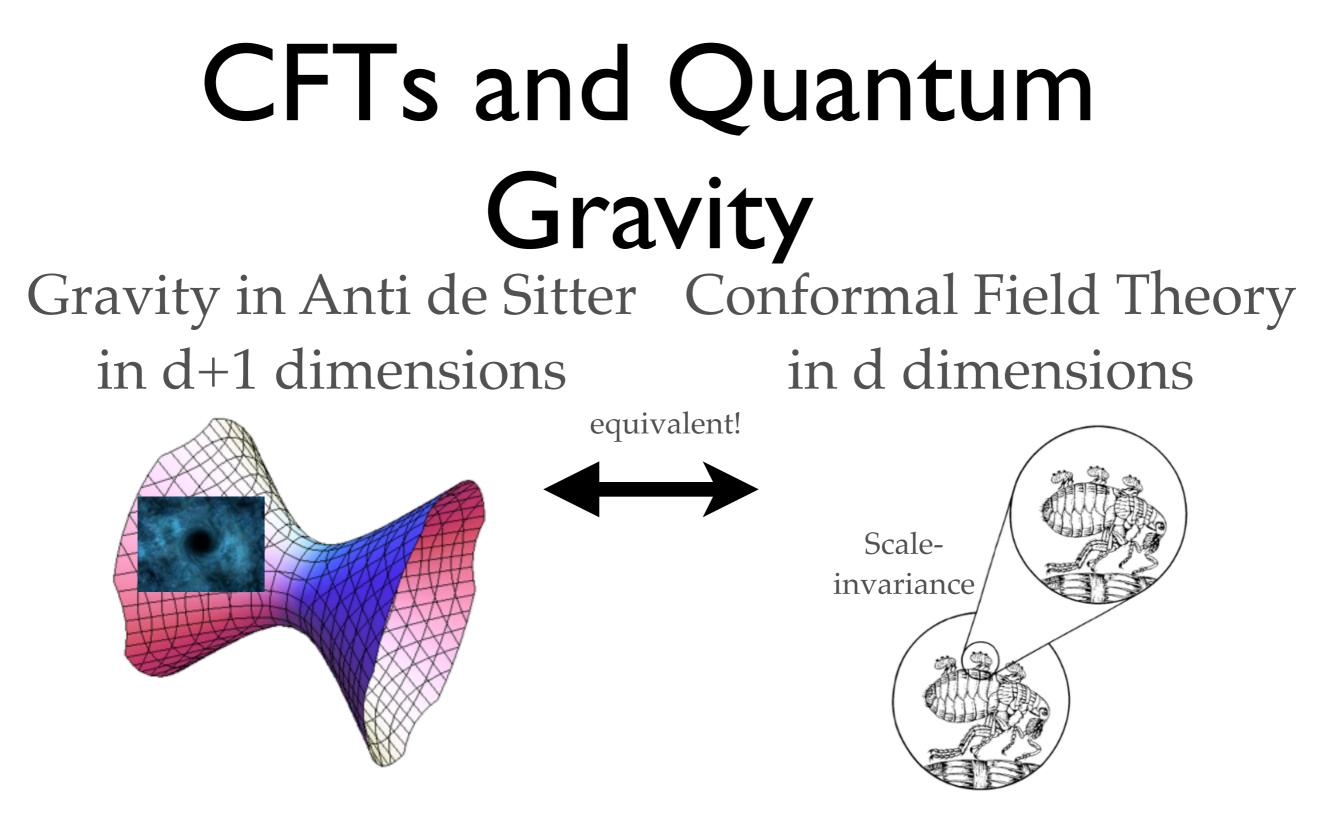
Strongly coupled theories are difficult to study.

Conformal invariance can give us a powerful tool to study their behavior.





They describe *some classes* of strongly-coupled theories.



So studying CFTs teaches us about gravity, and vice versa!

CFTs and Quantum Gravity

Why does an extra dimension emerge from a CFT?

When is physics *local* in this extra dimension?

Why do gravitational interactions emerge?

Do they look like Newtonian gravity/General Relativity VS modified theories of gravity?

CFTs and Quantum Gravity

What can we learn about black hole dynamics?

Hawking radiation: Semi-classical limit says black holes have a

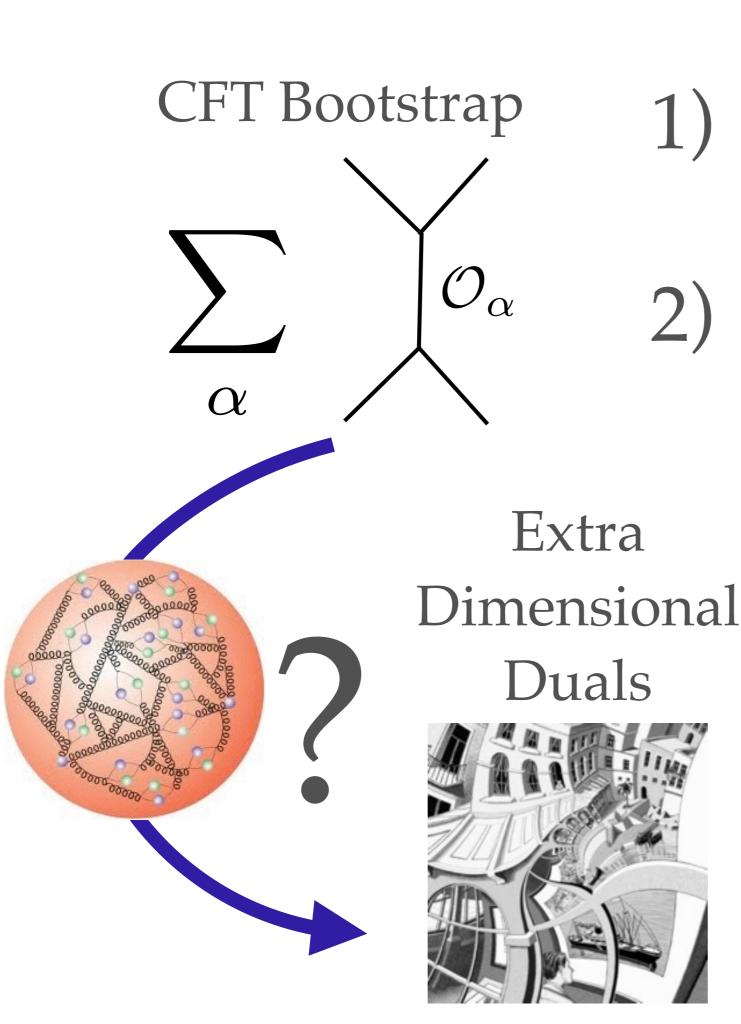
temperature.

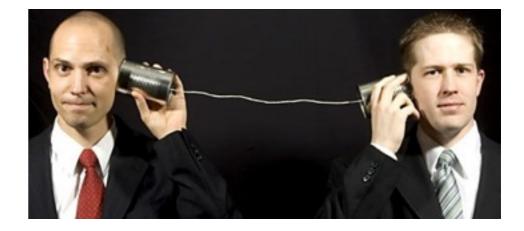
But if this is exactly true, then information is lost! Not consistent with Quantum Mechanics. Can we understand "pure" states mimicking a thermal states?

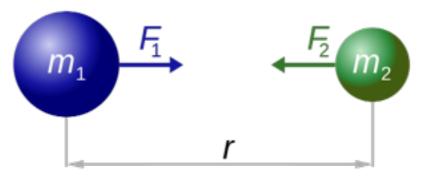
CFTs and Quantum Gravity

Goals:

- 1) Use the conformal bootstrap to learn about universal properties of CFTs
- 2) Interpret our results as statements about universal properties of the dual theories with extra dimensions. Specifically, want to derive *locality* and *gravitation* at long-distances.
 - I however *do not* assume anything a priori about the dual gravitational theory.





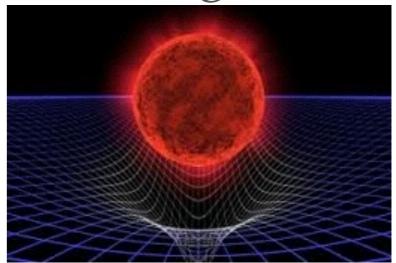


1)

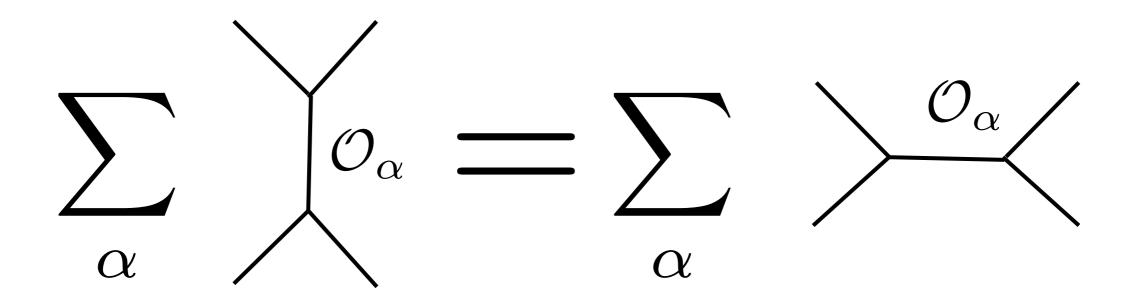
2)

 $F_1 = F_2 = G \frac{m_1 \times m_2}{r^2}$

3) 2d CFTs: Heavy states effectively deform CFT geometry



The Conformal Bootstrap



Operators

In conformal theories, a key role is played by "operators", which can be any local observable

Simple Example: density operator ho(x)

We study correlation functions among operators

 $\langle \rho(x)\rho(y)\rho(z)\rangle$

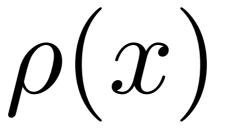
Operators and States

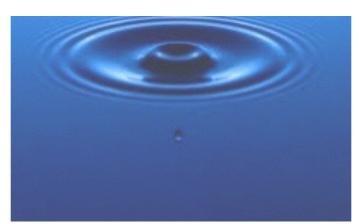
Every operator creates a unique state, and vice versa:

$$\rho(x)|0\rangle \leftrightarrow |\rho\rangle$$

By "measuring" ρ , we perturb the vacuum

and put it in a new state.





Operator Products Start with insertion of two operators



Decompose into a convenient basis at a fixed radius. E.g. Spherical harmonics Quantum: Decompose wavefunction $\psi(\theta, \phi) = \sum c_{\ell,m} Y_{\ell,m}(\theta, \phi)$

Operator Product Expansion

Products of operators can be expanded as sums of operators

$$\rho(x)\rho(y) = \sum_{i} c_{i}(x-y)\mathcal{O}_{i}(y)$$
Wilson '69
Zimmerman '70
This is very powerful when used inside correlation functions

$$\langle \rho(x_{1})\rho(x_{2})\rho(x_{3})\rho(x_{4})\rangle = \sum_{\alpha} c_{\alpha}^{2}g_{\Delta_{\alpha},\ell_{\alpha}}(x_{i})$$
"conformal block"

So all correlators are determined by the two- and three-point functions. Conformal symmetry reduces the terms in the sum to *unknown numbers* times *known functions*.

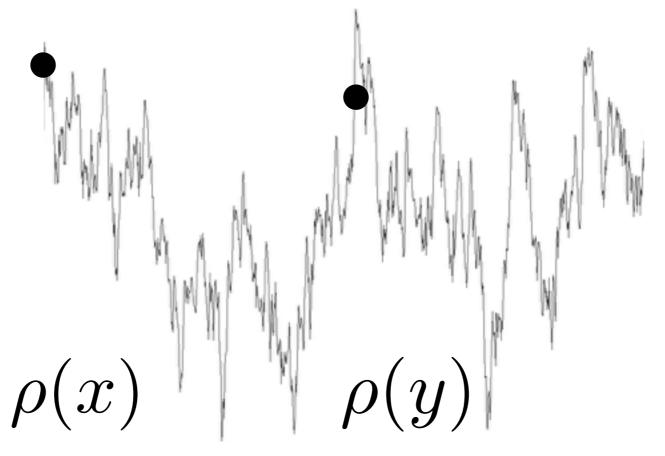
Conformal Bootstrap
We can contract in the "s-channel"

$$\langle \rho(x_1)\rho(x_2)\rho(x_3)\rho(x_4)\rangle = \sum_{\alpha} \checkmark^{\mathcal{O}_{\alpha}}$$

But we can also clearly contract in the "t-channel"
 $\langle \rho(x_1)\rho(x_2)\rho(x_3)\rho(x_4)\rangle = \sum_{\alpha} \checkmark^{\mathcal{O}_{\alpha}}$
The conformal bootstrap is the constraint that these give the
same answer. This simple statement contains an enormous
amount of physics.
 $\sum_{\alpha} \checkmark^{\mathcal{O}_{\alpha}} = \sum_{\alpha} \checkmark^{\mathcal{O}_{\alpha}}$

Conformal Bootstrap

Variance of fluctuations diverges on short distances

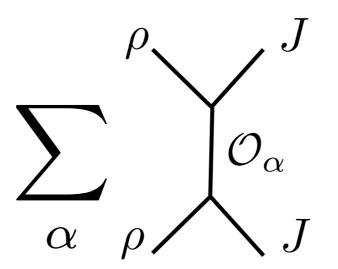


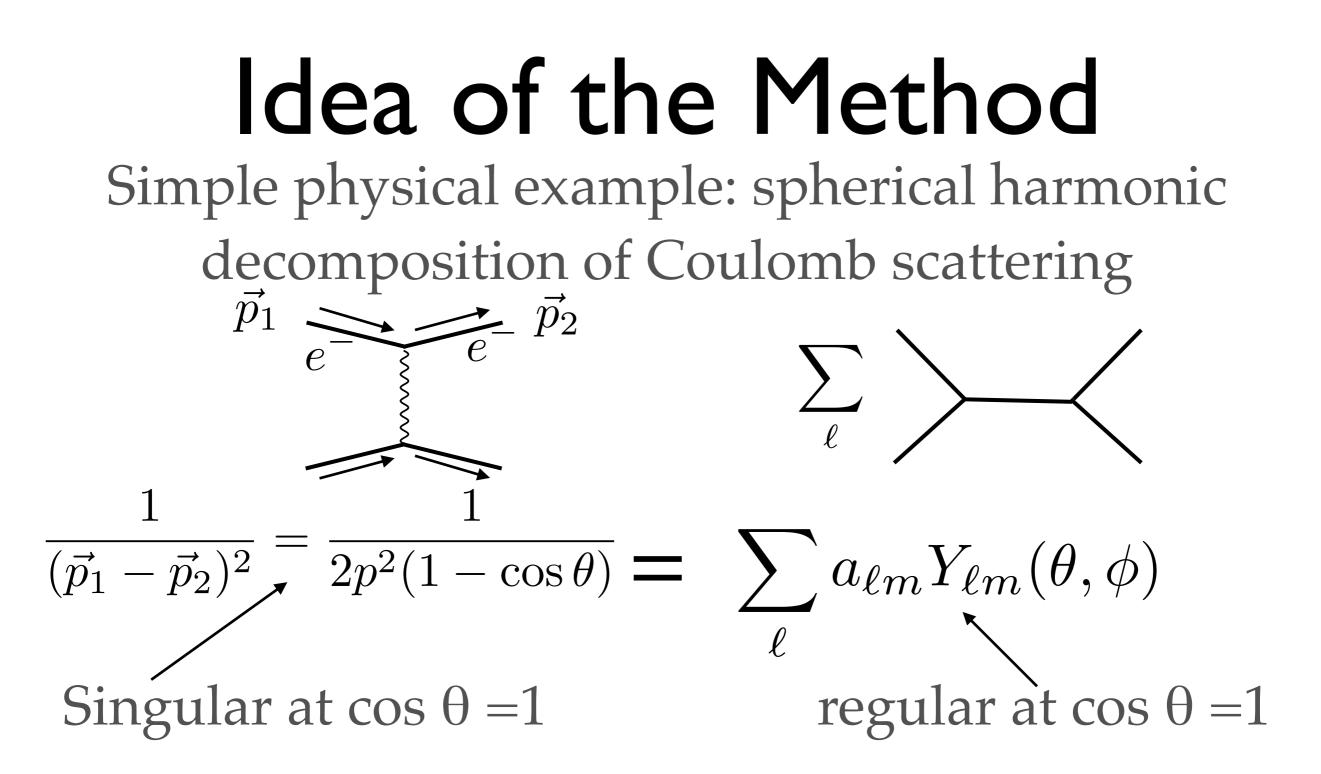
Rate of divergence is fixed by scaling dimensions

$$\langle \rho(x)\rho(y)\rangle \sim \frac{1}{|x-y|^{2\Delta_{\rho}}}$$

$\begin{array}{c} \textbf{Conformal Bootstrap}\\ \text{This divergence contributes to the correlation function of}\\ \text{four operators - e.g. two } \rho's \text{ and two } J's.\\ \langle \rho(x)\rho(y)J(z)J(w)\rangle \supseteq \langle \rho(x)\rho(y)\rangle \langle J(z)J(w)\rangle\\ \rho(y) \qquad J(z)) \end{array}$

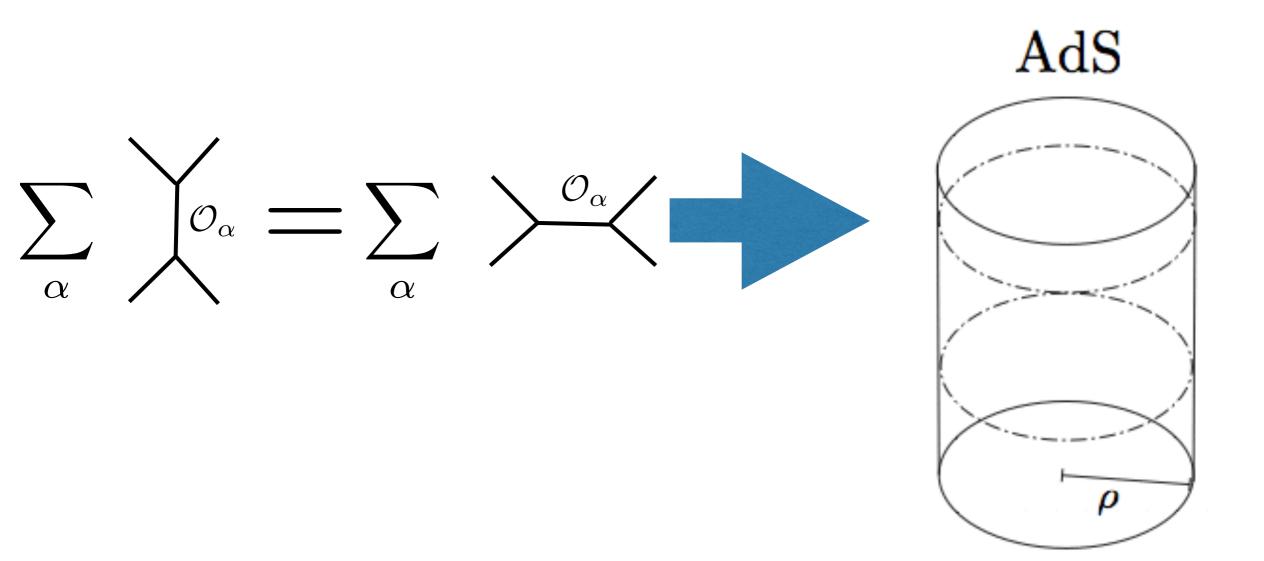
These singularities *must be reproduced* by the sums in the bootstrap equation.



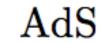


So we must have an infinite tower of spherical harmonics at large spin

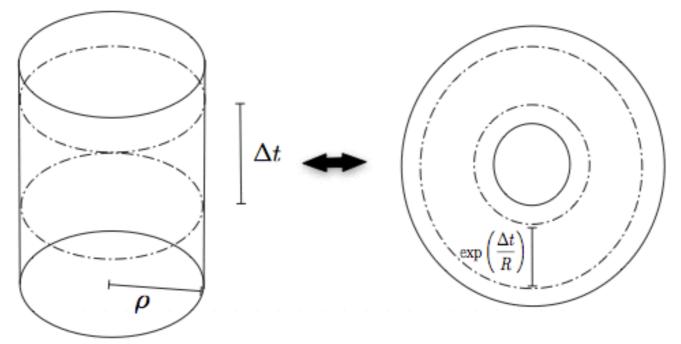
We use this type of analysis of the bootstrap to prove the existence of certain large spin operator in CFTs, as well as obtaining their dimensions and OPE coefficients. This will demonstrate important universal properties of the long-distance physics of their AdS duals.

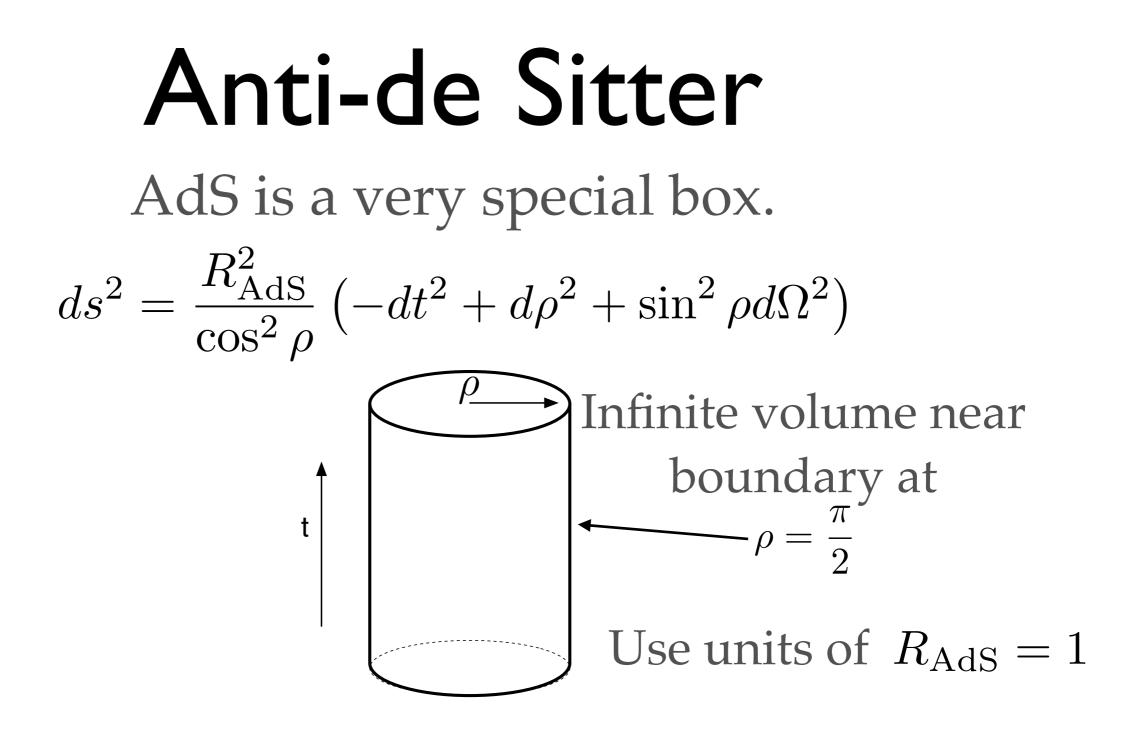


AdS Kinematics and CFT Kinematics



CFT





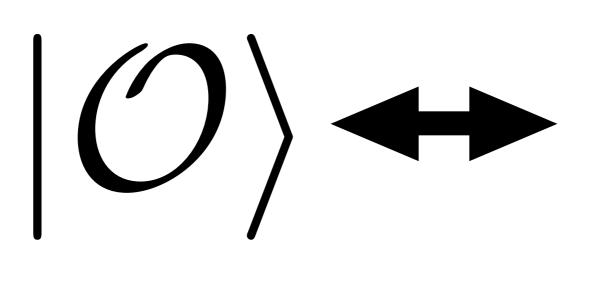
The isometries of AdS are in one-toone correspondence with the generators of the conformal group

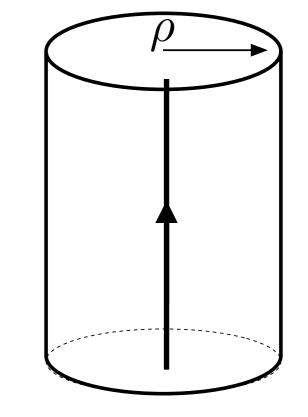
CFT Scaling AdS Energy Dimension HAdS LCFT CFT "Dilatation" **AdS Hamiltonian** Generates scaling Generates time SENATOR evolution $H_{\rm AdS}$ $D_{\rm CFT}$

States have (center-of-mass) wavefunctions in AdS that are completely determined by symmetry

E.g. Primary states = States at rest in the center of AdS

Primary (roughly, not total derivatives):

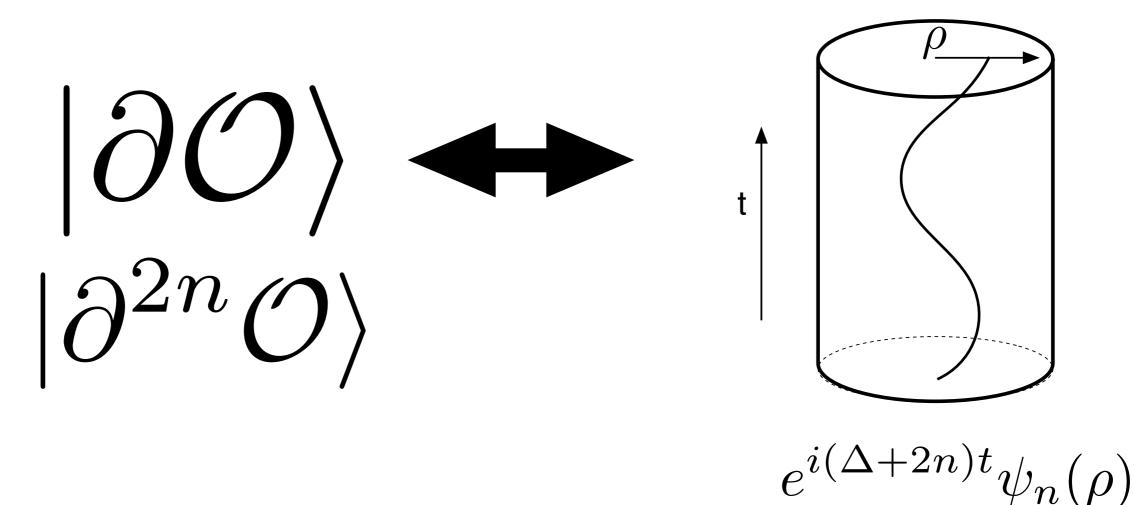




 $\psi(x) = e^{i\Delta t} \cos^{\Delta} \rho$

AdS Wavefunctions

Descendant states = states oscillating in AdS



AdS Wavefunctions

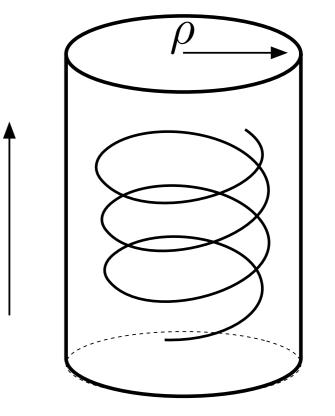
Descendant states with large spin

states orbiting center of AdS

 $|\partial_{\mu_1} \dots \partial_{\mu_\ell} \mathcal{O} \rangle$

They orbit center of AdS at a radius given by

 $r \sim L$



 $e^{i(\Delta+\ell)t}\psi_{\ell}(\rho)Y_{\ell}(\Omega)$

Locality

So far, coordinates in these wavefunctions are just labels. To really count as an extra dimension, there should be a sense of "locality" or "cluster decomposition":

Say we have a "rock" state:



and a "scissors" state:

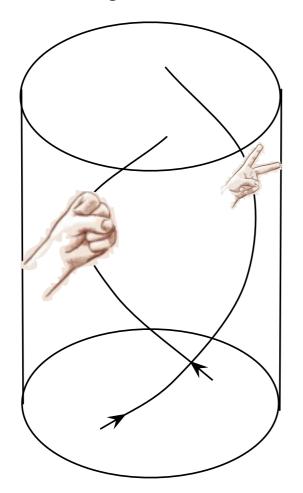


Can they co-exist peacefully? I.e. do there exist "rock+scissors" states where they are "far apart" and ignore each other?



Cluster Decomposition

How do can make these states "far apart in AdS"? Solution: give each of them large angular momentum so they orbit the center at large distance



So we want to look for states in the CFT with large angular momentum.

> ALF, Kaplan, Poland Simmons-Duffin

Cluster Decomposition

So by "cluster decomposition", I will mean that given any two primary operators ϕ_{rock} and ϕ_{sciss} in the CFT,

Theorem: all CFTs in $d \ge 3$

There exist an infinite number of operators at large spin, with dimensions $\Delta_0(L) + \gamma(L) \text{ where } \lim_{L \to \infty} \gamma(L) = 0$

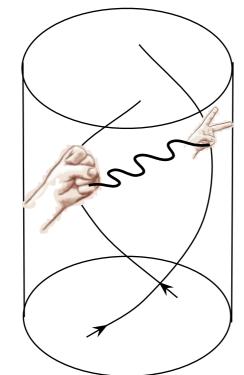
"free" part "interaction part"

Proof follows from matching singularities in
bootstrap equationALF, Kaplan, Poland
Simmons-Duffin

Binding Energy

Furthermore, binding energies for weak AdS interactions show up as a correction to this dimension i.e. an anomalous dimension $\Delta(L) = \Delta_0(L) + \gamma(L)$

This gives us a simple CFT handle on the strength of interactions at large distances



 $\gamma(L) \sim \gamma(r)$

ALF, Kaplan, Poland Simmons-Duffin

"Newtonian" Gravity in $d \ge 3$ The bootstrap also has a universal singularity from the

stress tensor in the CFT.

At long distances, this produces a weak "Newtonian"

binding energy $\gamma(r) \sim G_N \frac{m_{\rm rock} m_{\rm sciss}}{r^{d-2}}$

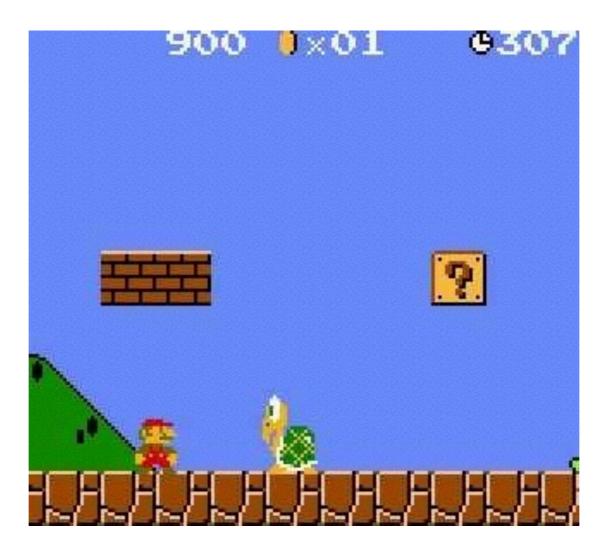
 $\frac{\text{AdS} \quad \text{CFT}}{m_{\text{rock}} \leftrightarrow \Delta_{\text{rock}}}$ $\frac{r \leftrightarrow L}{G_N \leftrightarrow \frac{1}{\text{Ndof}}}$

All interacting CFTs in d≥3 have this "Newtonian gravity" correction to their binding energy at large spin.



ALF, Kaplan, Poland Simmons-Duffin

2d CFTs



Why Focus on 2d?

Useful toy model: conformal symmetry is much bigger! Dual to 3d gravity in AdS: Gravitons have no degrees of freedom, but there are still black holes. Some other toy 2d Ising Model models:

Why Focus on 2d?

Useful toy model: conformal symmetry is much bigger! Dual to 3d gravity in AdS: Gravitons have no degrees of freedom, but there are still black holes. Some other toy 2d QCD at large N: models: the gluon has no DOFs, and the theory is solvable.

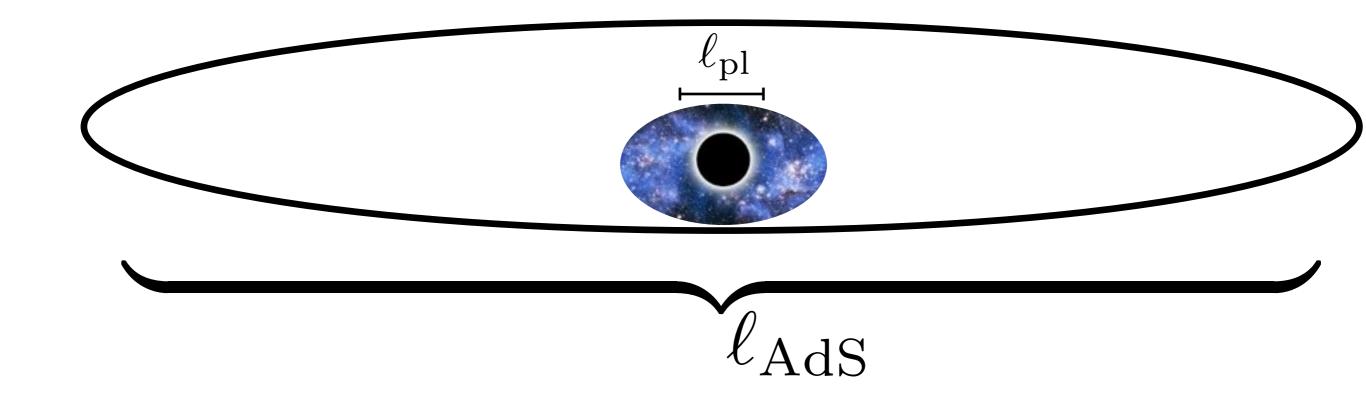


Large C

ALF, Kaplan, Walters

Consider large CFT central charge : essentially, large number of degrees of freedom. Like a classical limit.

The central charge is related to G_N in AdS by $c = \frac{3}{2G_N}$ so this is a "semi-classical" gravity limit Brown, Henneaux, '86



Large C How do we get interesting effects in gravity at $G_N \rightarrow 0$? Keep $G_N M \sim R$ fixed What does this limit do in the CFT?

$$G_N \sim \frac{1}{c} \qquad \Delta \sim M \implies G_N M \sim \frac{\Delta}{c}$$

and in fact, by conformal symmetry

$$\phi(x)\phi(y) \supset \frac{\Delta}{c}T_{\mu\nu} \qquad \phi(x)\phi(y) \supset \left(\frac{\Delta}{c}T_{\mu\nu}\right)$$

so at $\frac{\Delta}{c} \sim \mathcal{O}(1)$, we must include all powers of
the stress tensor.

 $\sim n$

This corresponds to resumming all multi-graviton contributions in AdS.

Large C and Classical Backgrounds Again, focus on four operators: $\phi^{\dagger}(x) \stackrel{1, T_{\mu\nu}}{\underset{T^{2}_{\mu\nu}}{\overset{\chi^{\dagger}(z)}{\overset{\chi^{\star}(z)}{\overset{\chi^{\star}(z)}{\overset{\chi^{\star}(z)$

We want to calculate the contribution from all powers of T_{μν}'s/h_{μν}'s This "dressed" contribution is a **2d conformal block**

 $\phi(0)$

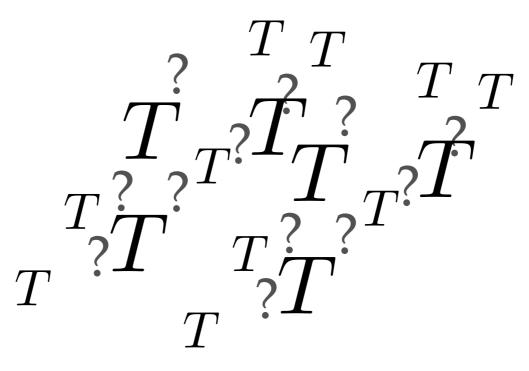
 $\chi(w)$

2D Conformal

Conformal group in 2d relates all of these contributions to each other!

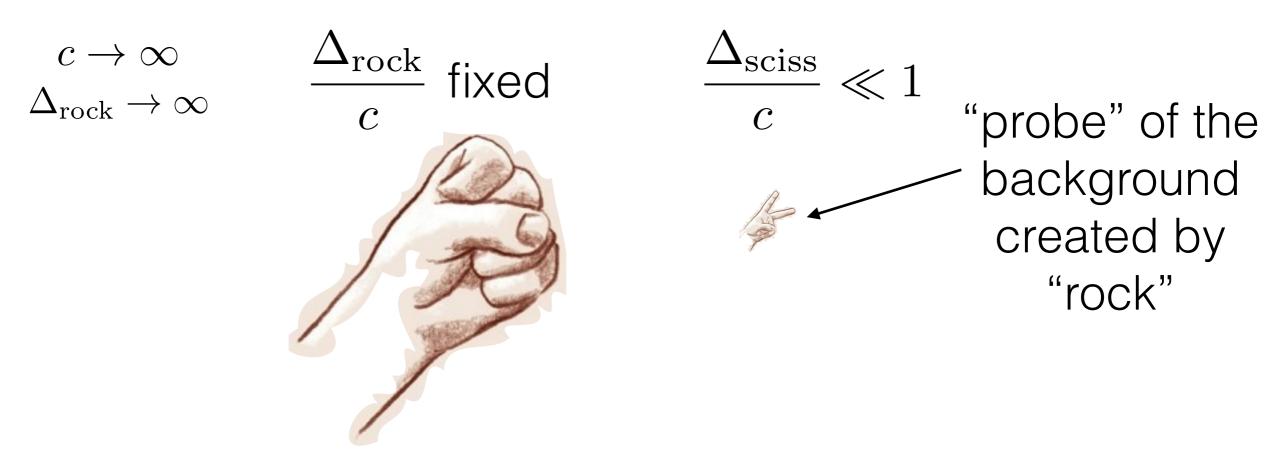
 $1, T, T^2, T^3, \ldots$ All related by symmetry.

So in principle, their contributions are known and computable! But not known in practice.



Large C and Classical Backgrounds ALF, Kaplan, Walters

Consider the following "probe" limit:



Total contribution of *all* Tⁿ can be computed and has a simple effect in this limit.

Large C and Classical Backgrounds

Idea: do a conformal transformation to put the CFT in a new metric.

Due to quantum effects, T_{µv} gets a vacuum expectation value:

 $\langle \phi_{\rm rock} \phi_{\rm rock} T \rangle \rightarrow \langle \phi_{\rm rock} \phi_{\rm rock} T \rangle - \langle \phi_{\rm rock} \phi_{\rm rock} \rangle \langle T \rangle$

Want this to cancel

It turns out that this can always be done. At large C, all Tⁿ contributions are completely canceled! ALF, Kaplan, Walters

Large C and Classical Backgrounds

So the "conformal blocks" can be calculated in this limit, and they are just the contribution from the state without Tⁿ, but in a new background metric.

We get this without making any reference to gravity!

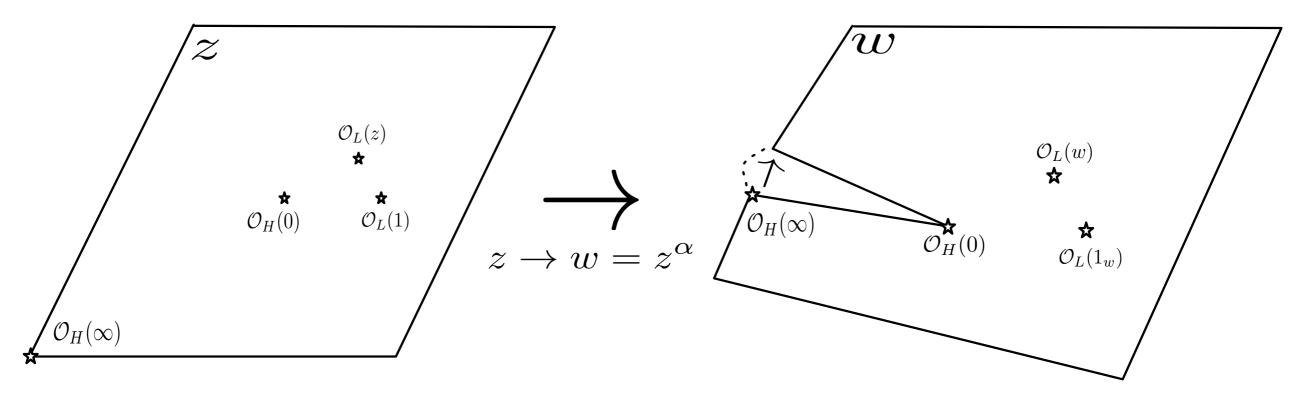
> ALF, Kaplan, Walters

Large C and Classical Backgrounds

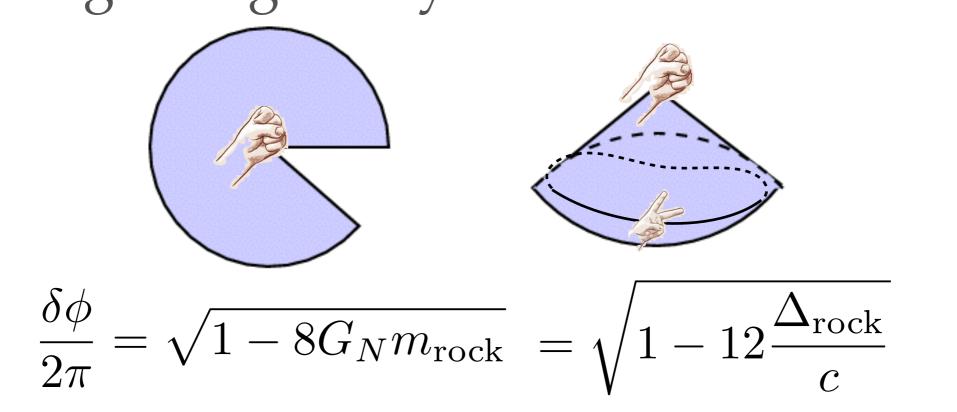
The coordinate transformation is very simple:

$$z \to w(z) = z^{\alpha}$$
 $\alpha_{\rm rock} = \sqrt{1 - \frac{12\Delta_{\rm rock}}{c}}$

Think of this as cutting out a wedge of space:



Relation to Gravity In 3-dimensional AdS, there is a minimum threshold for black hole masses: $m_{\min} = \frac{1}{8G_N} = \frac{c}{12}$ Below this threshold, a local mass just makes a deficit angle singularity. It matches the CFT result!



Classical Background
What about
$$\Delta_{\text{rock}} \ge \frac{c}{12}$$
? $\alpha_{\text{rock}} = \sqrt{1 - \frac{12\Delta_{\text{rock}}}{c}} \equiv 2\pi i T_{\text{rock}}$
What is the \swarrow two-point function in this
background?
The contribution from vacuum + Tⁿ is
 $\langle \phi_{\text{rock}} | \oint(t) \oint(0) | \phi_{\text{rock}} \rangle = \left(\frac{\pi T_{\text{rock}}}{\sin(\pi T_{\text{rock}}t)}\right)^{\Delta_{\text{sciss}}} \sum_{\text{Exactly thermal}}$
-Eigenstate Thermalization Hypothesis

$$|\phi_{\rm rock}\rangle =$$



Relation to Gravity

What about $\Delta_{\text{rock}} > \frac{c}{12}$? $(m > m_{\min})$ Above this threshold, a black hole horizon forms in AdS



 $T_{\rm rock}$ matches the temperature of the AdS black hole!

Future Directions $d \ge 3$

ALF, Kaplan,

Walters In 2d, we used the vacuum expectation value of T to cancel the single-T contribution $\langle \phi_{
m rock} \phi_{
m rock} T \rangle \rightarrow \langle \phi_{
m rock} \phi_{
m rock} T \rangle - \langle \phi_{
m rock} \phi_{
m rock} \rangle \langle T \rangle$ This can be done in 4d and 6d as well! The result depends only on the "a_d" anomaly coefficient. New coordinates are periodic in time, with period

$$\beta^{-d} \sim \frac{\Delta_{\text{rock}}}{a_d}$$

which is the right scaling in the limit of very large AdS black holes in Einstein gravity!

Future Directions Going beyond the Semi-Classical Limit in 2D

Additional tool: put together with information on scaling dimensions from modular invariance?

$$\rightarrow \Delta_{\alpha} \sum_{\alpha} \mathcal{O}_{\alpha} = \sum_{\alpha} \mathcal{O}_{\alpha} \rightarrow OPE$$

Extract non-perturbative information about gravity using the bootstrap?

Conclusions

There is an enormous amount of information contained in the bootstrap equation.

Recent results on conformal blocks give us a sharp tool for extracting information.

The constraints of large C + "only a few light operators" seem to be far-reaching, and still have much to teach us.

Old dream: AdS₃ gravity may be simple enough to have tractable solutions - i.e. 2d QCD (`t Hooft model) for gravity

The End