

Effective WIMPs

Jeff Hutchinson

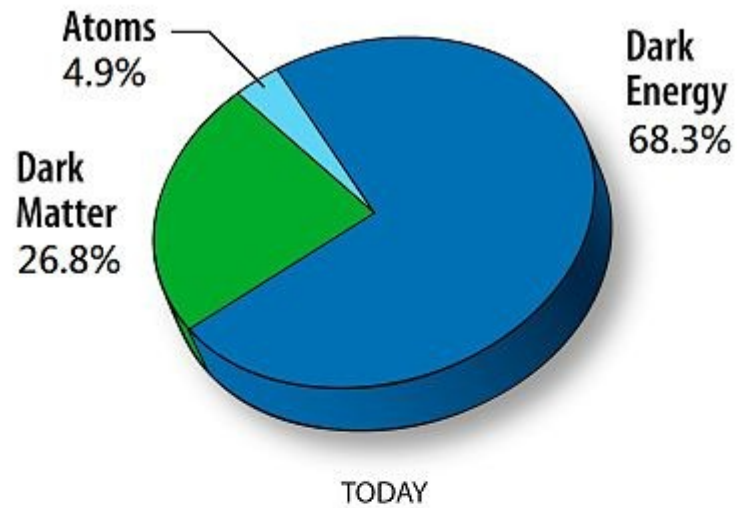
University of California, Davis

With

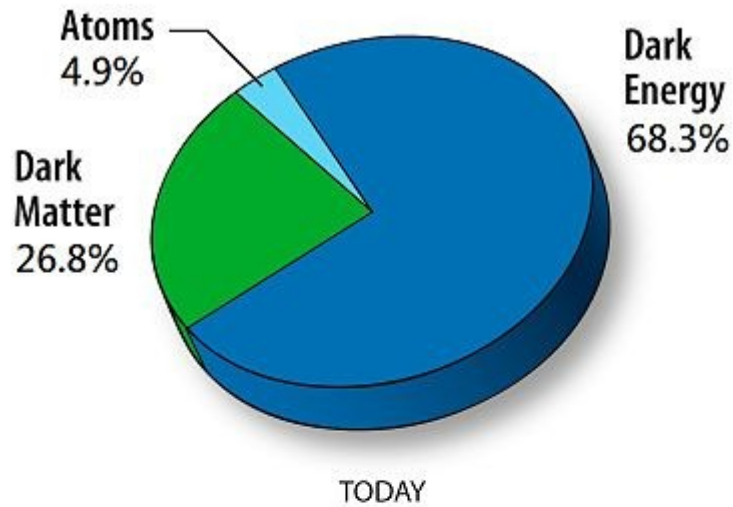
Spencer Chang, Ralph Edezhath, & Markus Luty

Based on: [arXiv:1307.8120](https://arxiv.org/abs/1307.8120)

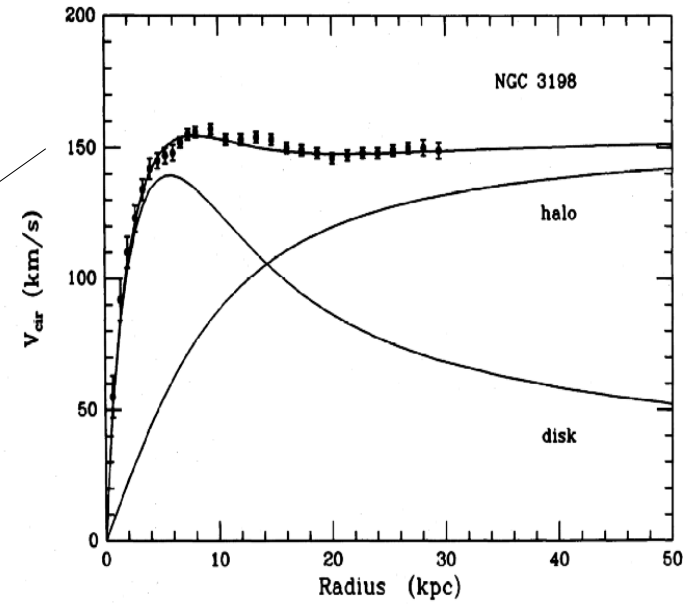
Dark Matter



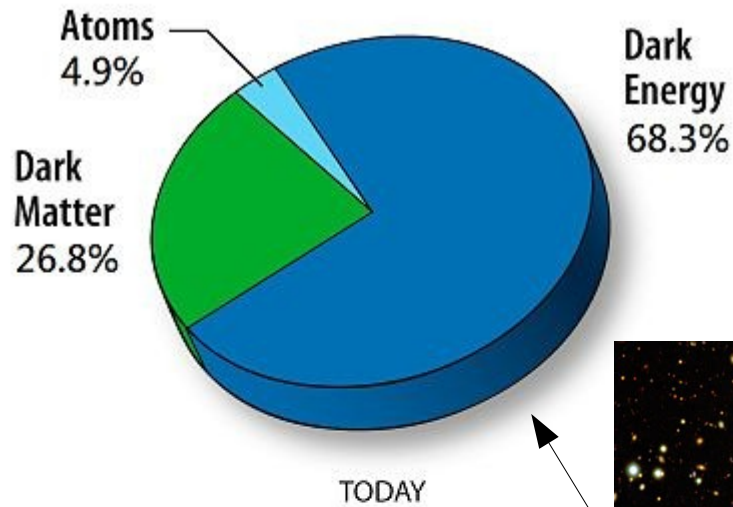
Dark Matter



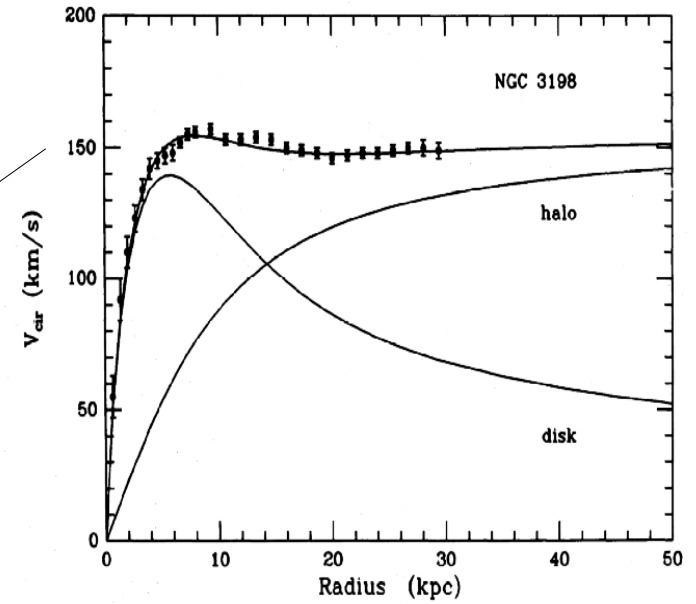
DISTRIBUTION OF DARK MATTER IN NGC 3198



Dark Matter



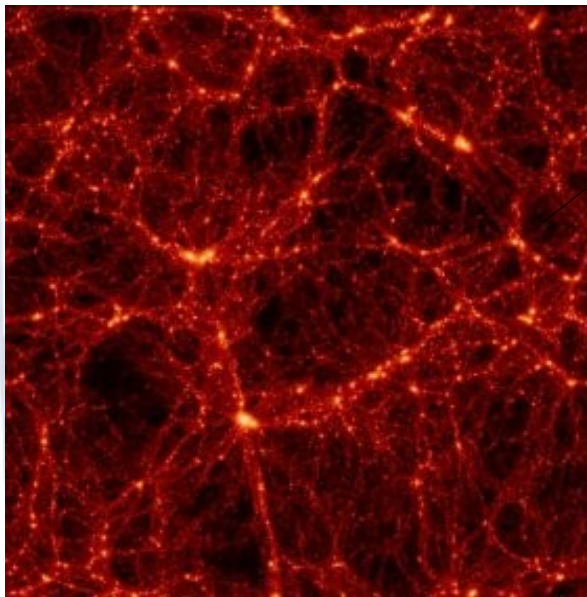
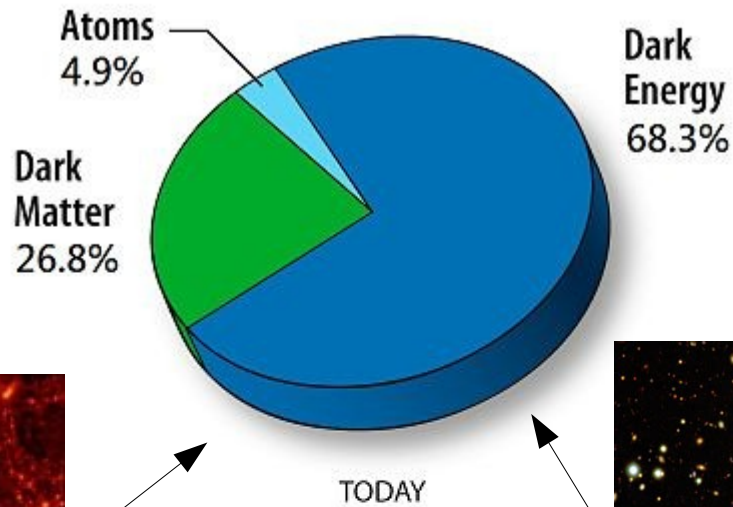
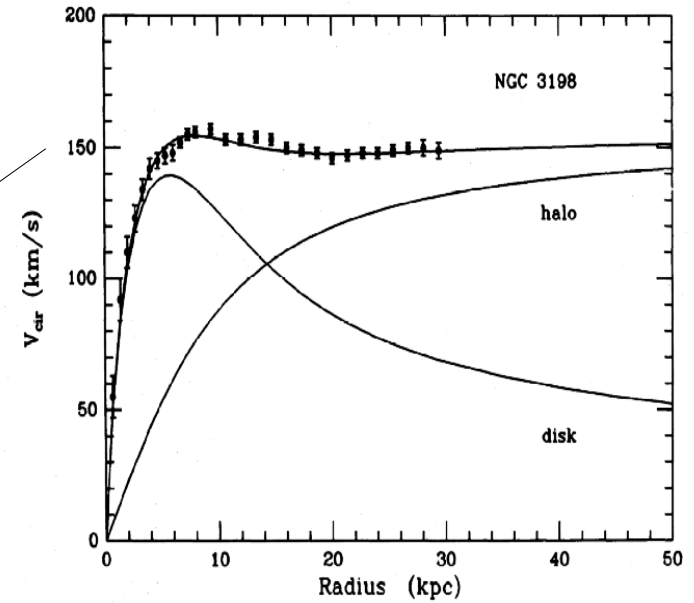
DISTRIBUTION OF DARK MATTER IN NGC 3198



Weak Lensing

Dark Matter

DISTRIBUTION OF DARK MATTER IN NGC 3198

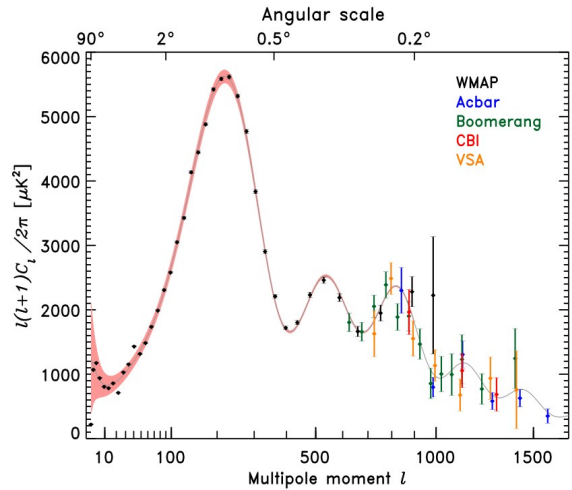
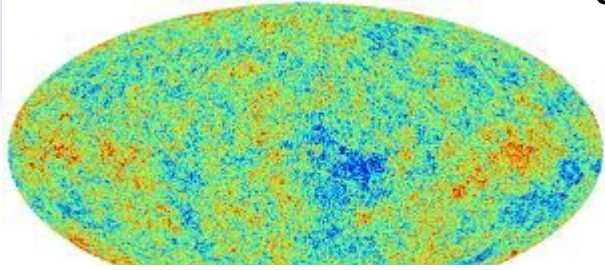


Large Scale Structure Formation

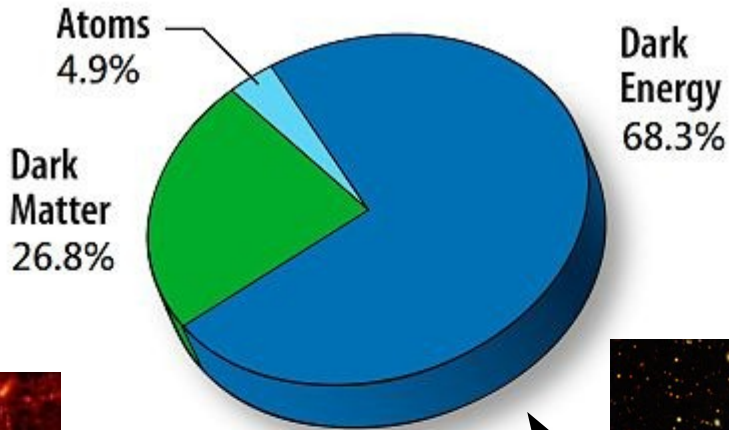


Weak Lensing

Cosmic Microwave Background



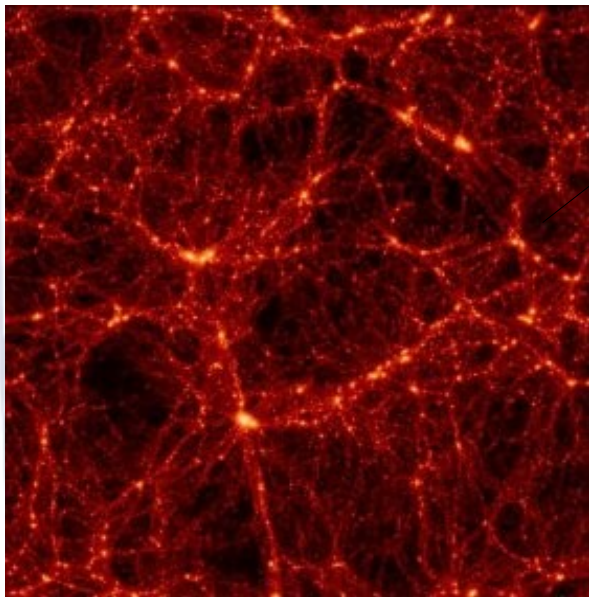
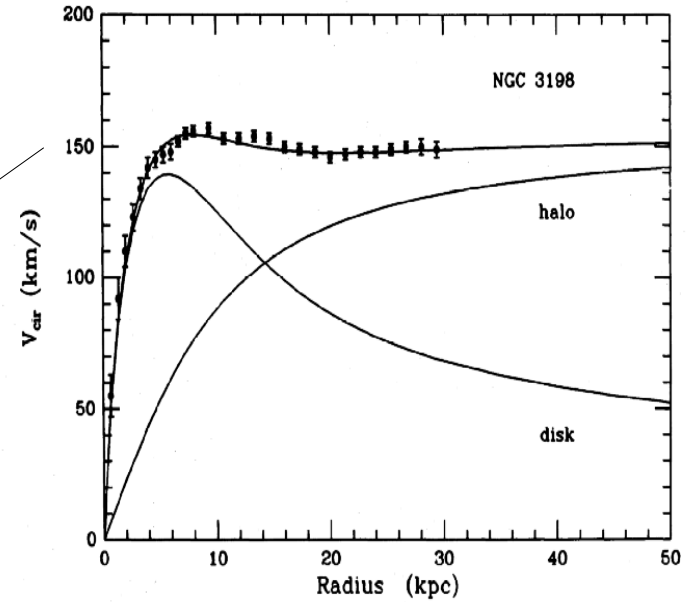
Dark Matter



TODAY

Galactic Rotation Curves

DISTRIBUTION OF DARK MATTER IN NGC 3198



Large Scale Structure Formation



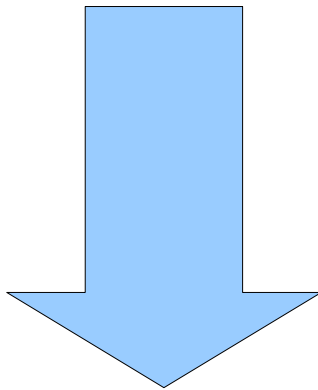
Weak Lensing

WIMPs

WIMP = “Weakly Interacting Massive Particle”

WIMPs

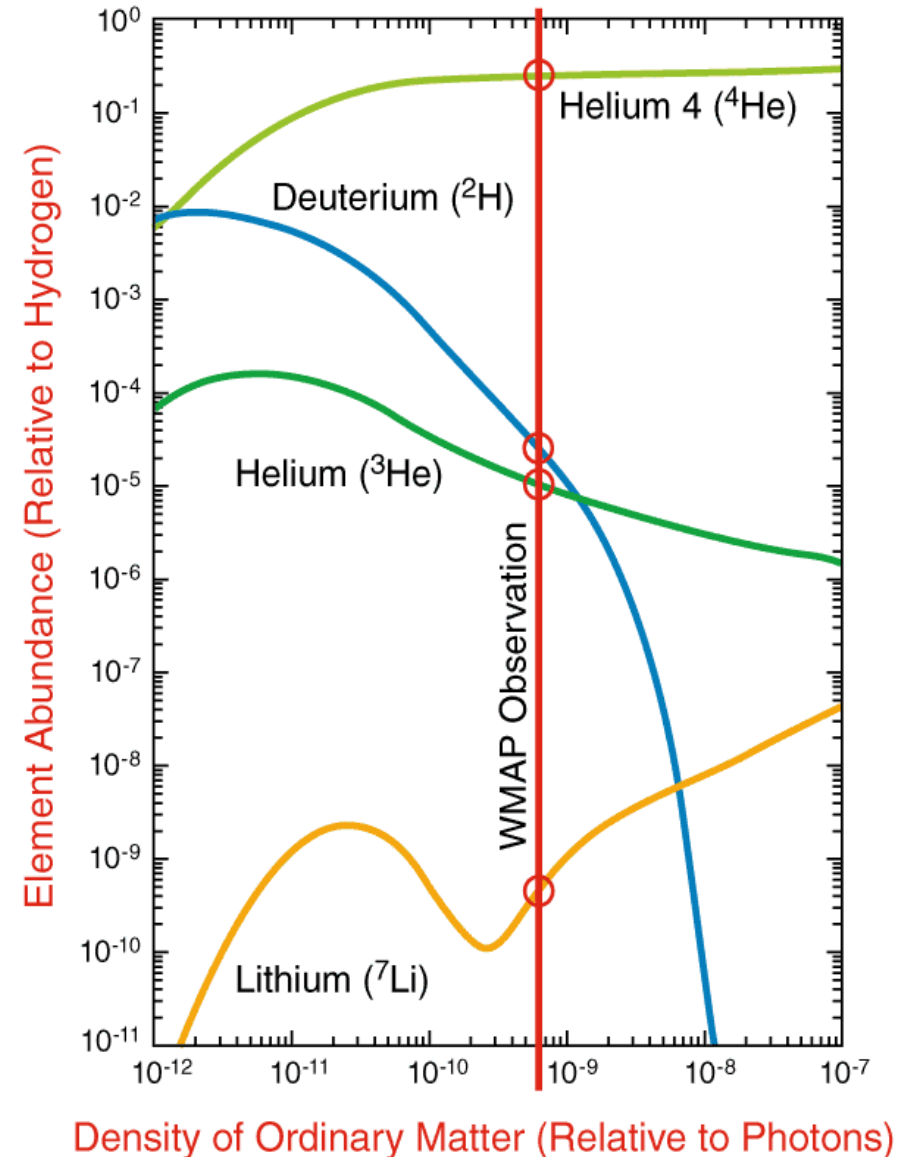
WIMP = “Weakly Interacting Massive Particle”



**New
Particles**

Why WIMPs?

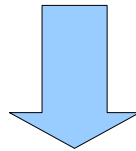
Too much Baryonic
Matter
↓
Wrong Primordial
Abundances



Why WIMPs?

Too much Baryonic

Matter



Wrong Primordial

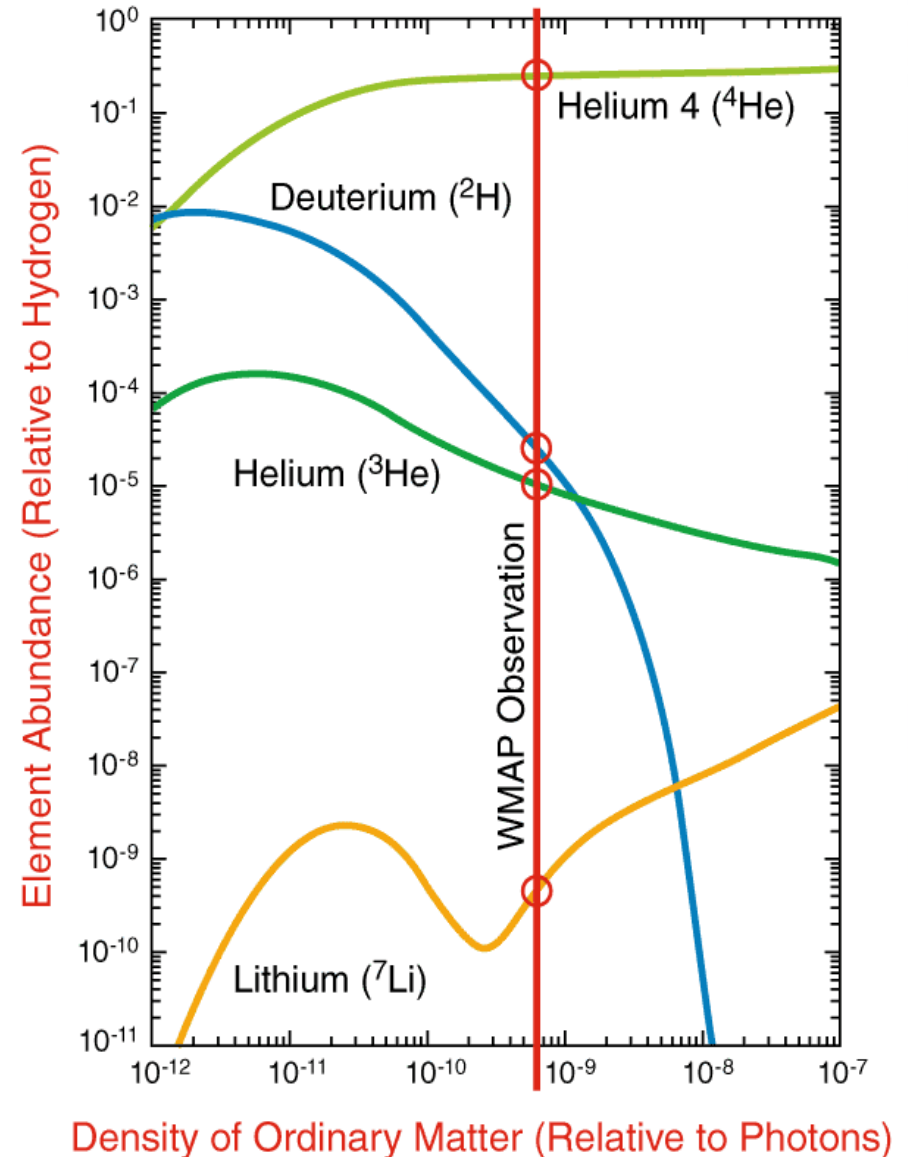
Abundances

Not:

Black Holes &

Non-Luminous Baryonic

Matter



Why WIMPs?

“WIMP Miracle”

Weak Scale Mass ($\sim 100\text{GeV}$)

+

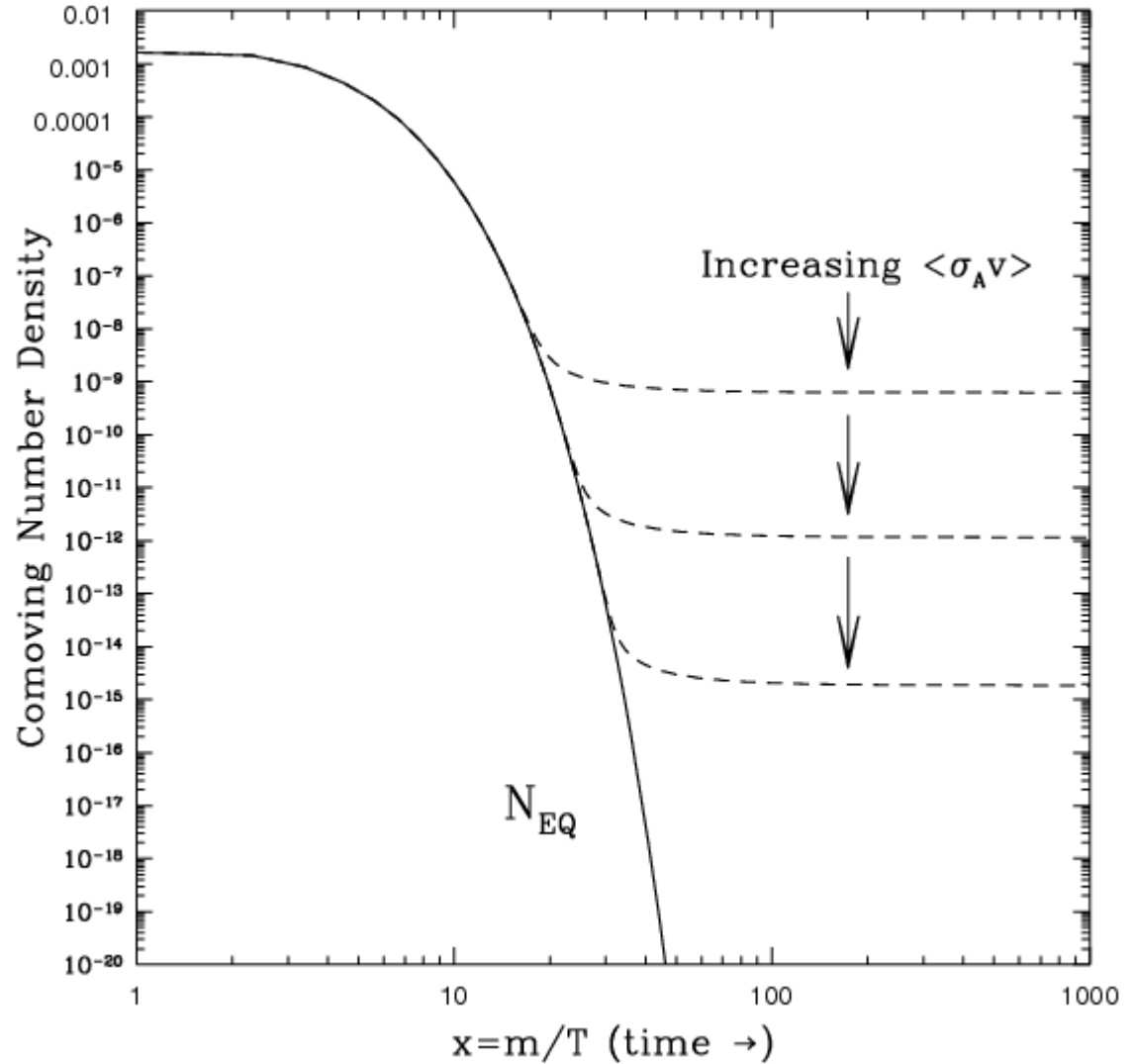
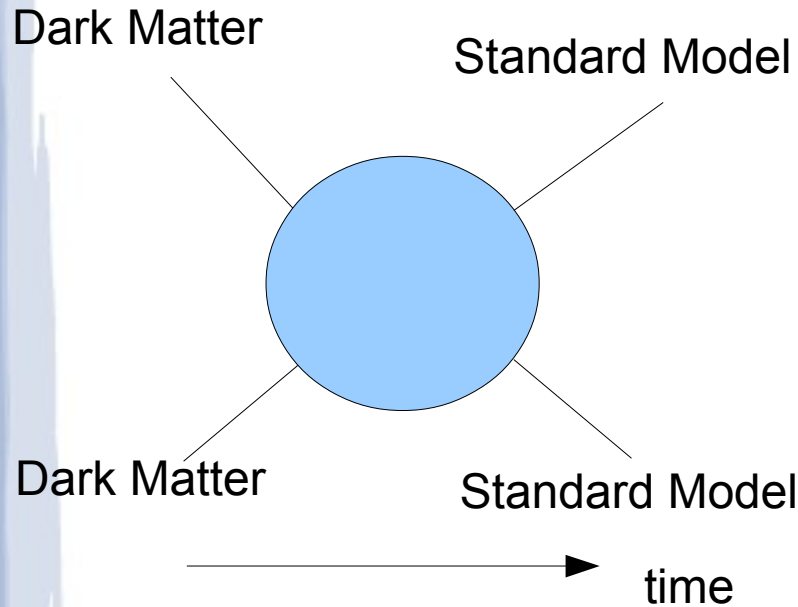
Weak Scale Coupling (~ 0.1)

\sim

Correct Abundance

Relic Abundance

Thermal Relics



Dark Matter Searches

1) Collider

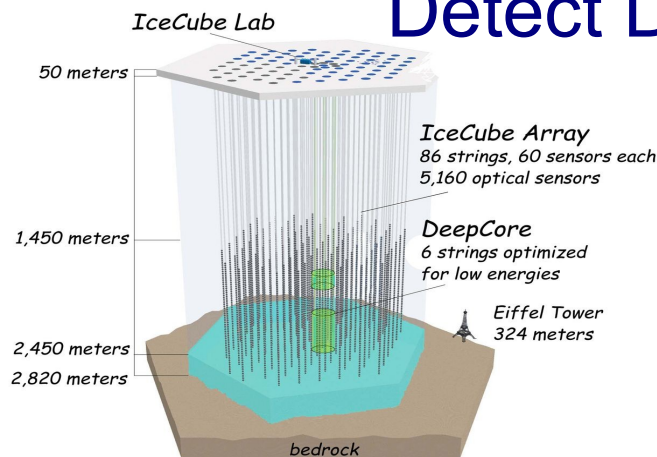
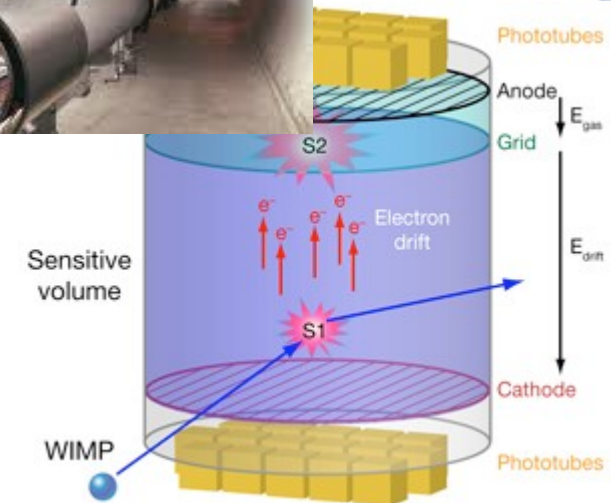
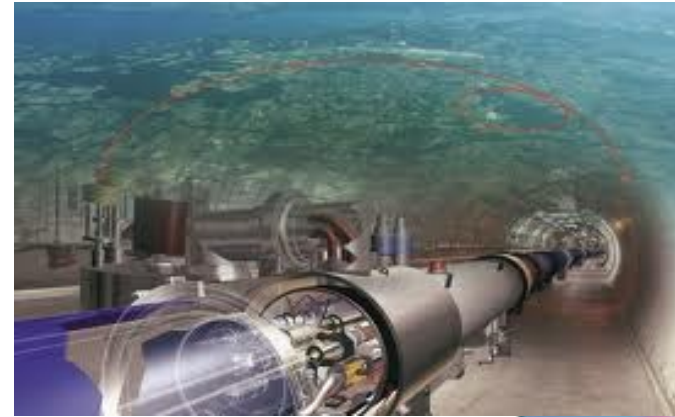
Make Dark Matter

2) Direct Detection

Detect Local Dark Matter

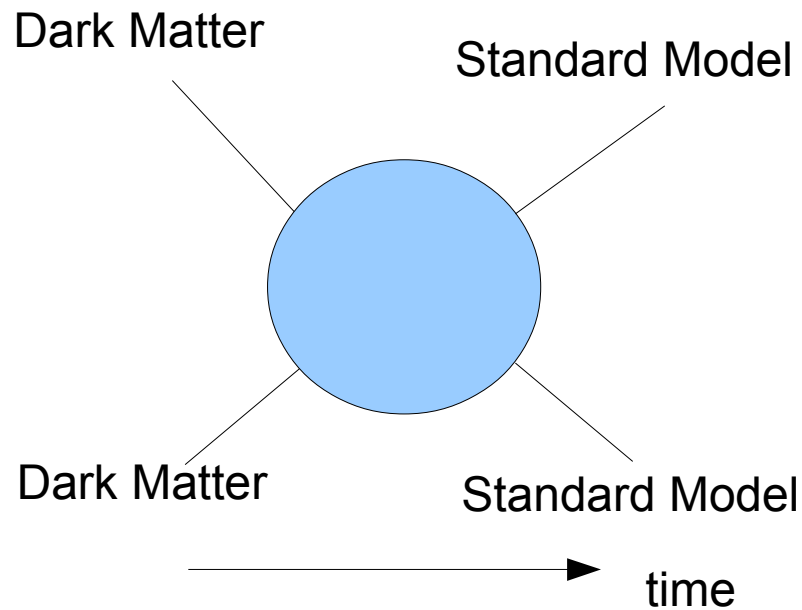
3) Indirect Detection

Detect Distant Dark Matter Annihilations



Indirect Detection

Dark Matter Annihilations

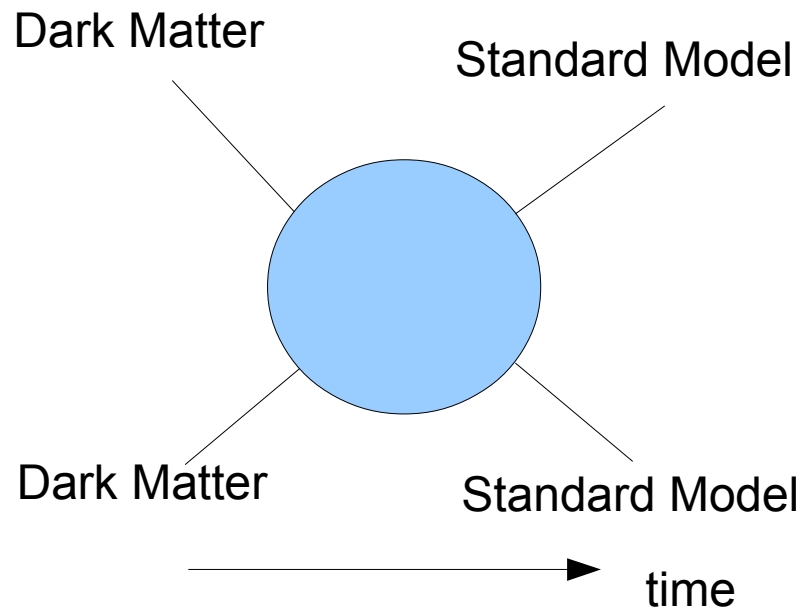


Same Basic Process
For Relic Abundance

Not Yet Sensitive
To Thermal Relics

Indirect Detection

Dark Matter Annihilations



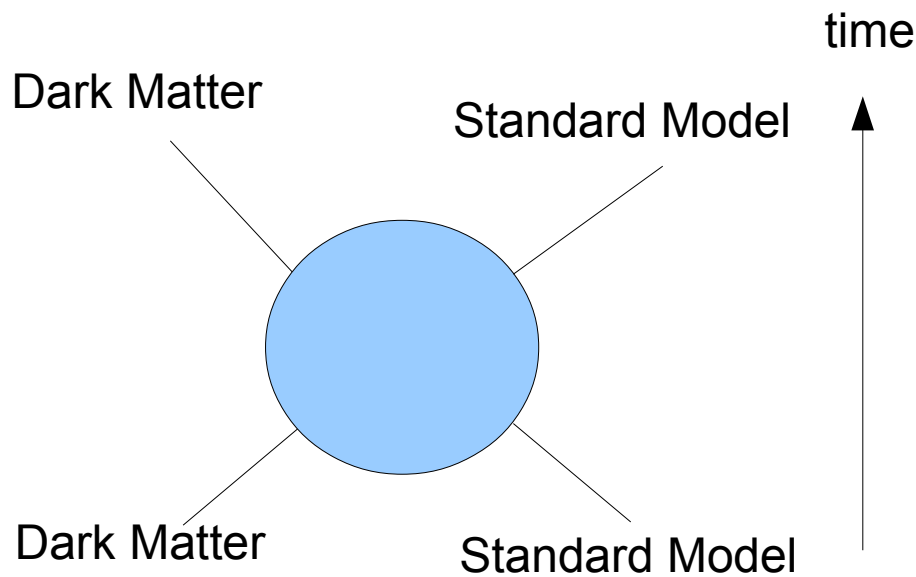
Same Basic Process
For Relic Abundance

Not Yet Sensitive
To Thermal Relics

Not Relevant

Direct Detection

Dark Matter – Nuclei Collisions



Spin Dependent:

Amplitude \sim Spin

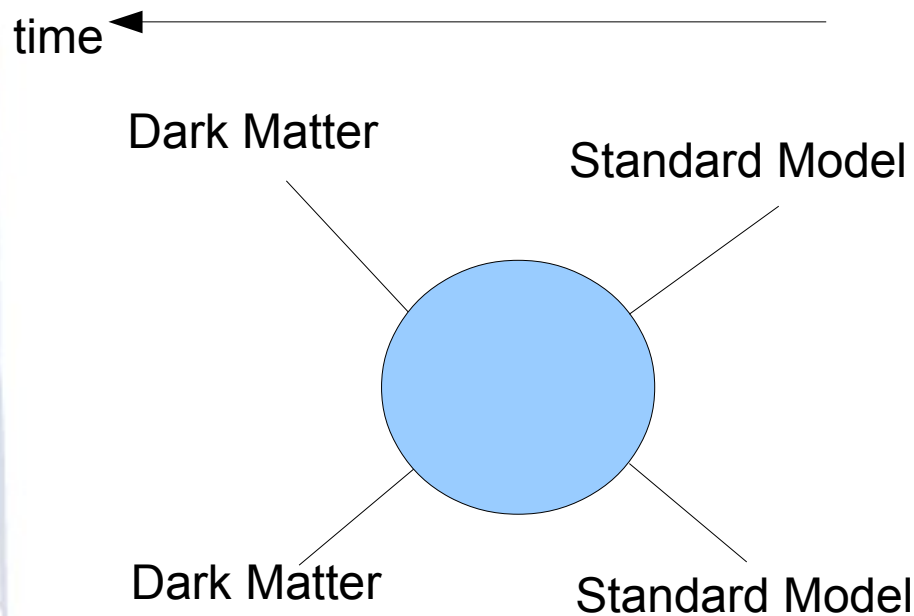
Spin Independent:

Amplitude Not \sim Spin

Cross Section $\sim A^2$

Collider Searches

Dark Matter Production



Dark Matter

=

Missing Energy (MET)

Monojet

(1 quark/gluon + MET)

&

Jets + MET

(2 quark/gluon + MET)

Dark Matter Searches

✓
1) Collider

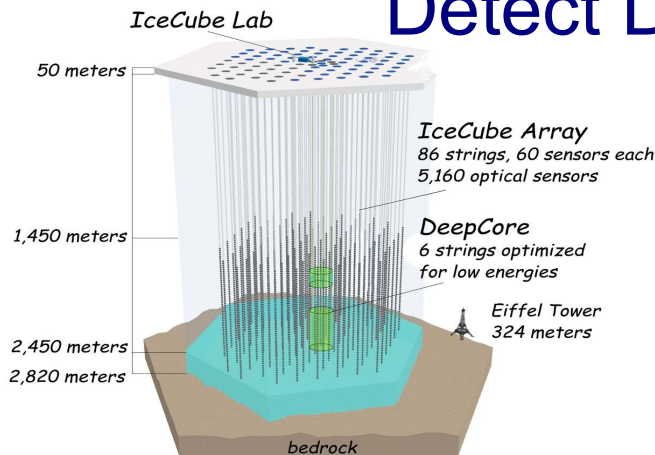
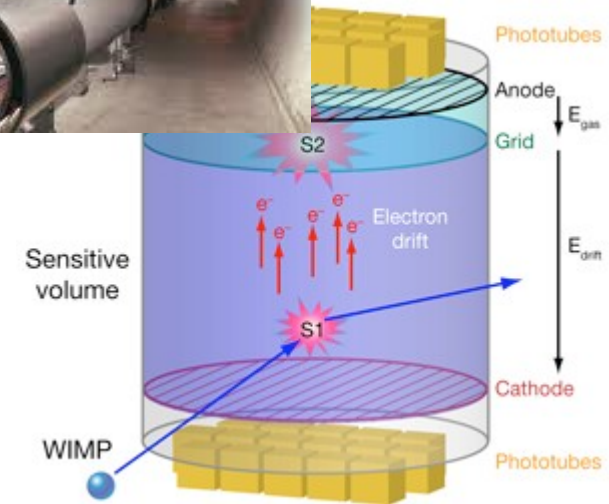
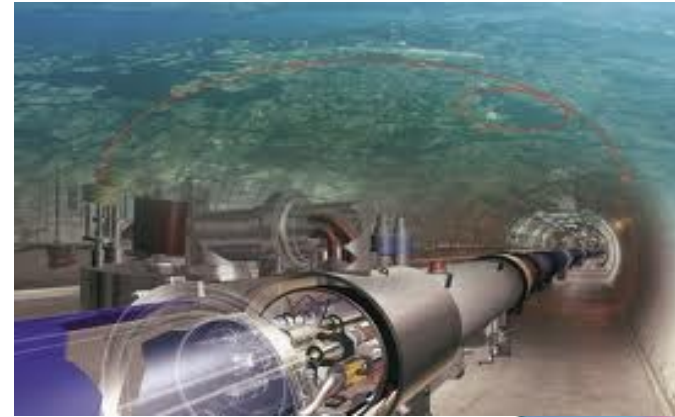
Make Dark Matter

✓
2) Direct Detection

Detect Local Dark Matter

✗
3) Indirect Detection

Detect Distant Dark Matter Annihilations



Our Models: Motivations

Preserve “WIMP Miracle”

Thermal Relic DM

Renormalizable Interactions

Minimal New Content

Generality

All Spin Combinations

Own Antiparticle vs. Not

Flavor Structures

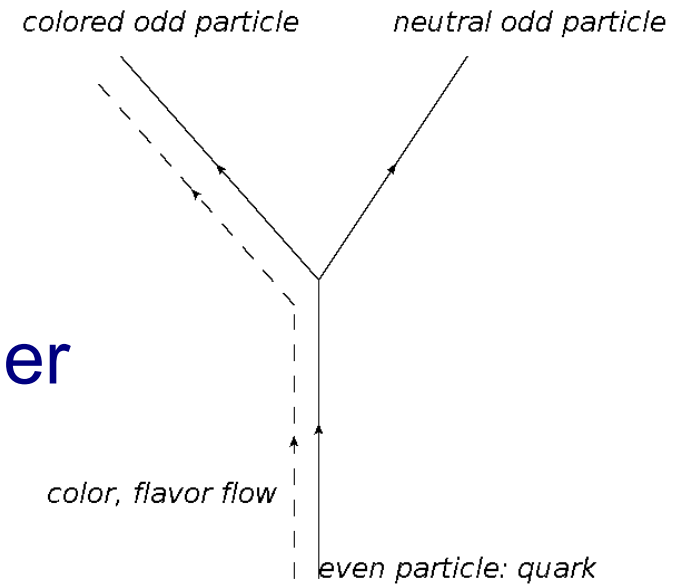
Our Models: Basics

2 New Particles:

χ , Dark Matter

Q, Standard Model Partner

1 New Interaction, λ



3 Parameters

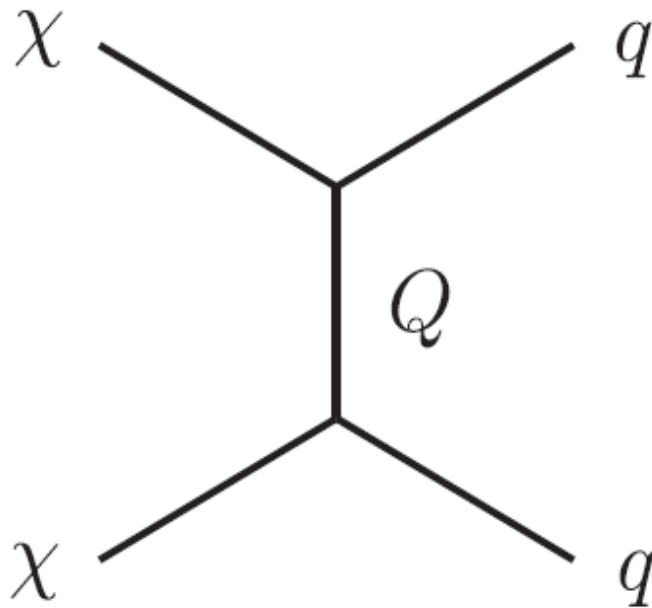


Relic Abundance

2 Parameters

Basic Processes

collider production



direct
detection

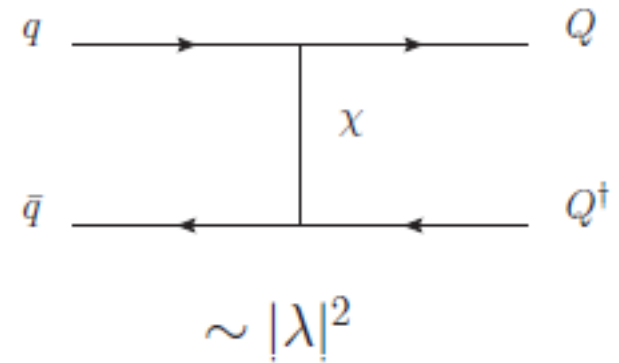
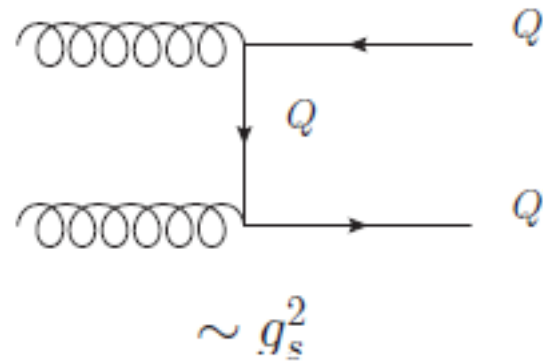
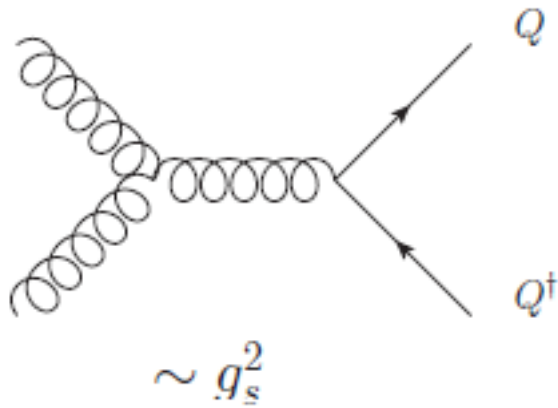


freeze out,
indirect detection

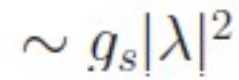
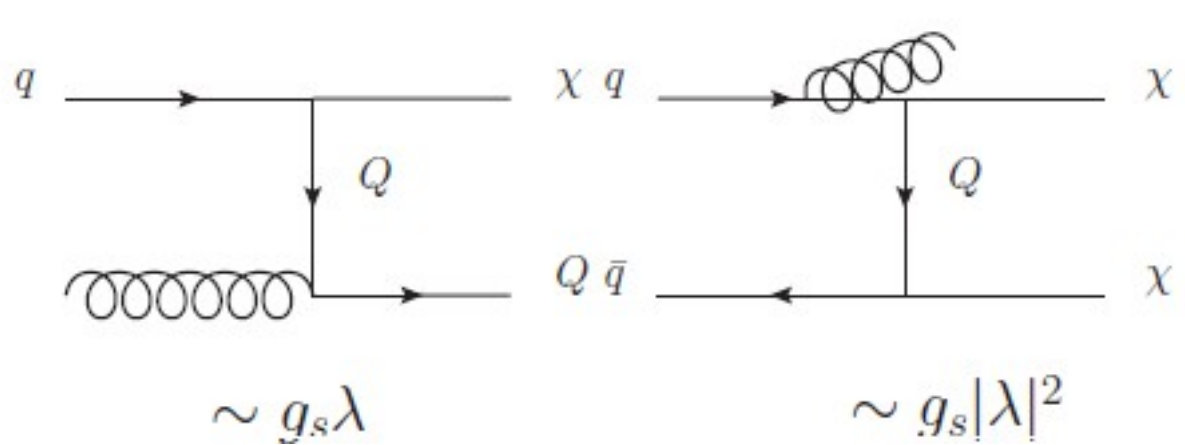
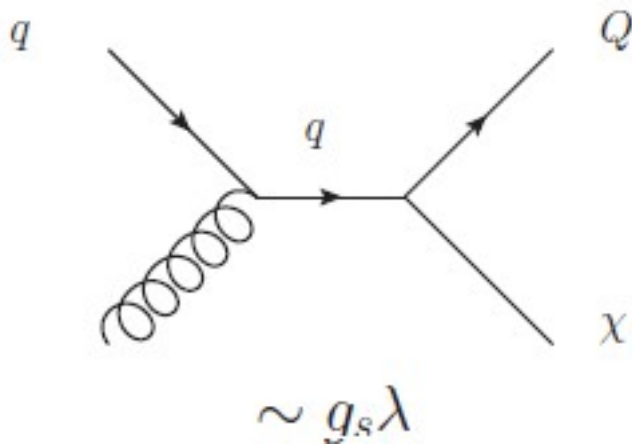


Collider Production

Jets+MET Production



Monojet Production



Our Model: Variations

	Spin		
	0	1/2	1
Own Antiparticle	χ : Real Scalar Q: Dirac Fermion	χ : Majorana Fermion Q: Complex Scalar	χ : Real Vector Q: Dirac Fermion
Not Own Antiparticle	χ : Complex Scalar Q: Dirac Fermion	χ : Dirac Fermion Q: Complex Scalar	χ : Complex Vector Q: Dirac Fermion

SM Particle = Quarks

All Generations

3 Flavor
Variations:

u d c s t b

2 Generations

3rd Generation

Our Model: Basic Features

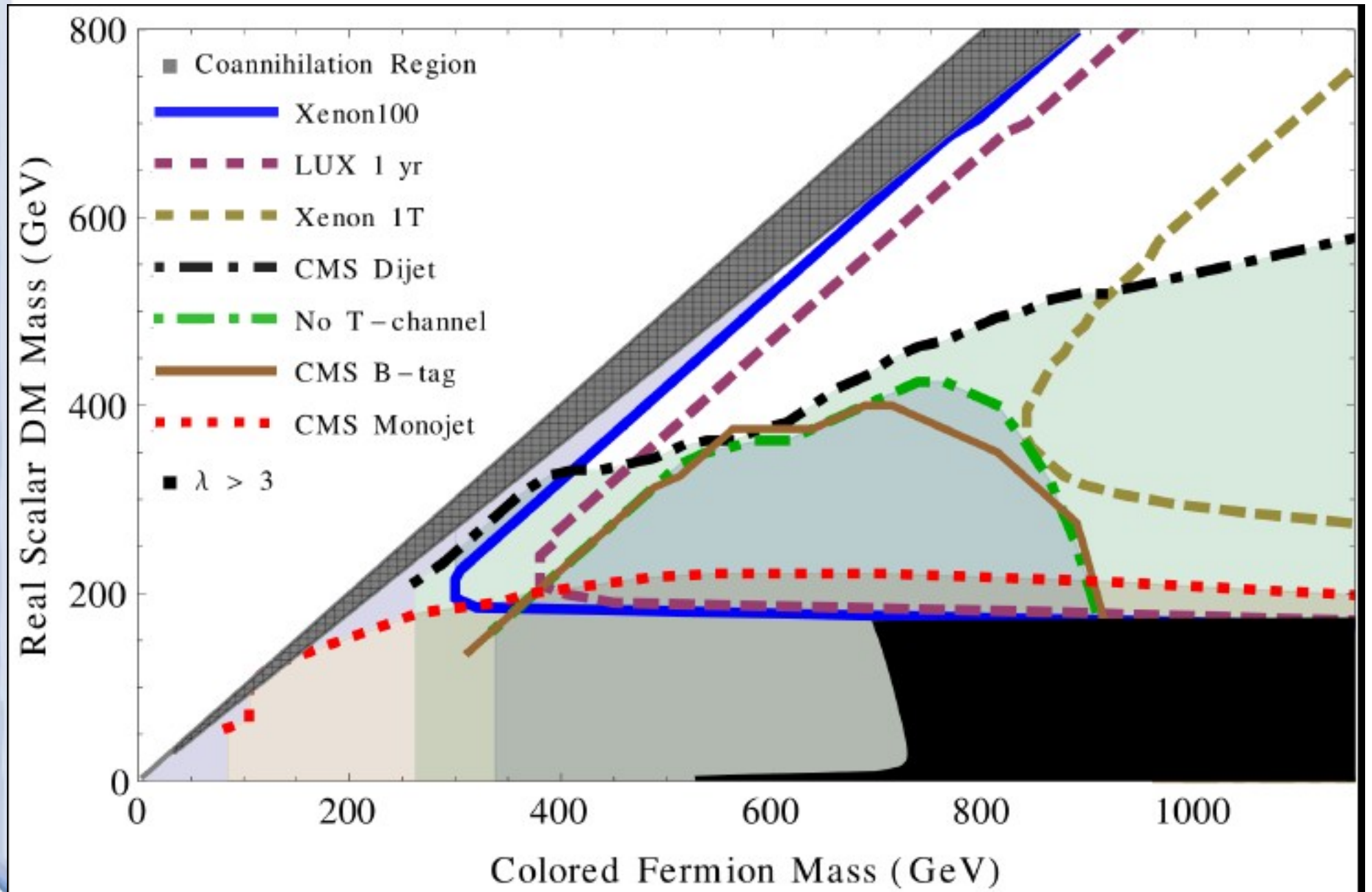
Enhanced Direct Detection Near Degeneracy

Larger λ Extends Limits to Larger m_Q

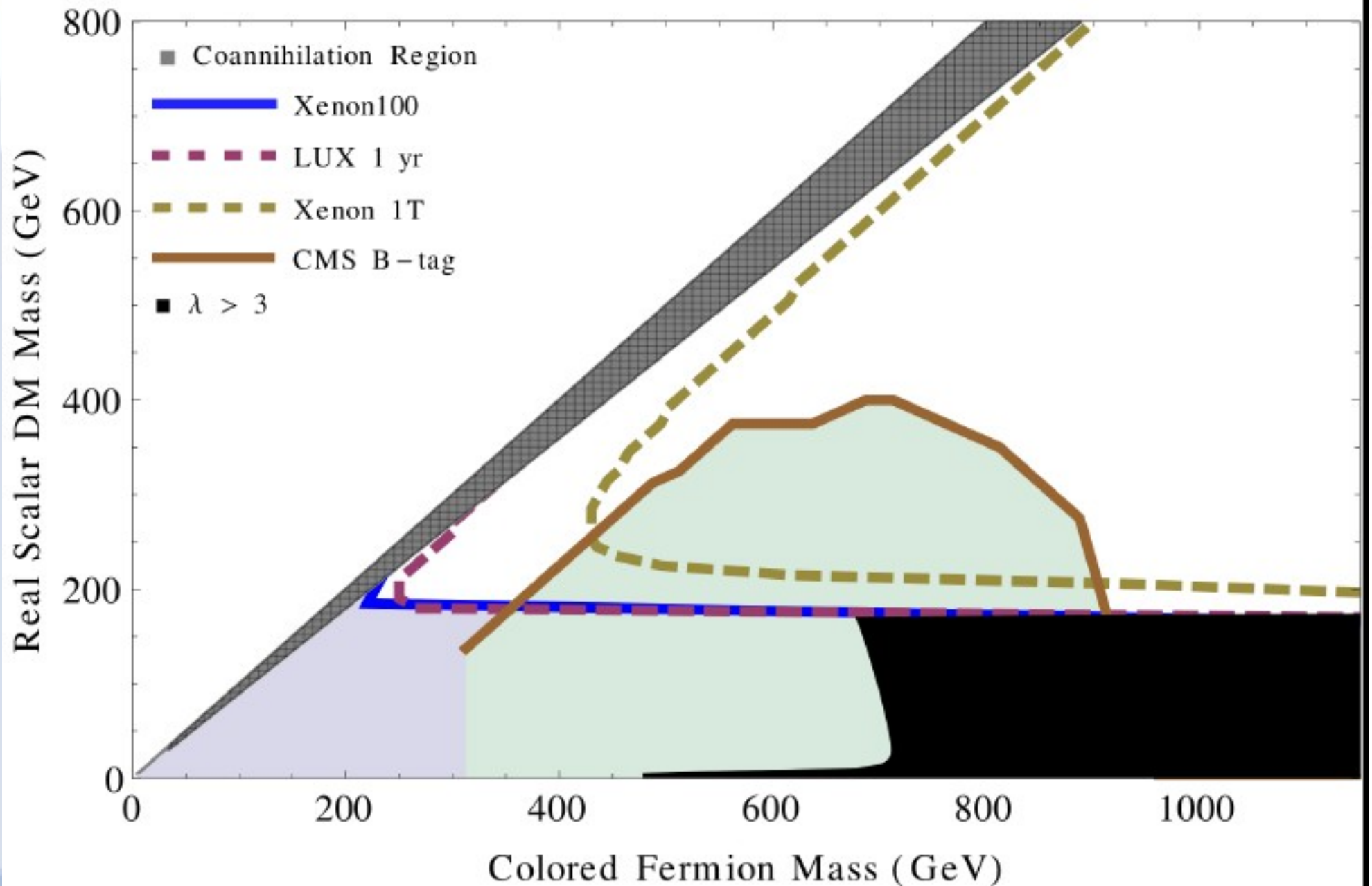
Direct Detection Rules Out:

Spin	0	1/2	1
Own Antiparticle	χ : Real Scalar Q: Dirac Fermion	χ : Majorana Fermion Q: Complex Scalar	χ : Real Vector Q: Dirac Fermion
Not Own Antiparticle	χ : Complex Scalar Q: Dirac Fermion	χ : Dirac Fermion Q: Complex Scalar	χ : Complex vector Q: Dirac Fermion

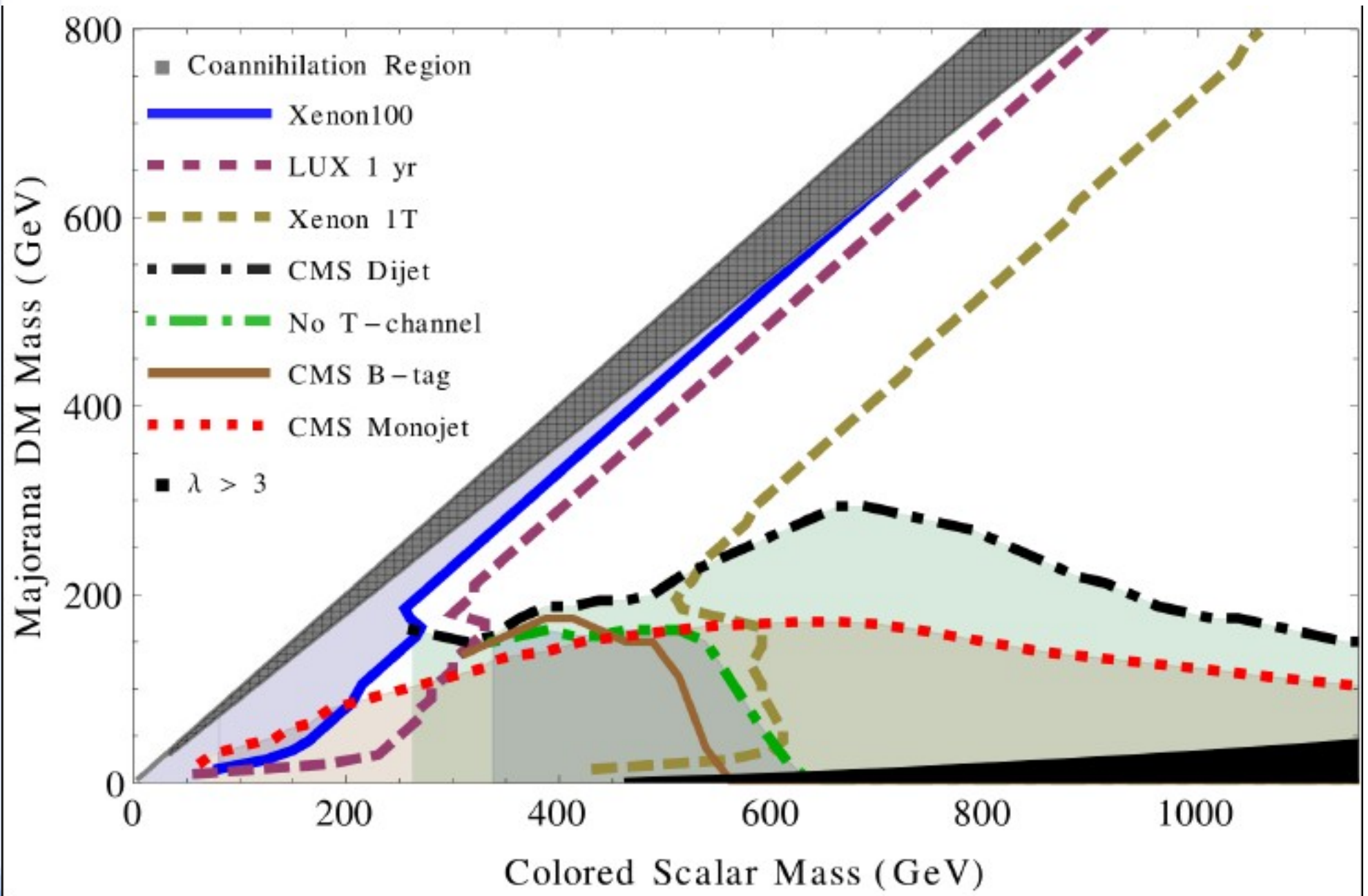
Real Scalar Dark Matter (All Gen.)



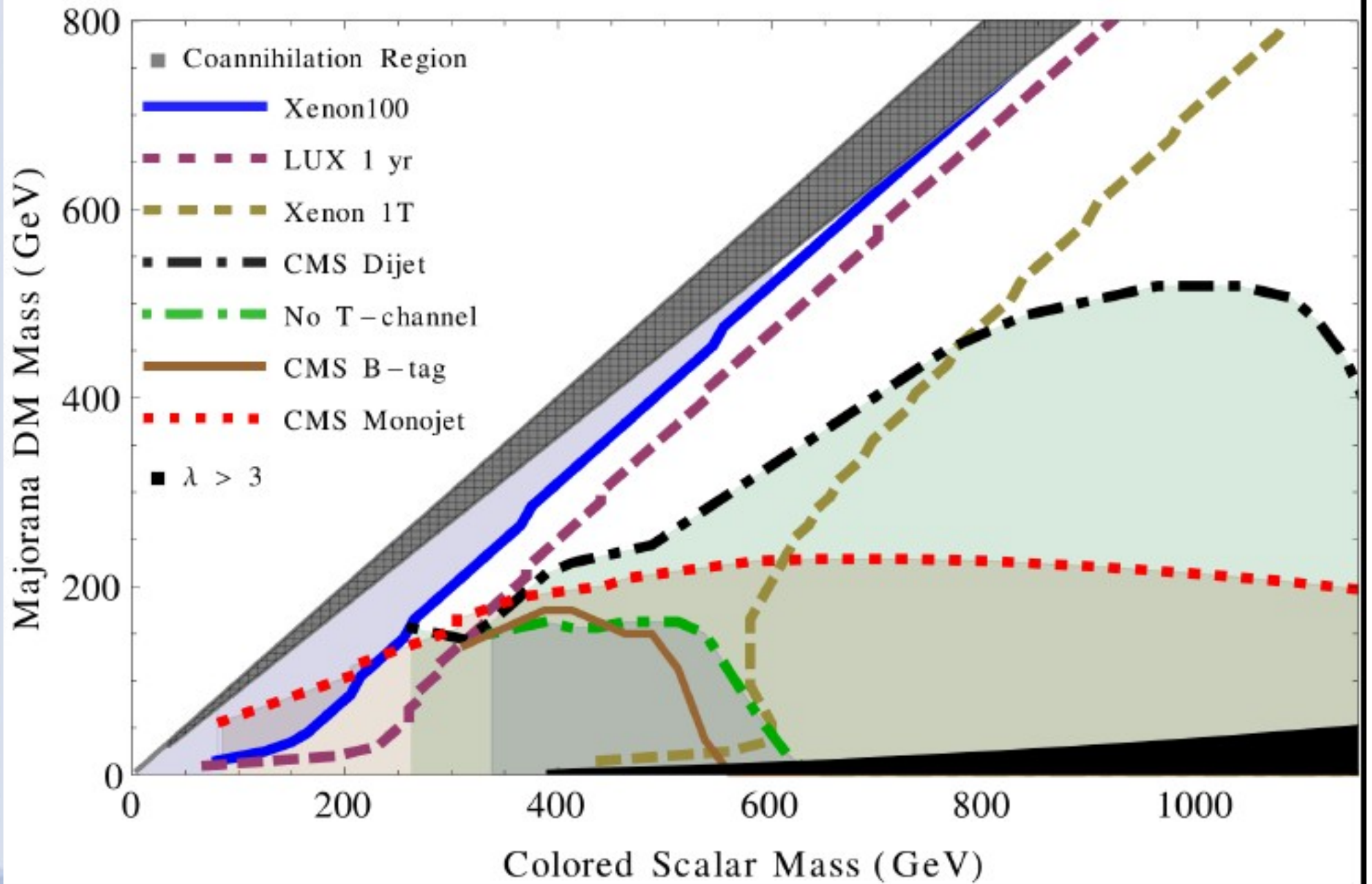
Real Scalar Dark Matter (3rd Gen.)



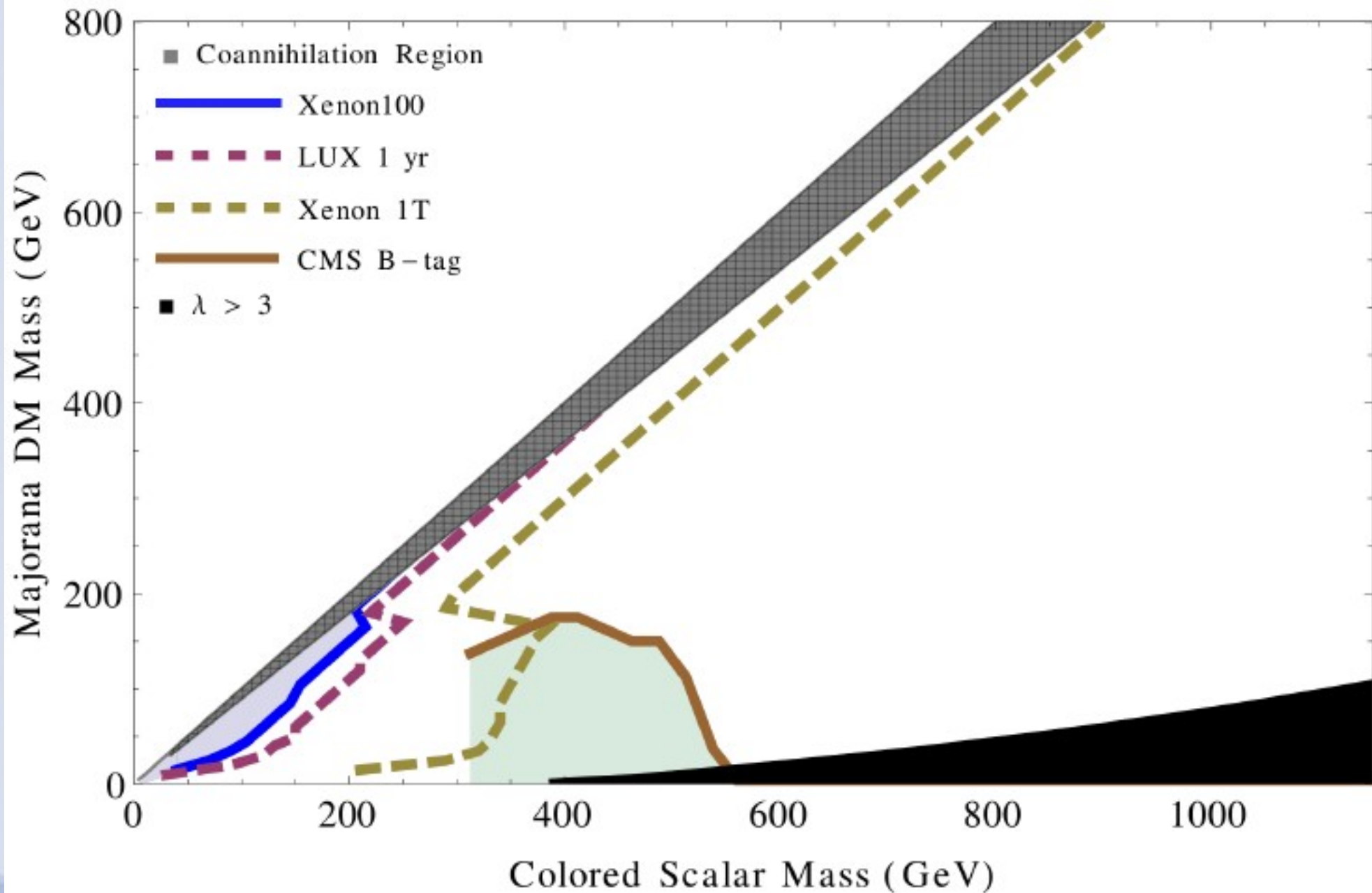
Majorana Dark Matter (All Gen.)



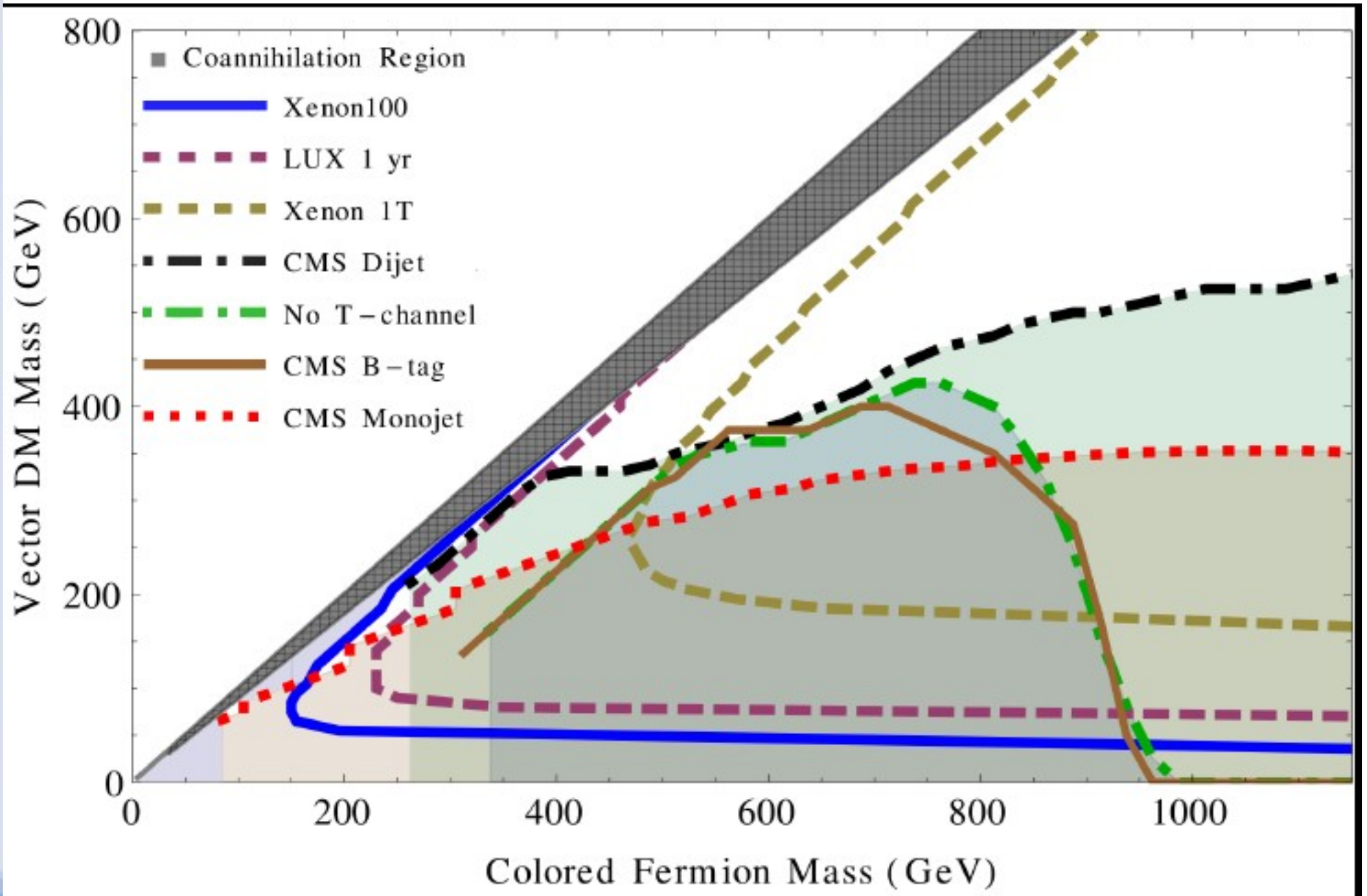
Majorana Dark Matter (2 Gen.)



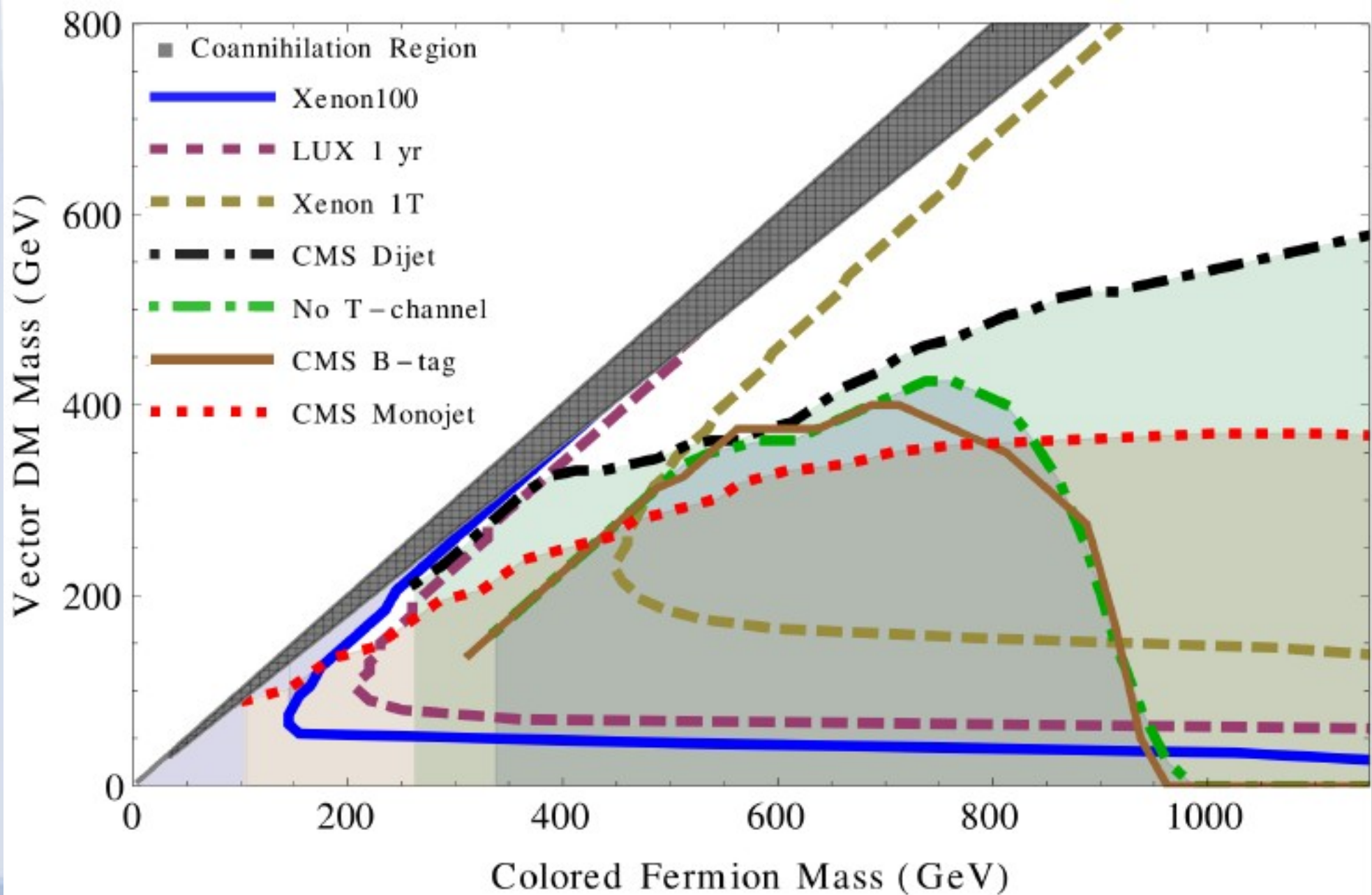
Majorana Dark Matter (3rd Gen.)



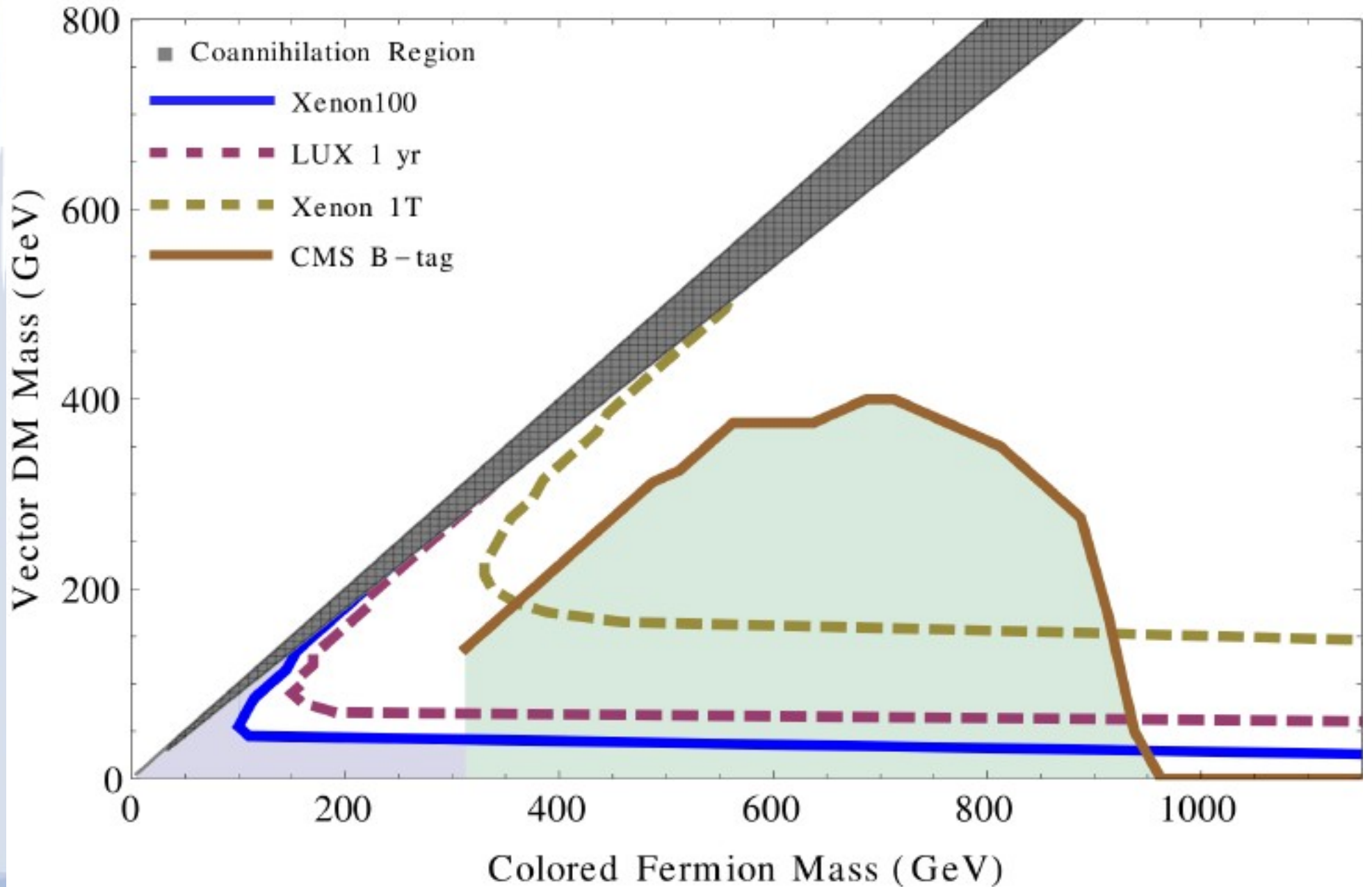
Real Vector Dark Matter (All Gen.)



Real Vector Dark Matter (2 Gen.)



Real Vector Dark Matter (3rd Gen.)



Summary

Simplified Models provide straightforward comparisons of collider and direct detection

Collider & Direct Detection are complementary

Dark Matter must be its own antiparticle to avoid direct detection

Jets + MET are more sensitive than Monojet searches to simple models of dark matter