Light DM Search with **NV Centers in Diamonds**

arXiv: 2302.12756

So Chigusa

SC, M. Hazumi, D. Herbschleb, N. Mizuochi, K. Nakayama





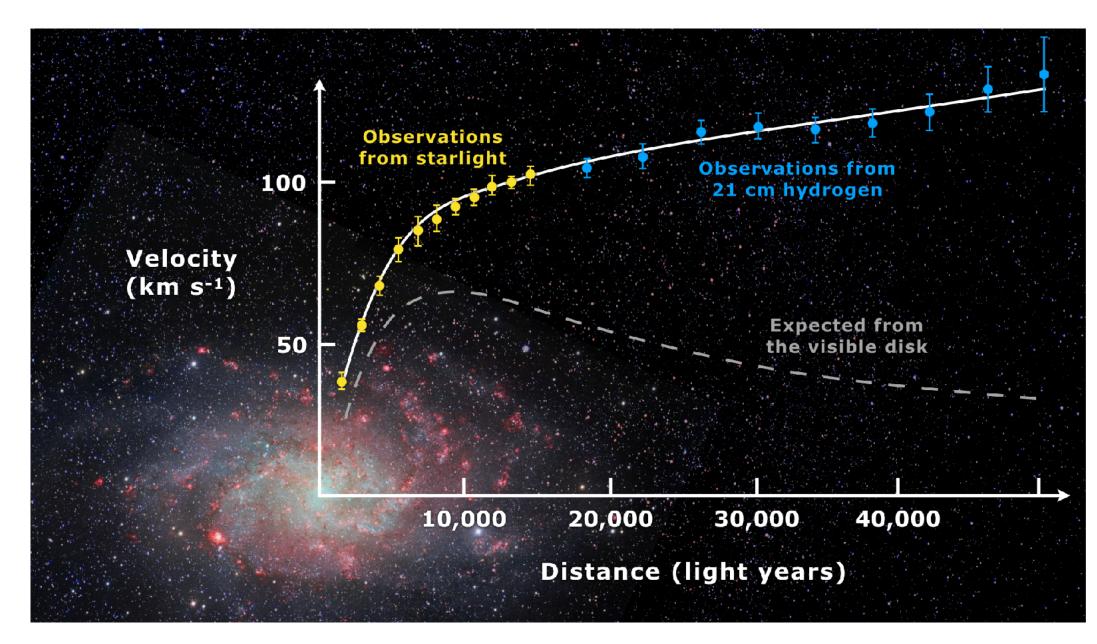
International Center for Quantum-field Measurement Systems for Studies of the Universe and Particles WPI research center at KEK

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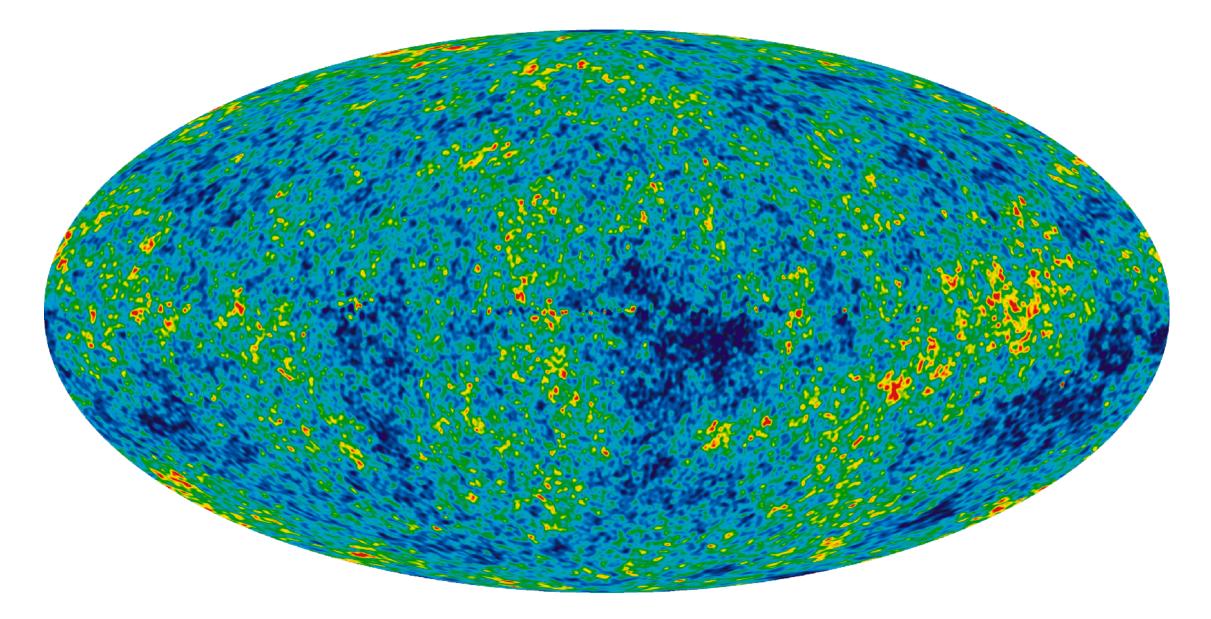
Dark Matter as a hint of new physics



Wikipedia "Galaxy rotation curve", E. Corbelli, P. Salucci (2000)

"Known"

✓ DM existence, abundance Has gravitational interaction



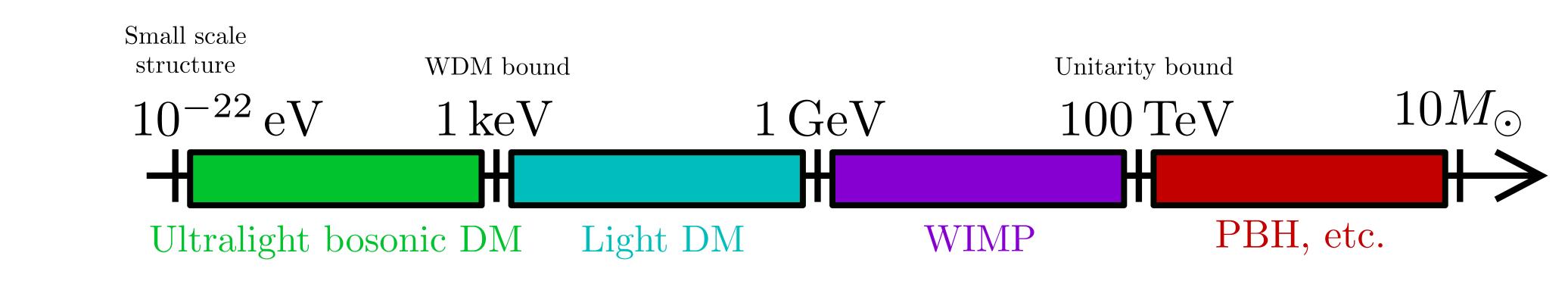
Wikipedia "Cosmic microwave background", 9 years of WMAP data

"Unknown"

✓ DM mass Von-gravitational interactions



DM mass window



- WIMP miracle w/ thermal production of $\mathcal{O}(1)$ TeV
- Other interaction and production mechanisms allow a broader mass range
 - freeze-in
 - misalignment mechanism for light bosonic DM

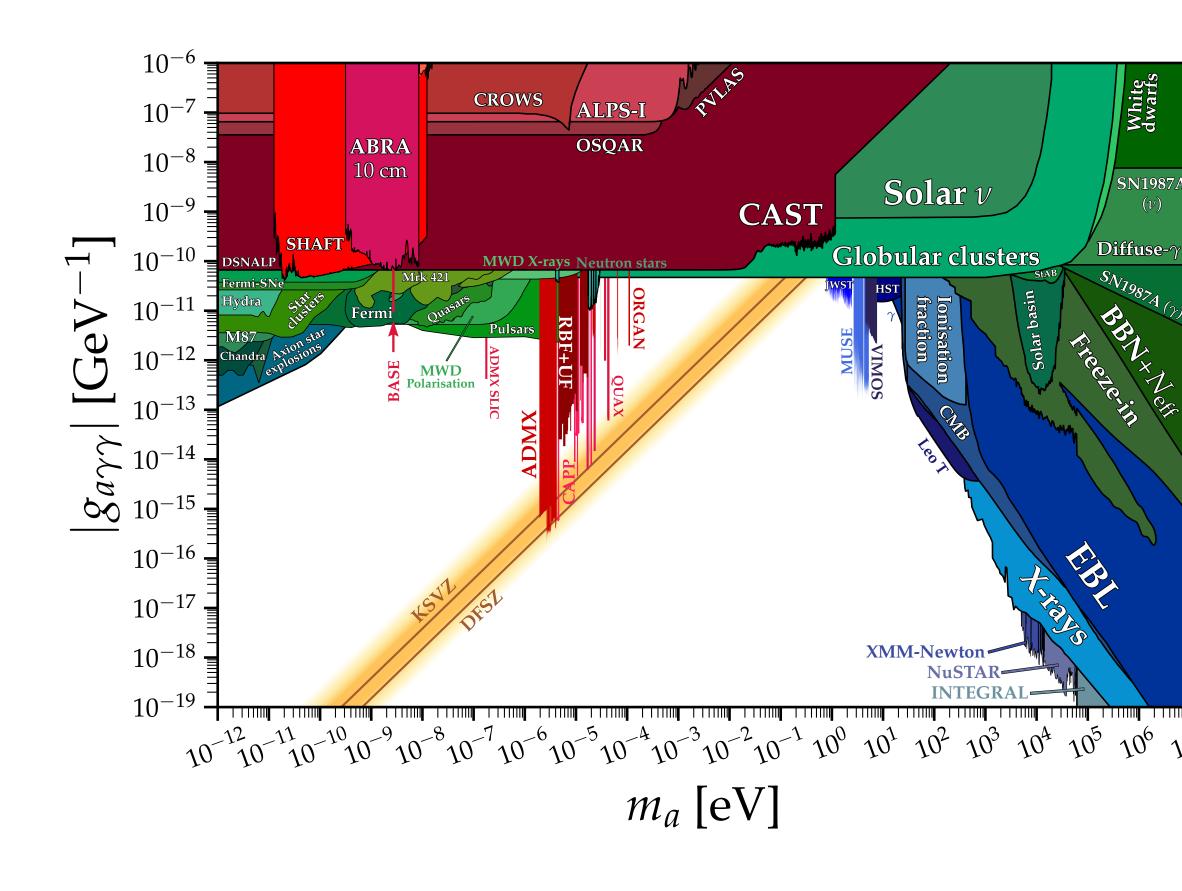
Nelson+ '11, Arias+ '12



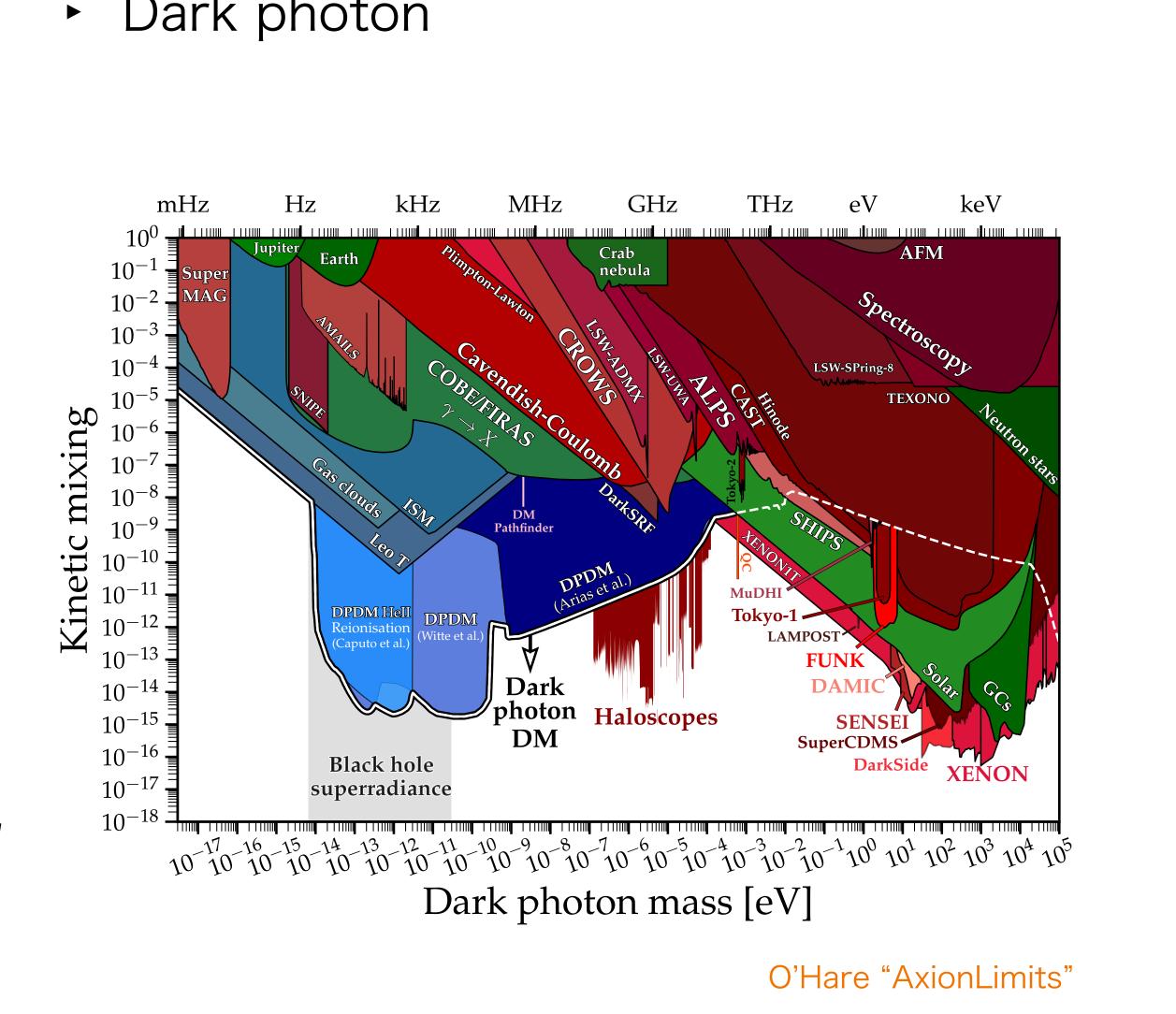


Light bosonic DM

- QCD axion
- Axion-like particles (ALPs)



Dark photon



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DM-induced electromagnetic field

QCD axion / axion-like particles (ALPs)

$$\mathscr{L} = g_{aff} \frac{\partial_{\mu} a}{2m_f} \bar{f} \gamma^{\mu} \gamma_5 f \rightarrow H_{\text{eff}} = \frac{g_{aff}}{m_f} \nabla a \cdot \mathbf{S}_f$$

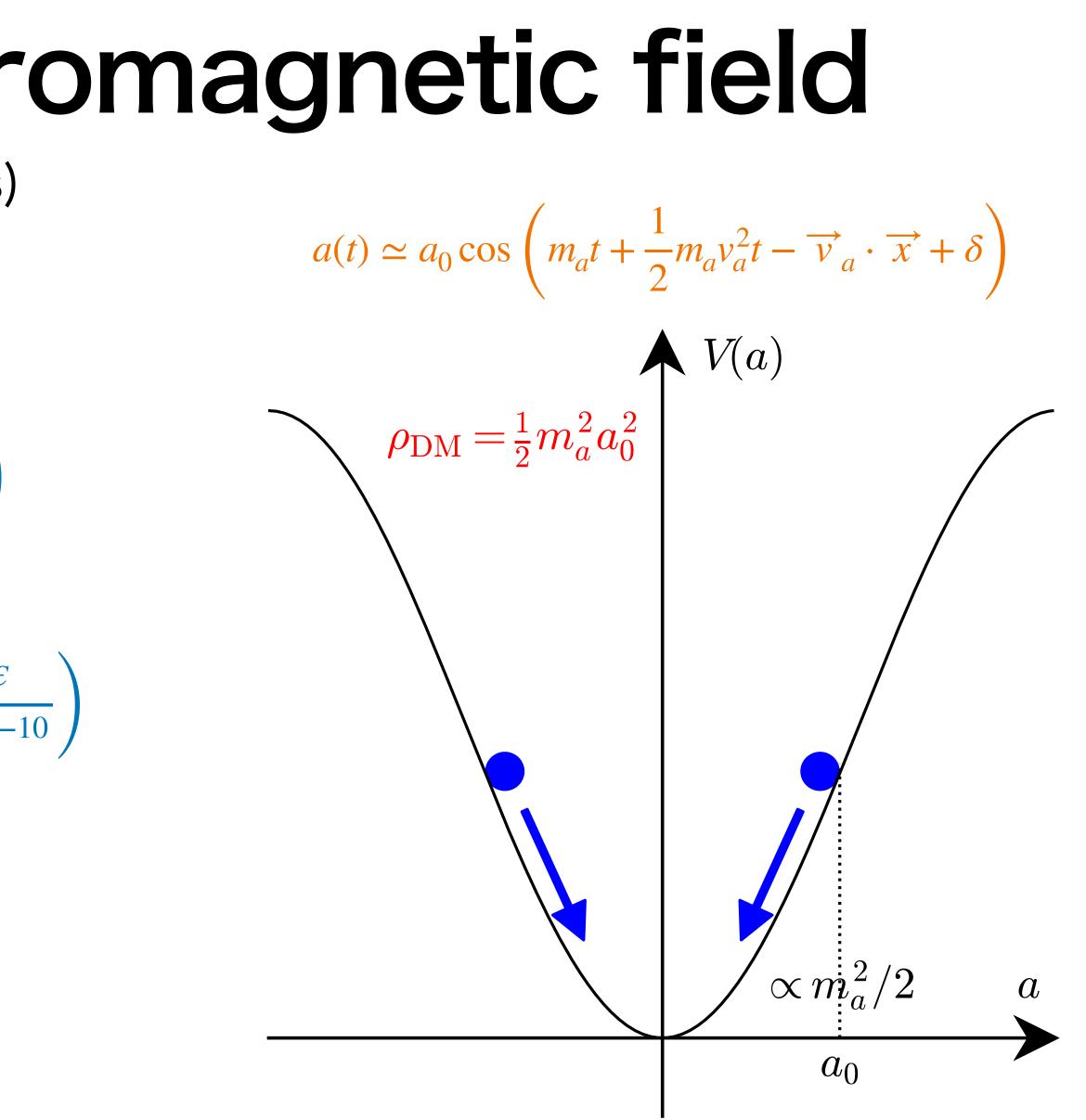
$$\mathbf{B}_{\text{eff}} \simeq \sqrt{2\rho_{\text{DM}}} \frac{g_{aff}}{e} \mathbf{v}_{\text{DM}} \cos(mt + \delta) \sim 3 \text{ aT} \left(\frac{g_{aff}}{10^{-10}}\right)$$

Dark photon

$$\mathbf{B}_{\text{eff}} = \sqrt{2\rho_{\text{DM}}} \,\epsilon \left(\mathbf{v}_{\text{DM}} \times \hat{H} \right) \cos(mt + \delta) \sim 1 \,\text{aT} \left(\frac{\epsilon}{10^{-10}} \right)$$

Behaves as an effective EM field with coherence time

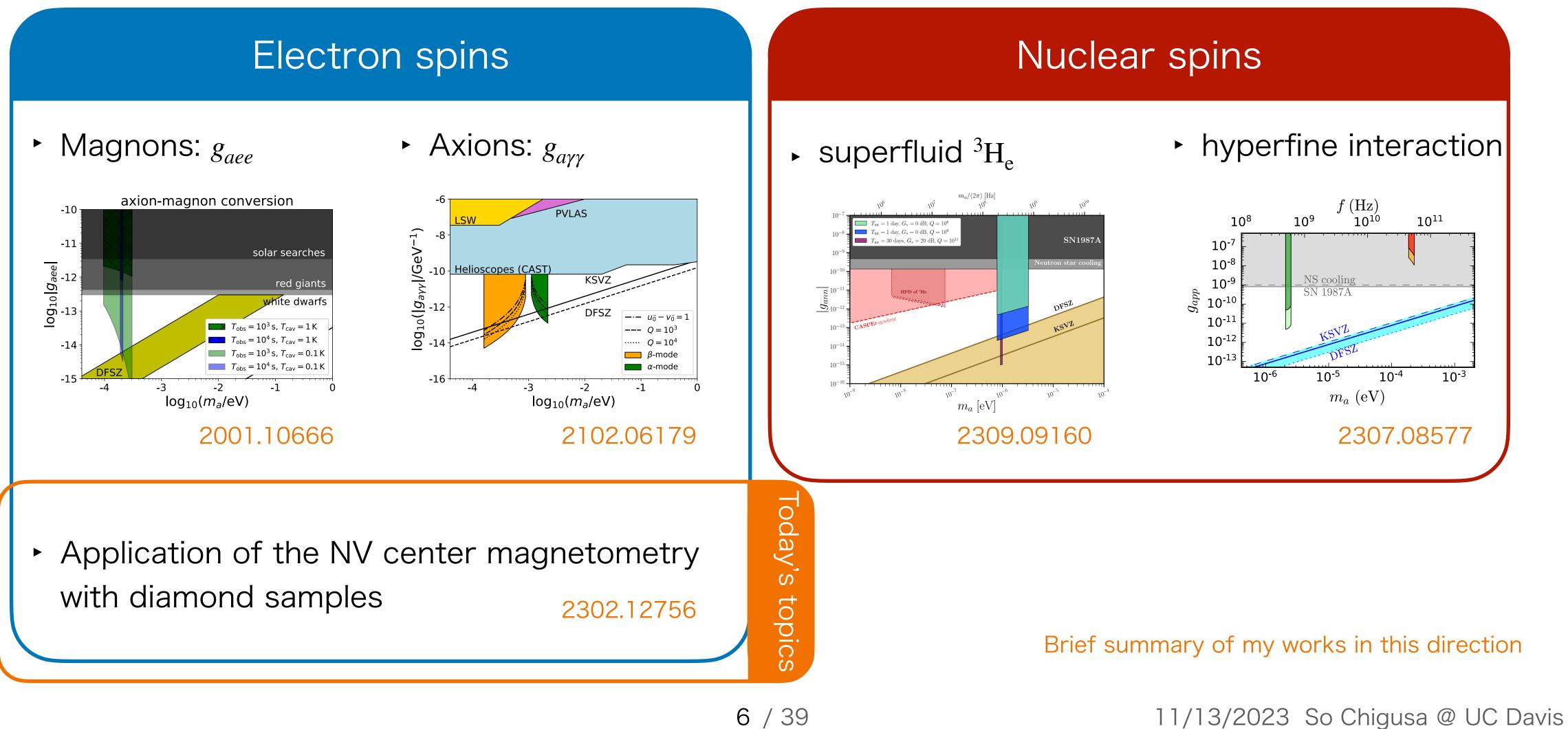
$$\tau_{\rm DM} = \frac{2\pi}{m_{\rm DM} v_{\rm DM}^2} \sim 6s \left(\frac{10^{-10} \,\mathrm{eV}}{m_{\rm DM}}\right)$$





Spin dynamics for DM search

Spin dynamics in various condensed matter systems can be used





Magnon as a DM signal Light bosonic DM converts into a collective excitation of spin = magnon

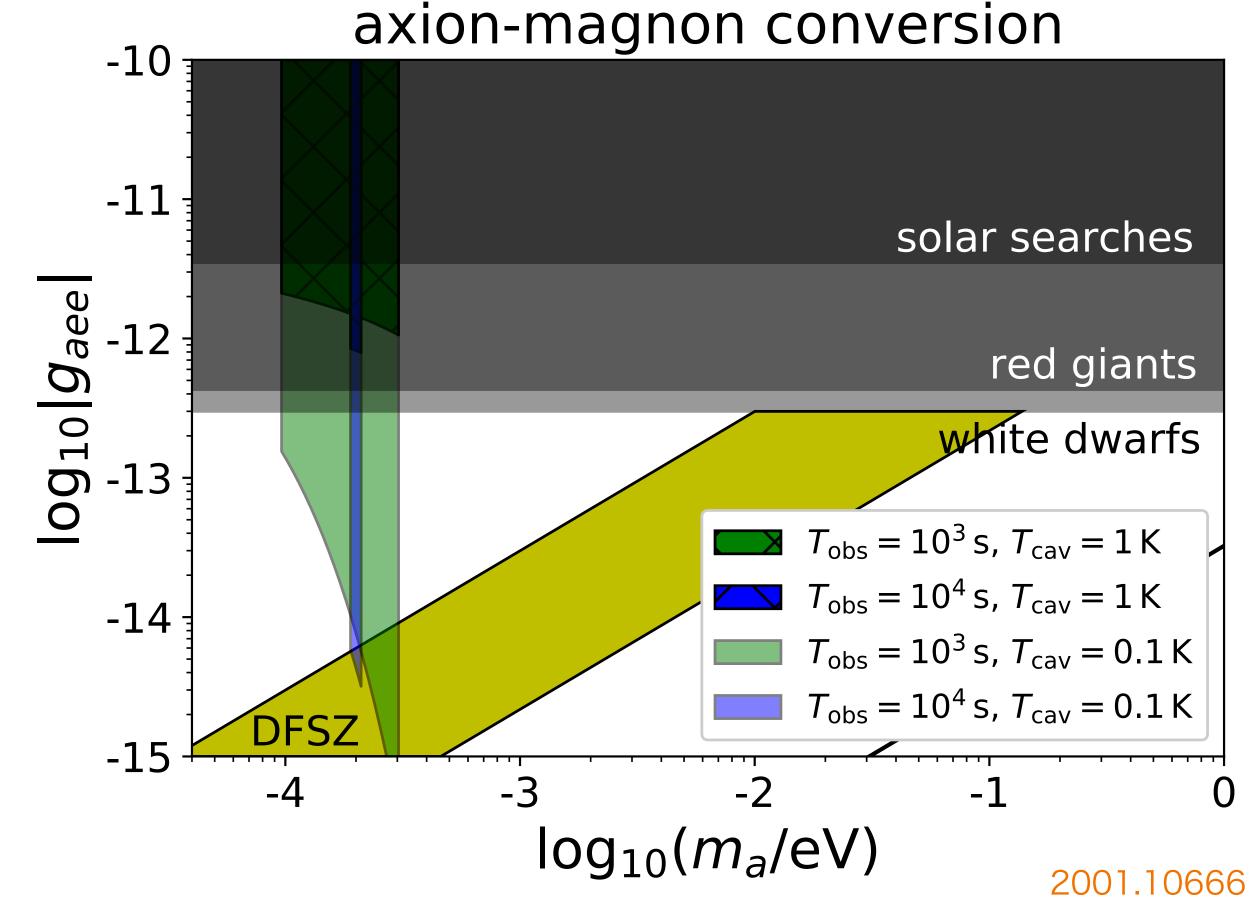




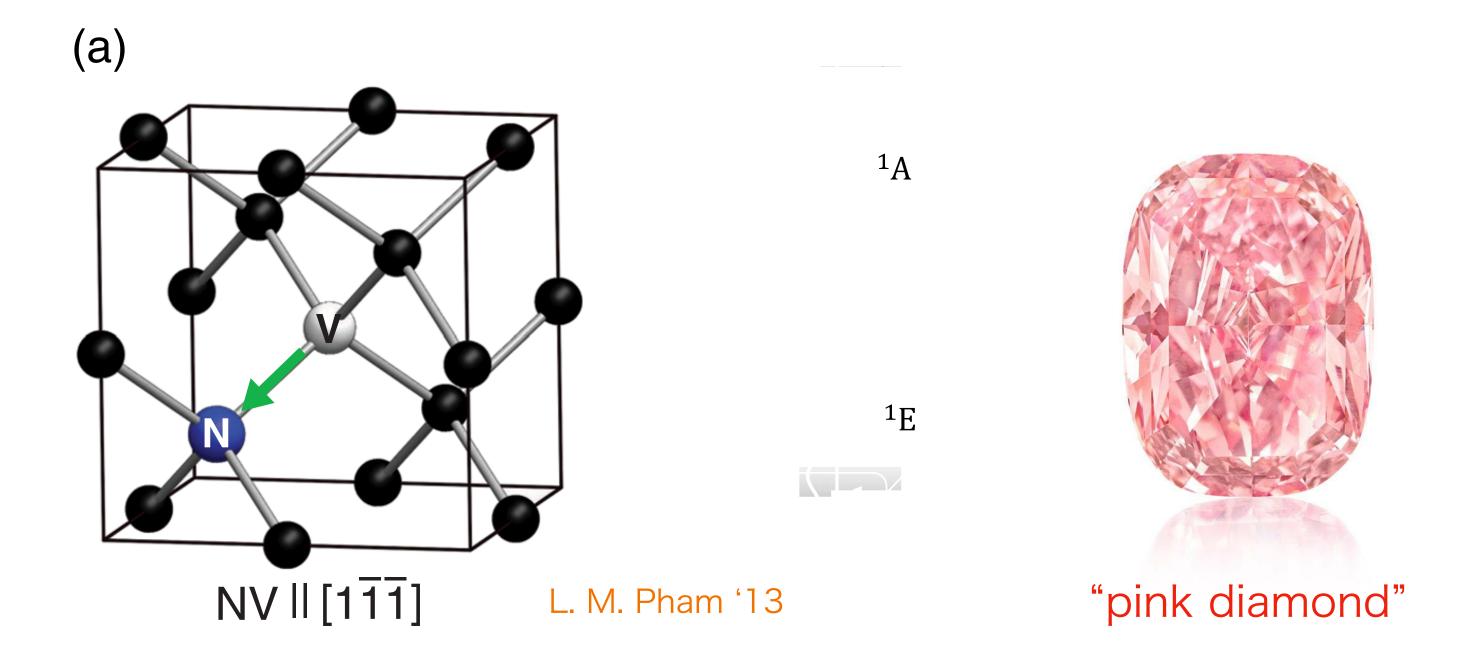
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- Introduction to NV center
 - What is it? How does it work as a quantum sensor?
- NV center magnetometry for DM detection
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 - Shielding effect for dark photon DM
- Experimental status
- Conclusion



Introduction to NV center

NV center in diamond



- The stable complex of substitutional nitrogen (N) and vacancy (V) in diamond
- The charged state NV⁻ has two extra e⁻s localized at V
- The ground state: e^- orbital singlet, e^- spin triplet S = 1 system

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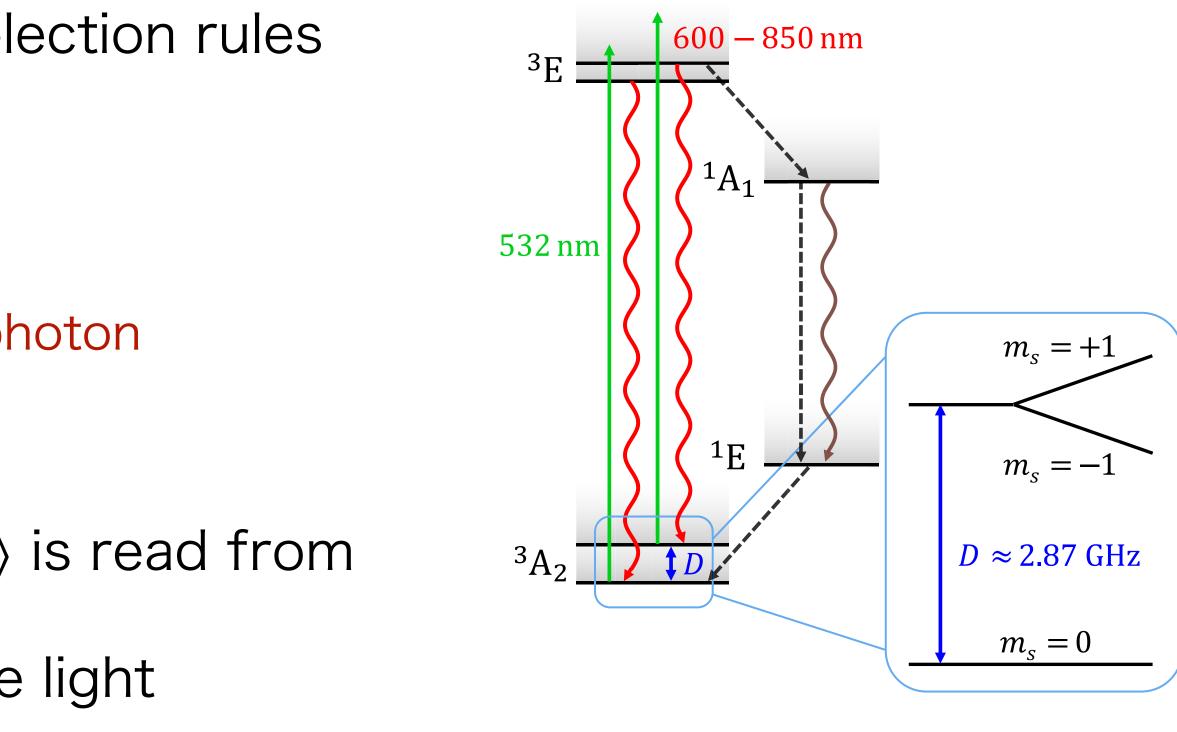
Fluorescence

- Governed by following processes + selection rules

• ${}^{3}A_{2}$ + 532 nm photon $\rightarrow {}^{3}E$

- ${}^{3}E \rightarrow {}^{3}A_{2} + 600 850$ nm photon
- ${}^{3}E_{S\neq 0} \rightarrow ({}^{1}A_{1} \rightarrow {}^{1}E) \rightarrow |m_{s} = \pm \rangle + \text{infrared photon}$
- The spin state $|\psi\rangle = \cos\frac{\theta}{2}|0\rangle + \sin\frac{\theta}{2}|\pm\rangle$ is read from strength of the red (pink) fluorescence light

• Can distinguish spin states $|m_s = 0\rangle$ and $|m_s = \pm\rangle$ by fluorescence measurement

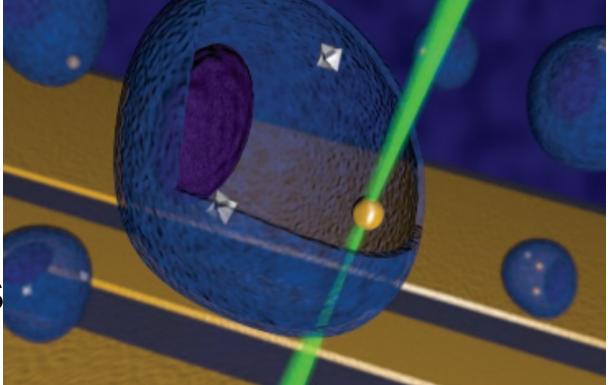


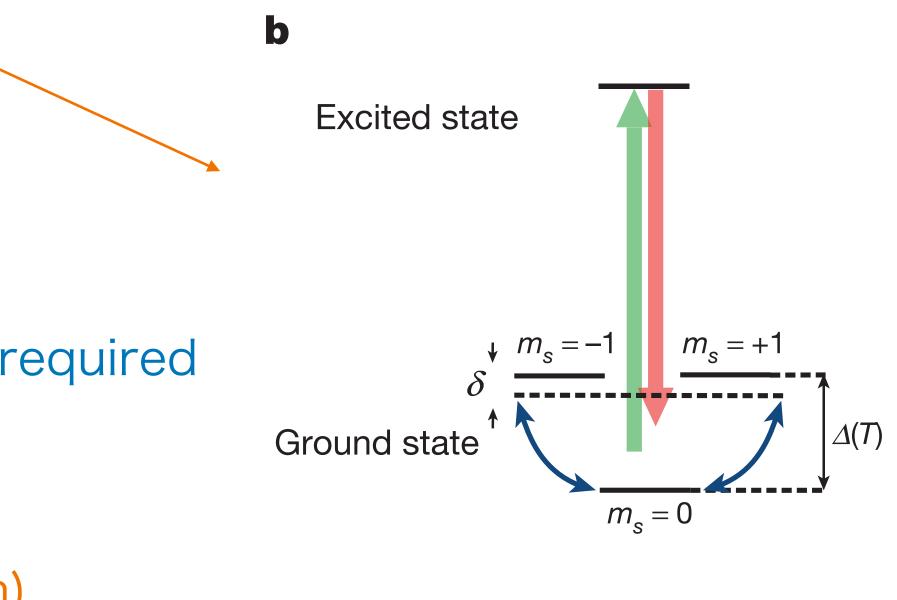
J. F. Barry+ '20



NV center as a quantu

- NV center works as a multimodal quantum sens
 - 1. Temperature G. Kucsko+ '13
 - 2. Electric field F. Dolde+ '11
 - 3. Strain M. Barson+ '17
 - 4. Magnetic field (explain later)
 - No cryogenics
 - No vacuum system
 - No tesla-scale applied bias fields are required
- Two options
 - Single NV center (high spacial resolution)
 - . Ensemble of NV centers (high sensitivity) with $\sim 1 20 \,\mathrm{ppm}$ concentration





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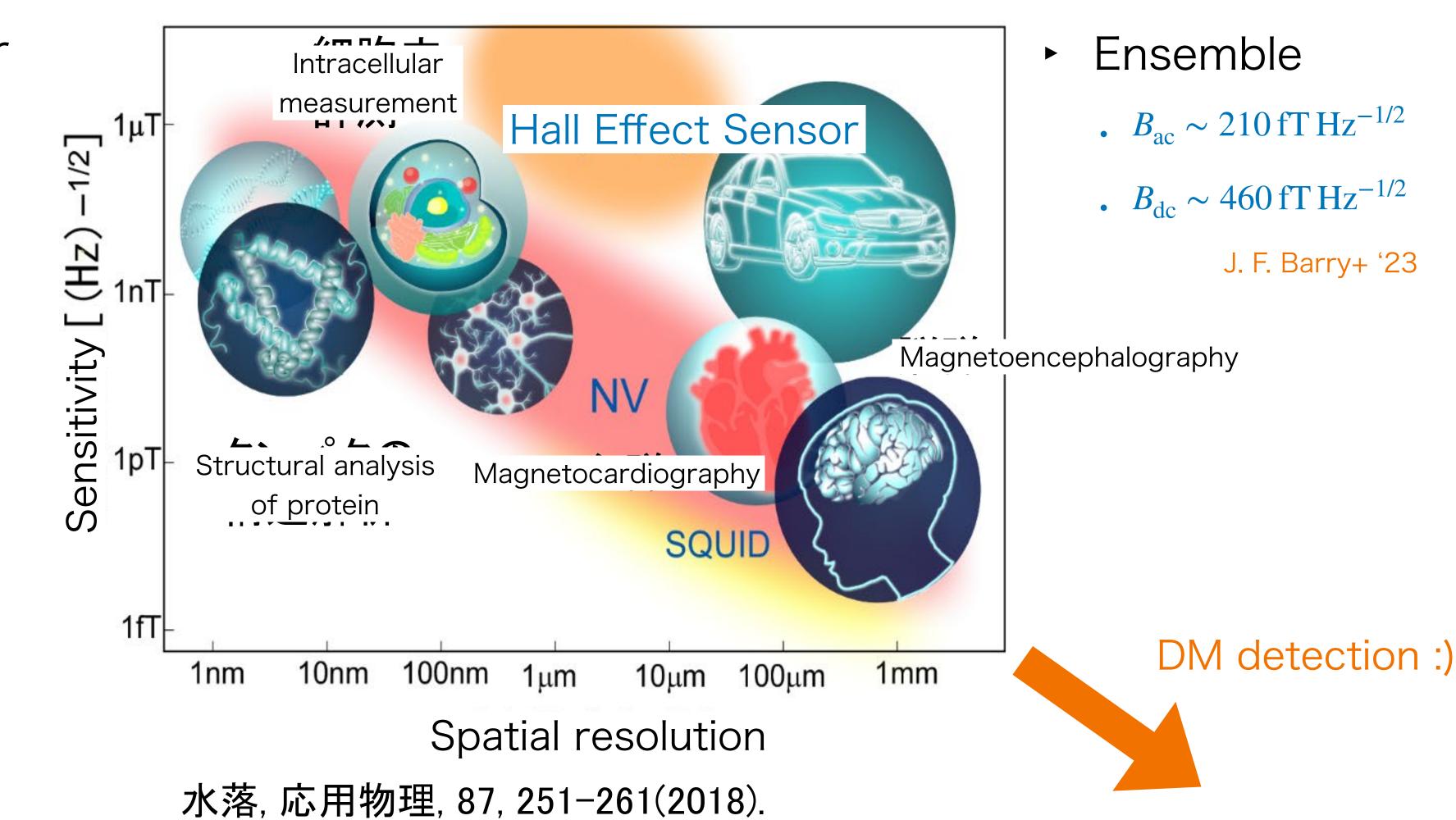


Applications of NV center magnetometory

- Single NV center
 - $B_{\rm ac} \sim 9.1 \,\mathrm{nT \, Hz^{-1/2}}$
 - $B_{\rm dc} \sim 10 \, \rm nT \, Hz^{-1/2}$

D. Herbschleb+ '19

 n_{NV}





DC magnetometry

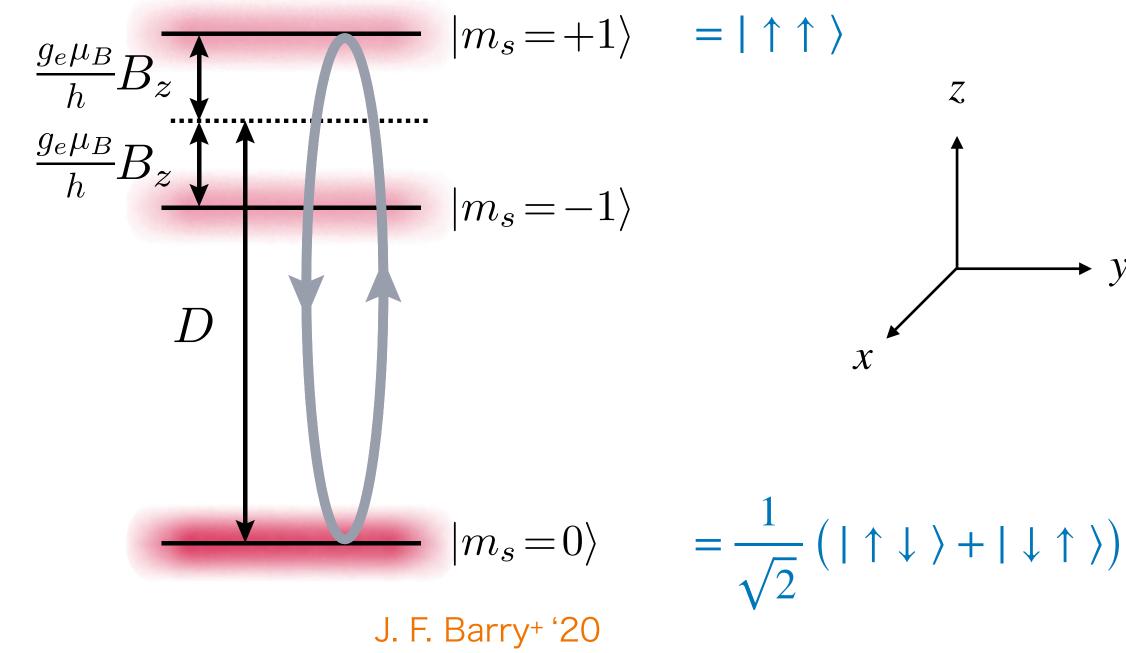


Rabi cycle

- Energy gap $\Delta E \sim 2\pi \times 2.87 \,\text{GHz}$
- Inject oscillating driving field with frequency $f = D + \frac{1}{2\pi} \gamma_e B_z$
 - $|-\rangle$ is irrelevant
 - qubit system of $|0\rangle$ and $|+\rangle$
- Under the transverse magnetic field $\mathbf{B}_1 = B_{1v} \,\hat{\mathbf{y}} \sin(2\pi f t),$

Time evolution is described by the Rabi cycle

$$|\psi(t)\rangle = \cos\left(\frac{1}{\sqrt{2}}\gamma_e B_{1y}t\right)|0\rangle + \sin\left(\frac{1}{\sqrt{2}}\gamma_e B_{1y}t\right)|0\rangle$$



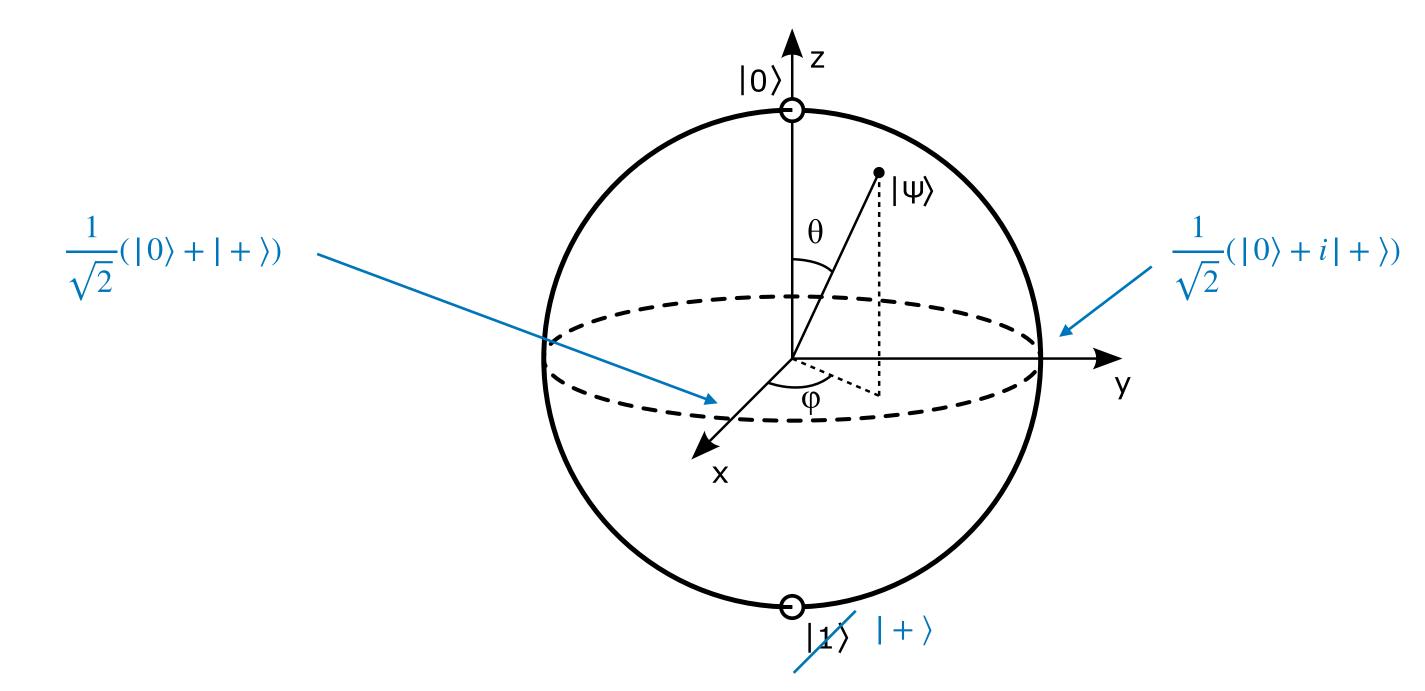
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Bloch sphere

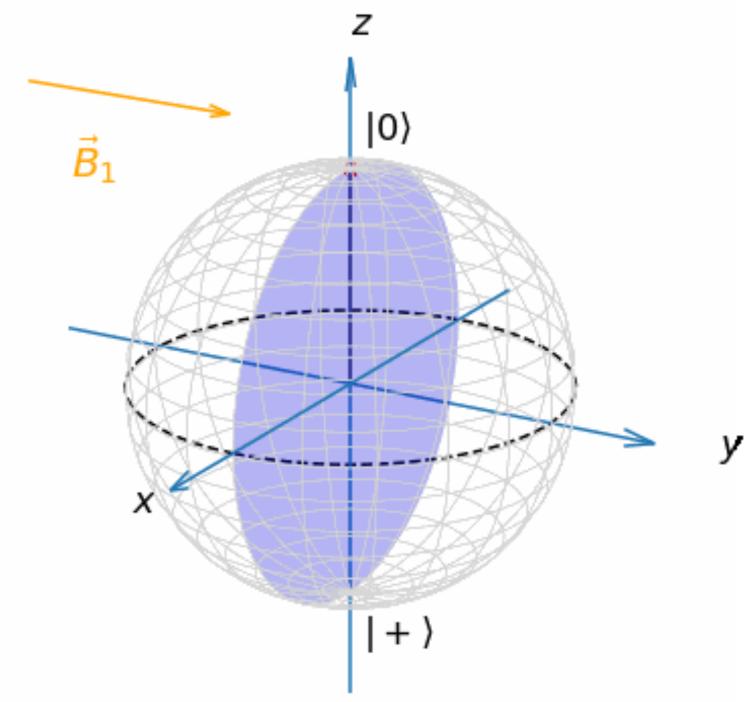


• Map from SU(2) group elements to the sphere S^2

$$|\psi\rangle = \cos\frac{\theta}{2}|0\rangle + \sin\frac{\theta}{2}e^{i\varphi}|+\rangle = (|0\rangle |+\rangle) \begin{pmatrix} \cos\frac{\theta}{2} & -\sin\frac{\theta}{2}e^{-i\varphi} \\ \sin\frac{\theta}{2}e^{i\varphi} & \cos\frac{\theta}{2} \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$



Rabi cycle on Bloch sphere



• Rotation around $\vec{B}_1 \propto \hat{\mathbf{y}}$

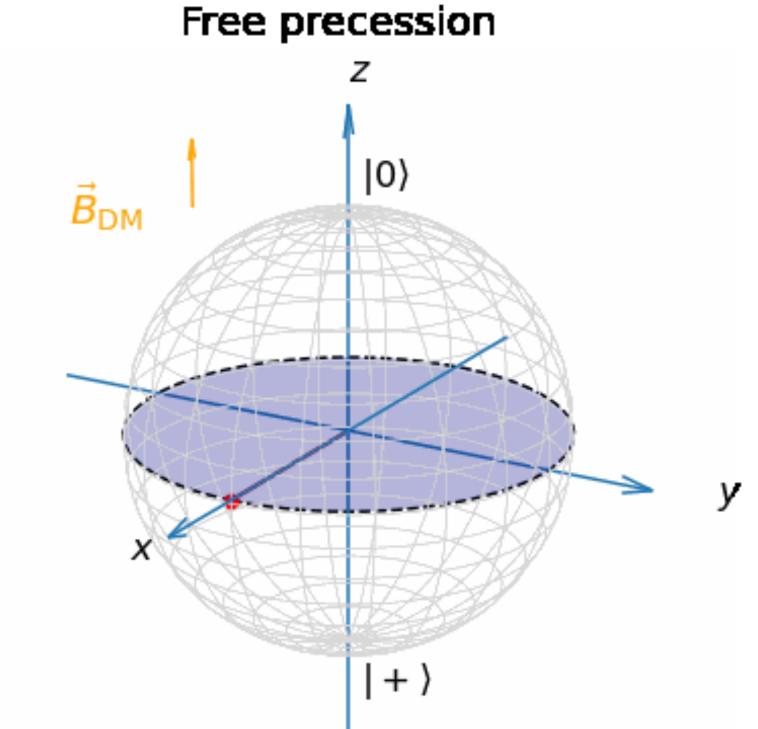
$$|\psi(t)\rangle = \cos\frac{\theta(t)}{2}|0\rangle + \sin\frac{\theta(t)}{2}|+\rangle$$
 with $\theta(t) = \sqrt{2\gamma_e}B_{1y}t$

Rabi cycle

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Free precession



• Magnetic field $\overrightarrow{B} \propto \hat{z}$ causes free precession = rotation around \hat{z} $|\psi(\tau)\rangle = \frac{1}{\sqrt{2}} \left(|0\rangle + e^{i\varphi(\tau)}|+\rangle\right)$ with $\varphi(\tau) = \gamma_e \int_0^{\tau} dt B_{\rm DM}^z(t) \simeq \gamma_e B_{\rm DM}^z \tau$ (for DC-like signal)

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Ramsey sequence **Ramsey sequence for DC magnetometry** 1. $(\pi/2)_v$ pulse

• Rabi cycle with $\theta = \sqrt{2}\gamma_e B_{1v}t = \pi/2$

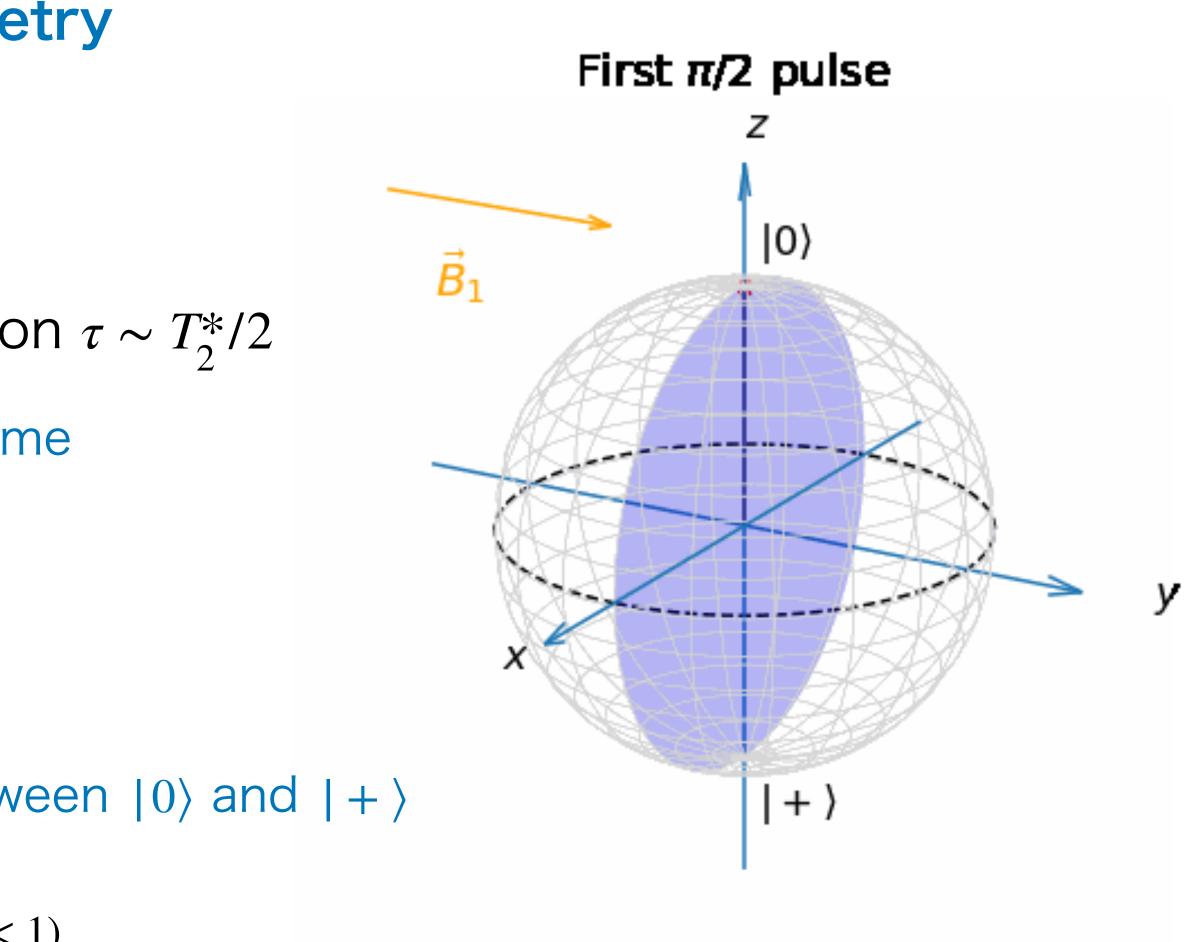
2. Free precession under **B**_{DM} for duration $\tau \sim T_2^*/2$

• $T_2^* \sim 1 \,\mu s$: spin relaxation (dephasing) time

- 3. $(\pi/2)_r$ pulse
- 4. Fluorescence measurement
 - DM signal is population difference between $|0\rangle$ and $|+\rangle$

 $S \equiv \frac{1}{2} \langle \psi_{\text{fin.}} | \sigma_z | \psi_{\text{fin.}} \rangle \propto \varphi(\tau) = \gamma_e B_{\text{DM}}^z \tau \quad (\varphi(\tau) \ll 1)$





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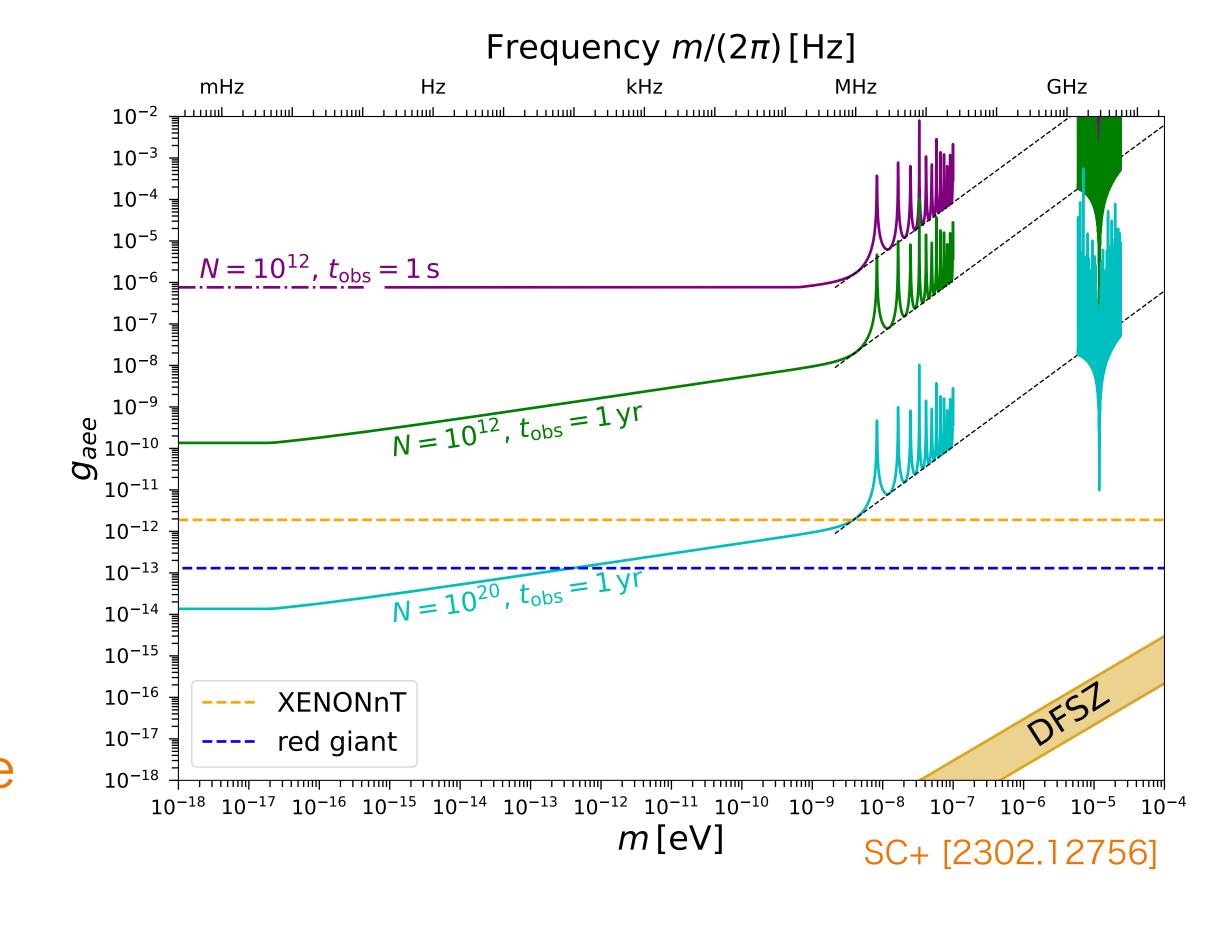


Sensitivity on axion DM Assume spin-projection noise limits the sensitivity

$$|x\rangle \equiv \frac{1}{\sqrt{2}} (|0\rangle + |+\rangle)$$

$$\Delta S \equiv \frac{1}{2} \left[\langle x | \sigma_z^2 | x \rangle - (\langle x | \sigma_z | x \rangle)^2 \right]^{1/2} = \frac{1}{2}$$

- Large statistics reduce noise
 - *N* : # of NV centers
 - t_{obs} : total observation time
- (Roughly) universally sensitive to dc-like signal with $m \leq 2\pi/\tau \sim 10^{-8} \,\mathrm{eV}$



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Effects of DM coherence time

- $B_{\rm DM}^z$ and δ change randomly with $\tau_{\rm DM} \sim 2\pi/m_{\rm DM} v_{\rm DM}^2$
- For $t_{\rm obs} \ll \tau_{\rm DM}$
 - Fixed $B_{\rm DM}^z$ and δ
 - (# of observations) $\simeq N(t_{\rm obs}/\tau)$
 - (Sensitivity) $\propto N^{1/2} (t_{\rm obs}/\tau)^{1/2}$
- For $t_{\rm obs} \gg \tau_{\rm DM}$
 - We measure the variance of $S_{\rm obs}$
 - Comparison of $\Delta S_{\rm DM}$ and $\Delta S N^{-1/2} (\tau_{\rm DM}/\tau)^{-1/2}$
 - (Sensitivity) $\propto N^{1/2} (\tau_{\rm DM}/\tau)^{1/2} (t_{\rm obs}/\tau_{\rm DM})^{1/4}$ Consistent with Dror+ [2210.06481] in the context of CASPEr

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10⁻³ 10^{-4} 10^{-5} $N = 10^{12}$, $t_{obs} = 1$ s 10^{-7} 10^{-8} $N = 10^{12}$, $t_{obs} = 1$ yr 10^{-5} 10-11 10^{-12} $N = 10^{20}, t_{obs} = 1y$ 10^{-13} 10^{-14} 10^{-15} 10^{-16} **XENONnT** 10^{-17} red giant 10^{-16} 10^{-15} 10^{-14} 10^{-13} 10^{-12} 10^{-11} 10^{-10} 10^{-9} 10^{-8} 10^{-7} 10^{-6} 10^{-5} 10^{-4} m[eV] SC+ [2302.12756]

Ηz

mHz

Frequency $m/(2\pi)$ [Hz]

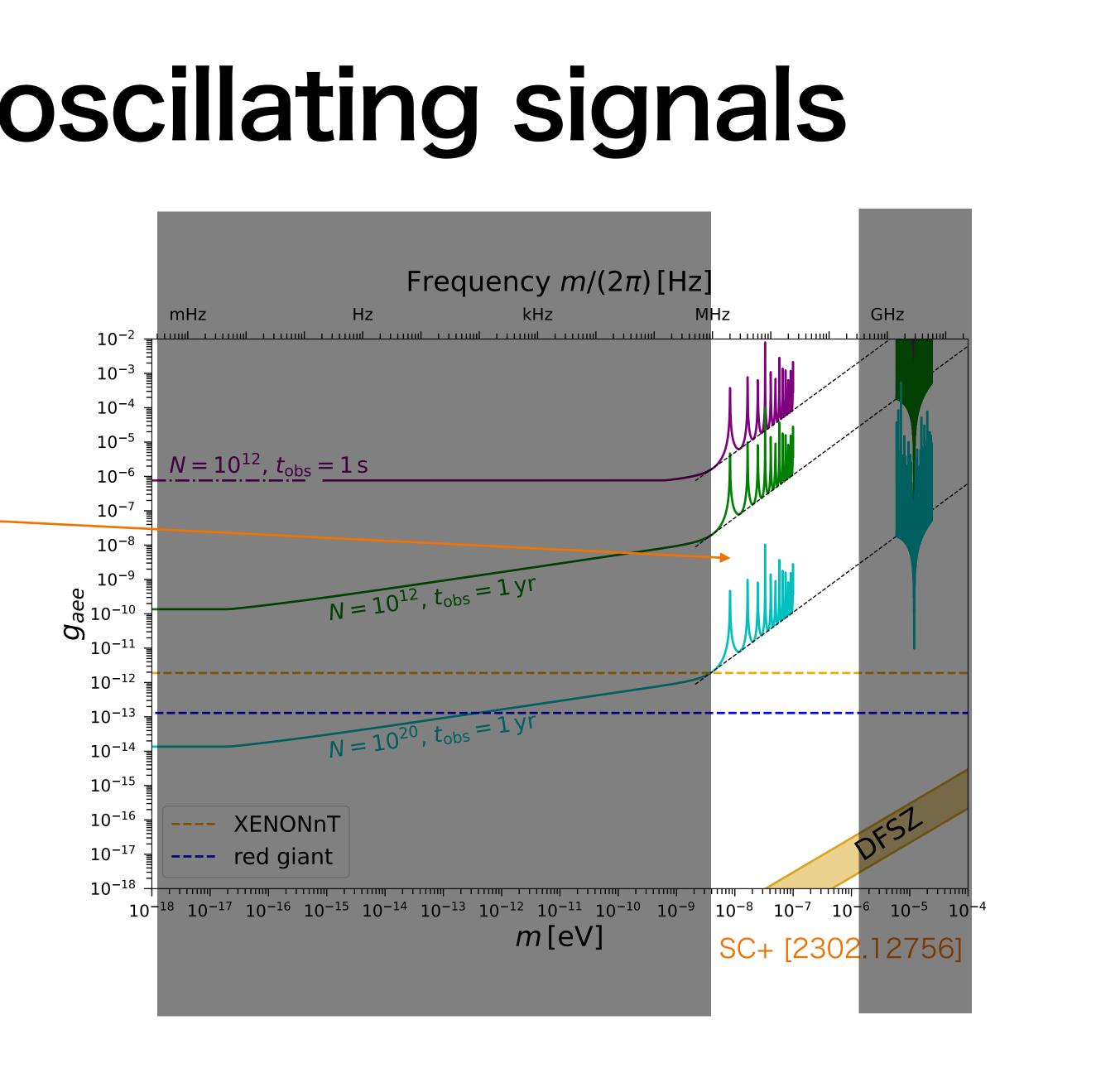


GHz

Insensitive to fast-oscillating signals

Fast oscillation leads to cancellation

$$S \sim \int_0^{\tau} dt B_{\rm DM}^z \sin(mt) \propto \frac{1 - \cos(m\tau)}{m\tau}$$



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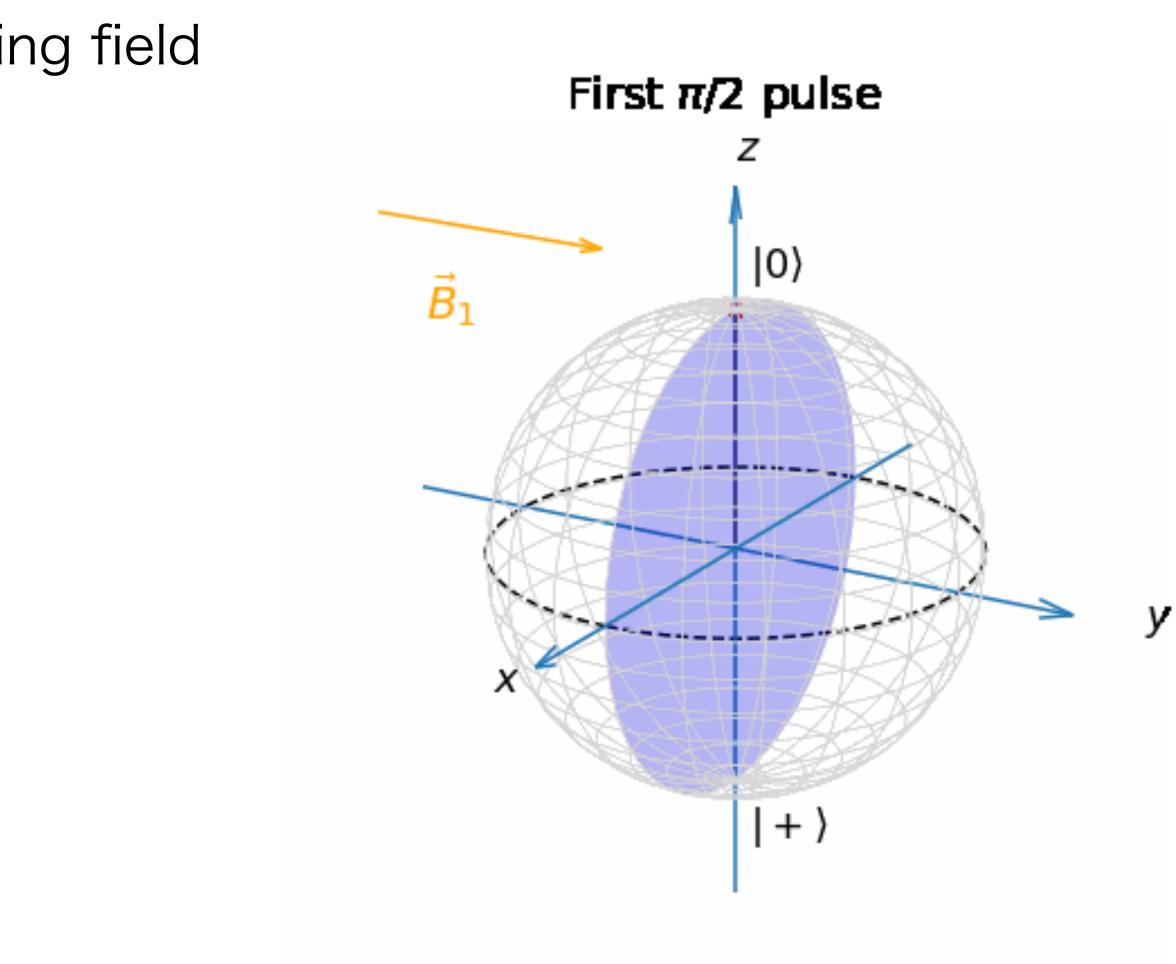


DM on resonance

If $m/2\pi \simeq f$, DM field itself works as a driving field

"Resonance" sequence for $m/2\pi \simeq f$

- 1. $(\pi/2)_v$ pulse
- 2. Free precession for duration $\tau \sim T_2^*/2$
- 3. Fluorescence measurement $S \propto B_{\rm DM}^y \tau$

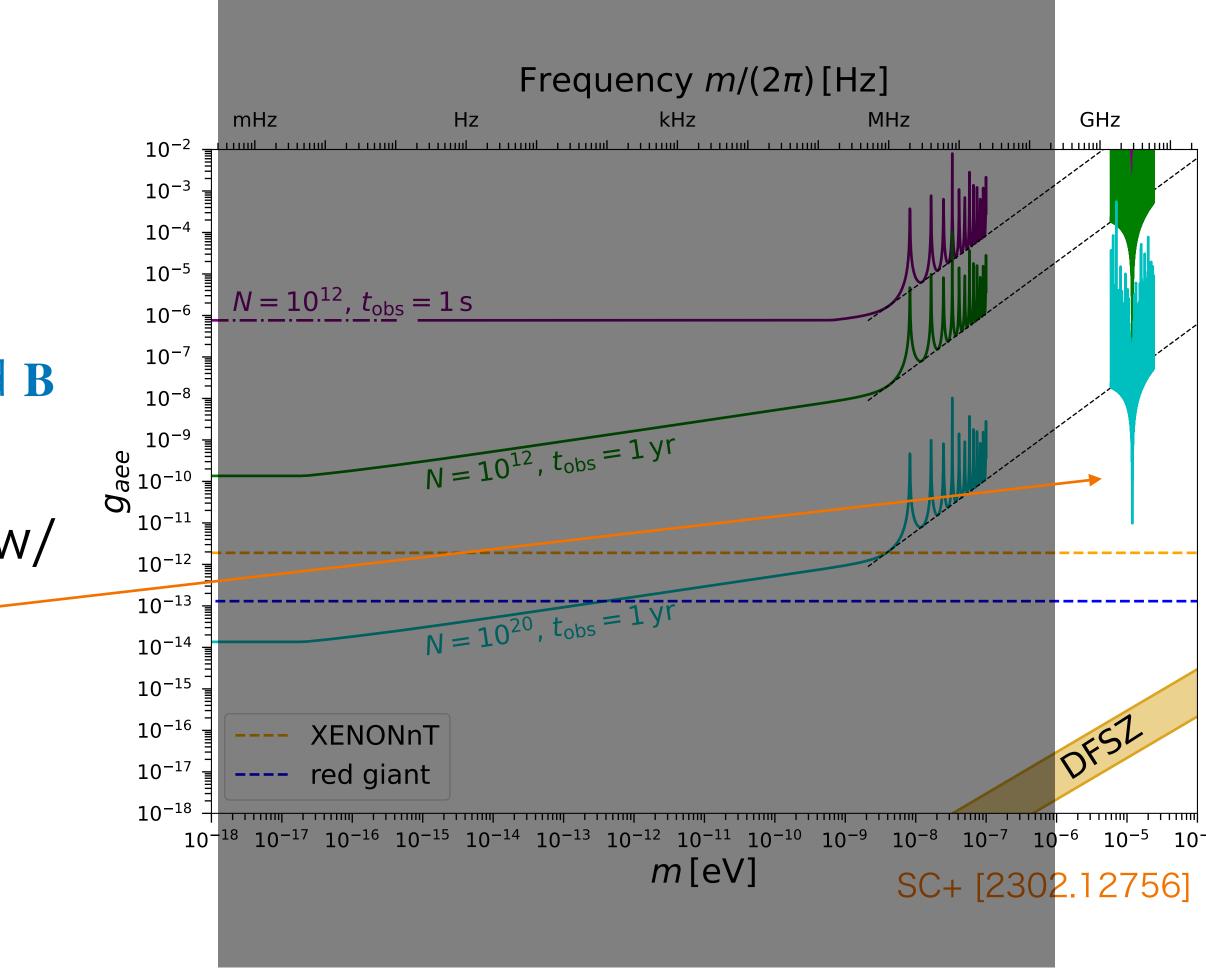




On resonance sensitivity

- Resonance position $\frac{m}{2\pi} \simeq 2.87 \,\text{GHz} \Leftrightarrow m \simeq 11.9 \,\mu\text{eV}$
 - Tunable with e.g., external magnetic field **B**
- Resonant enhancement of sensitivity w/

$$m\tau \sim 2 \times 10^4 \left(\frac{\tau}{1\,\mu s}\right)$$



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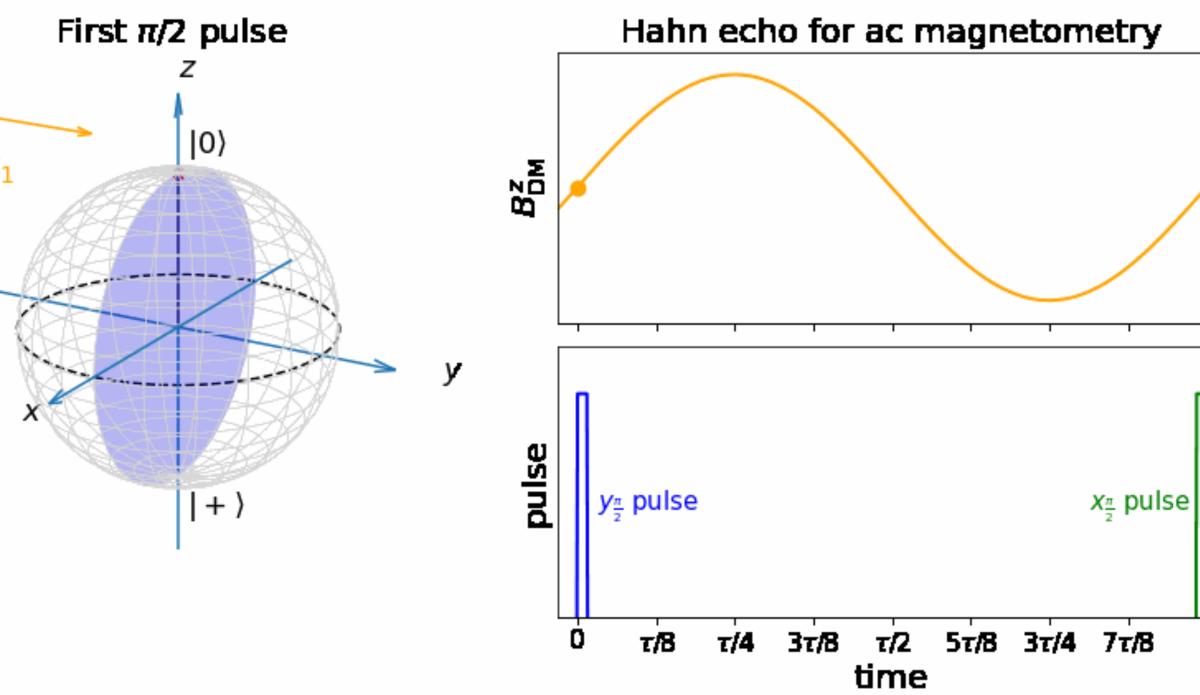
AC magnetometry



Insensitive to fast-oscillating signals

 Fast oscillation leads to cancellation when $m \leq 2\pi/\tau$

 \vec{B}_1





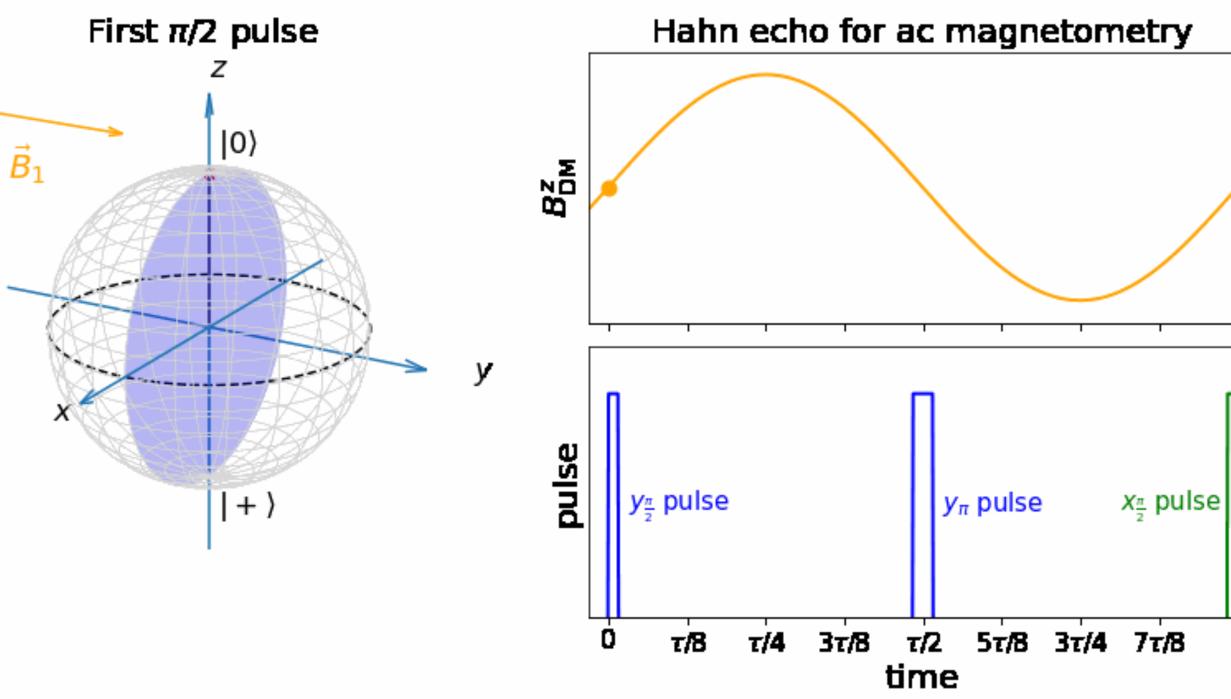


Hahn echo

Hahn echo for ac magnetometry

- 1. $(\pi/2)_{y}$ pulse
- 2. Free precession for $\tau/2$
- 3. π_v pulse
- 4. Free precession for $\tau/2$
- 5. $(\pi/2)_x$ pulse
- 6. Fluorescence measurement

$$\varphi(\tau) = \gamma_e \left(\int_0^{\tau/2} dt \, B_{\rm DM}^z(t) \, - \, \int_{\tau/2}^{\tau} dt \, B_{\rm DM}^z(t) \right) \Longrightarrow$$

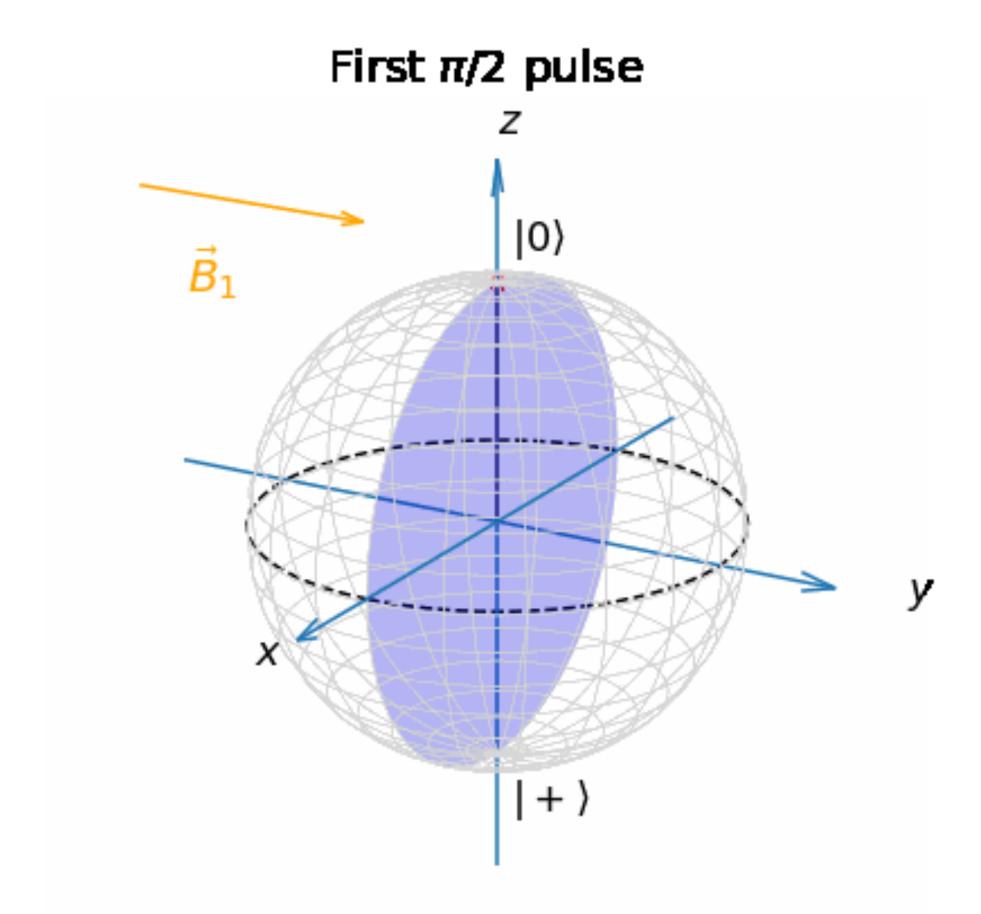


Targeted at the frequency $\sim 1/\tau$

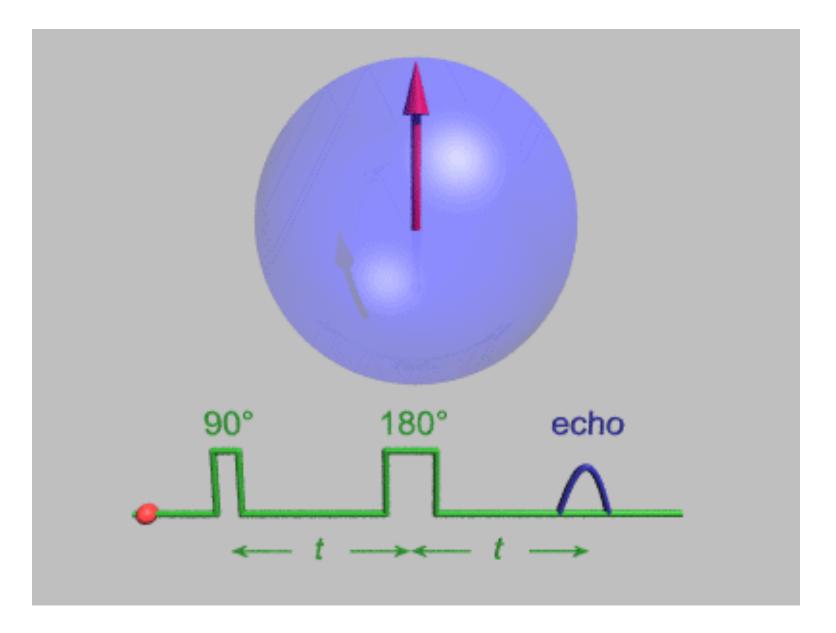




Prolonged relaxation time



• Any DC effect cancels out from $\varphi(t)$



- No dephasing from inhomogeneous DC fields
- Relaxation time $T_2 \sim 50 \,\mu s \gg T_2^* \sim 1 \,\mu s$
- Optimized choice $\tau \sim T_2/2$

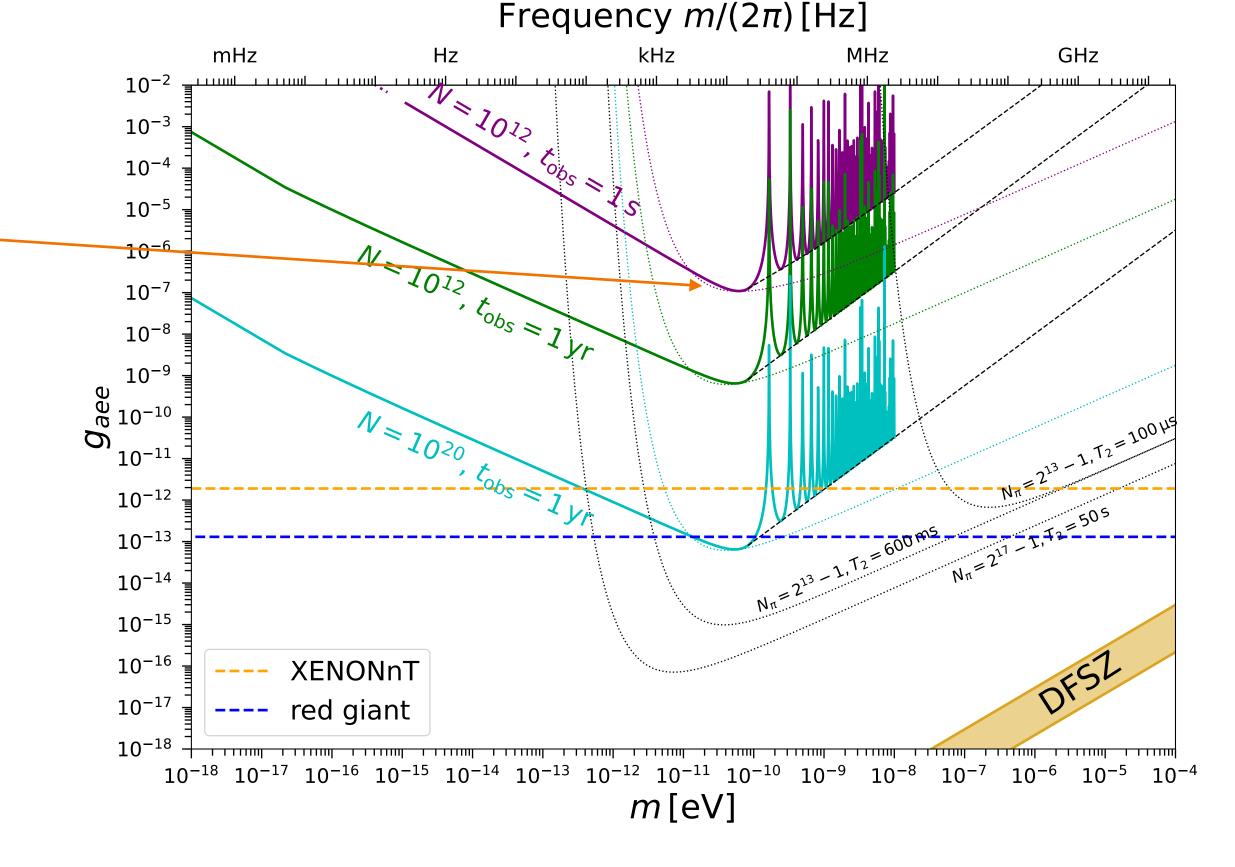


Sensitivity on axion DM

- Peak position $\frac{m}{2\pi} \sim \frac{1}{\tau} \sim 20 \,\mathrm{kHz}$
 - Better sensitivity around the peak than DC thanks to $T_2 \gg T_2^*$

• Tunable peak position with shorter τ

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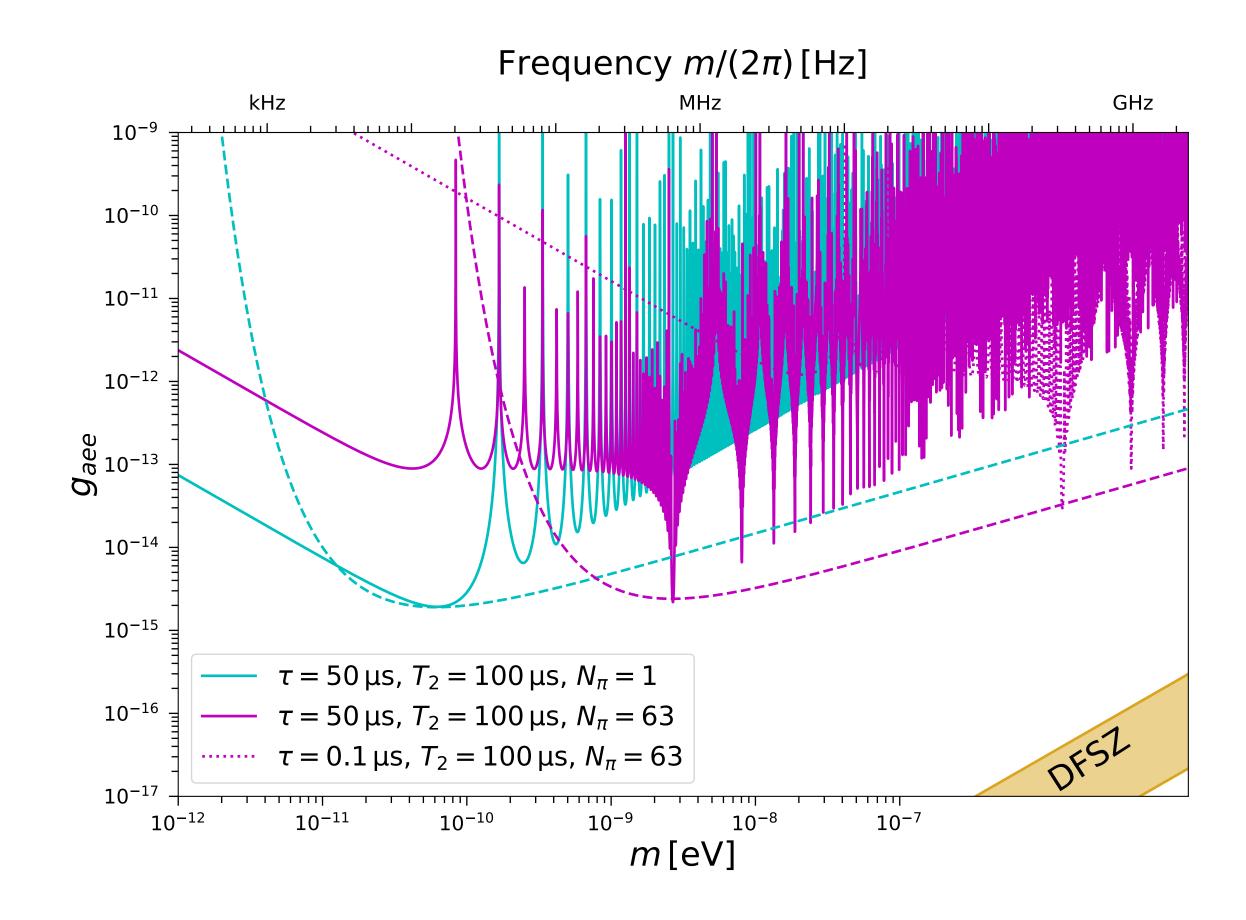




Towards sensitivity improvement

- Using More π_v pulses prolongs T_2
 - . Upper limit on $T_2 < T_1$
 - . target frequency $\times N_{\pi}$
- Lower temperature prolongs T_2, T_1 (With $N_{\pi} = 1023$)
 - 300 K : $T_2 = 100 \,\mu\text{s}, T_1 \sim 1 \,\text{ms}$
 - 77 K : $T_2 = 1 \text{ ms}, T_1 \sim 1 \text{ s}$
 - 4 K : $T_2 = 10 \text{ ms}, T_1 \gg 1 \text{ s}$
 - $0.1 \text{ K} : T_2 = 0.1 \text{ s}, T_1 \gg 1 \text{ s}$

D. Herbschleb, private communication

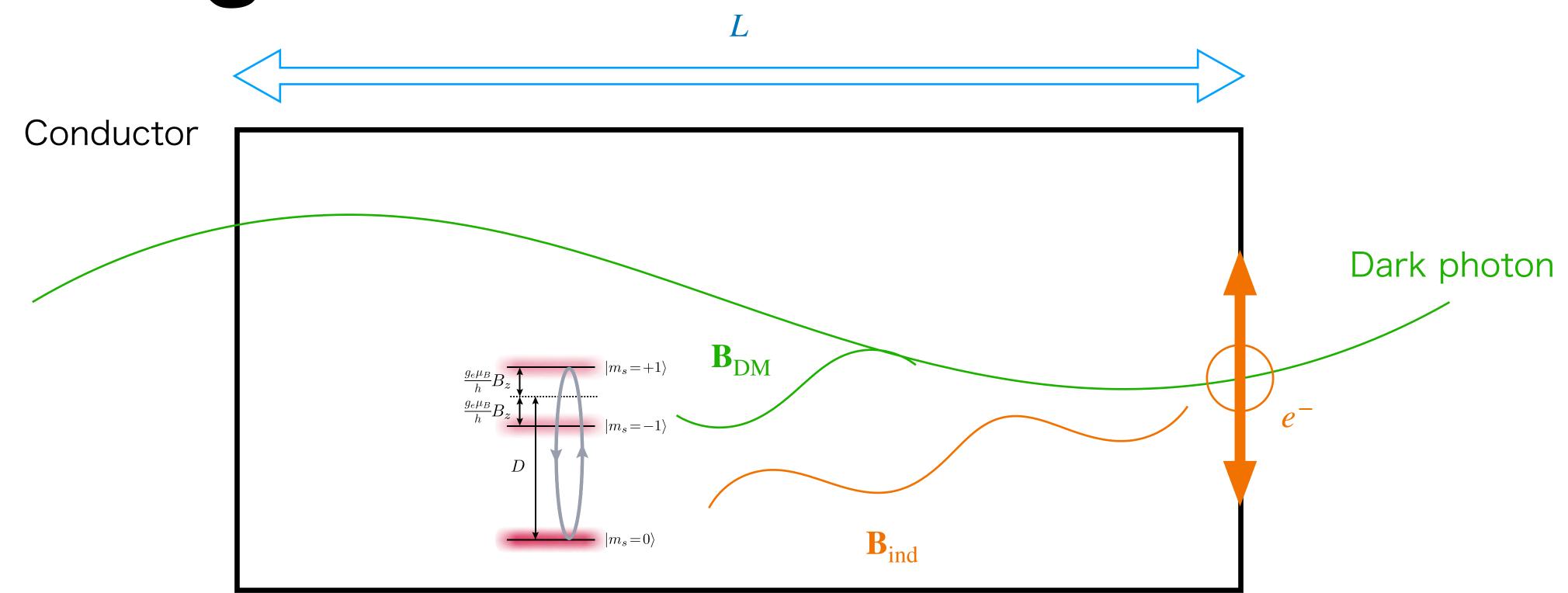


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Dark photons

Shielding effect



- Electric interaction of the dark photon creates current in the conductor and induces a magnetic field **B**_{ind}
- The effective magnetic field may be canceled and "shielded" if $\lambda_{DM} > L$

S. Chaudhuri+ [1411.7382] "DM Radio" paper 11/13/2023 So Chigusa @ UC Davis



NV center works without shielding

RO signal (arb. units) 🖲

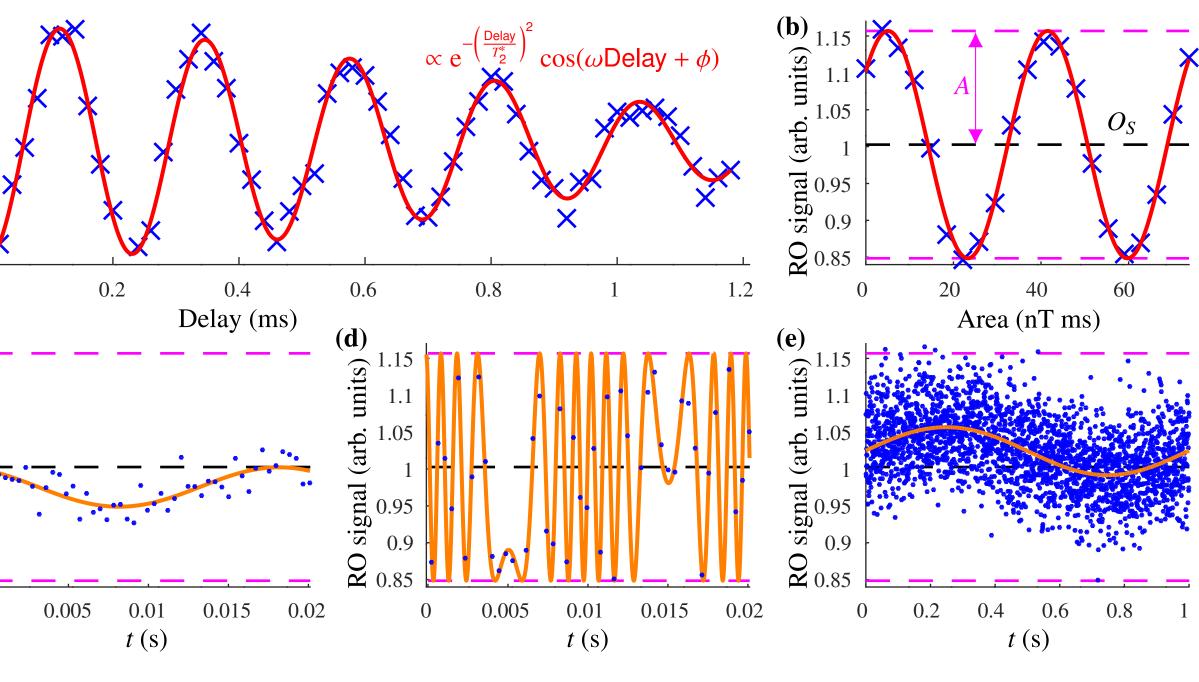
AC magnetometry is insensitive to DC(-like) noises

• Applicable for $\frac{m}{2\pi} \gtrsim 1 \, \text{kHz}$

"Low-frequency quantum sensing" signal (arb. 1 0.92 0.9 possible with Fourier analysis

• Applicable for $\frac{m}{2\pi} \gtrsim \frac{1}{t_{\rm the}}$

Must be careful with AC magnetic noises with the target frequency



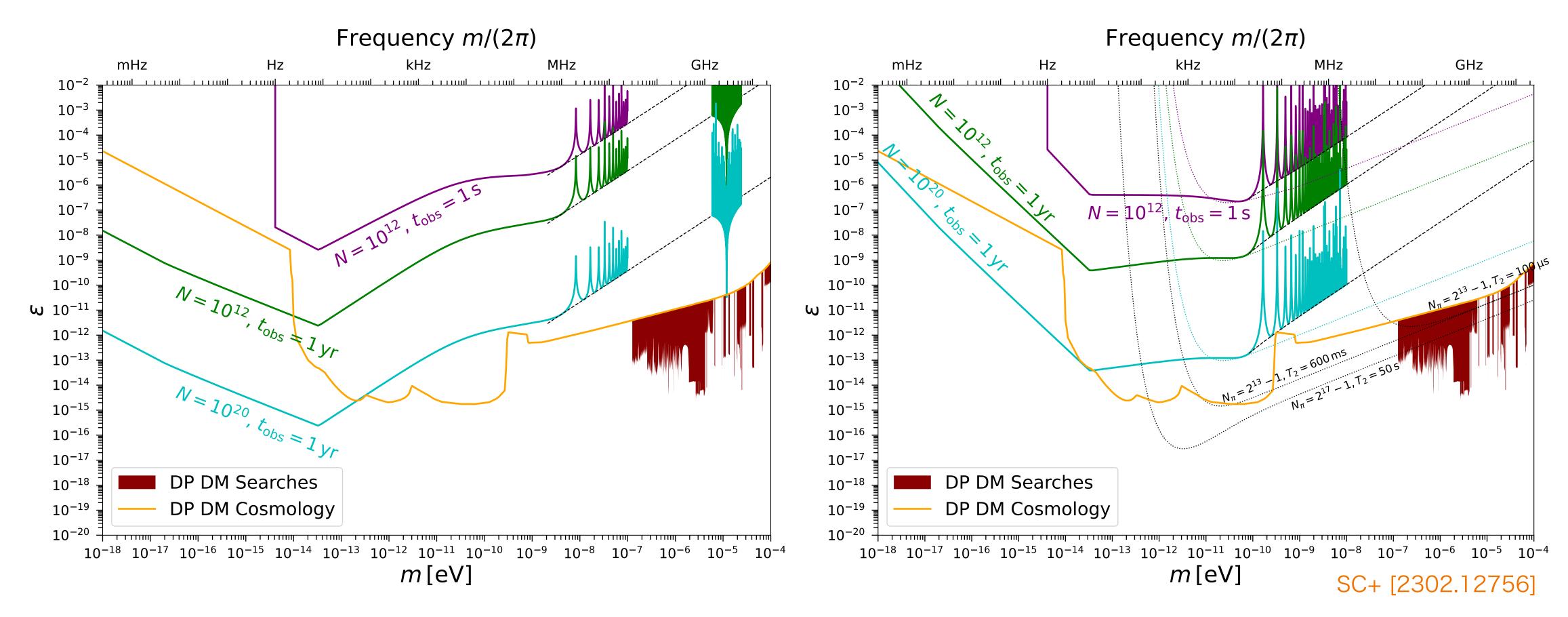
D. Herbschleb+ [2209.13870]

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Sensitivities on dark photon DM

DC magnetometry

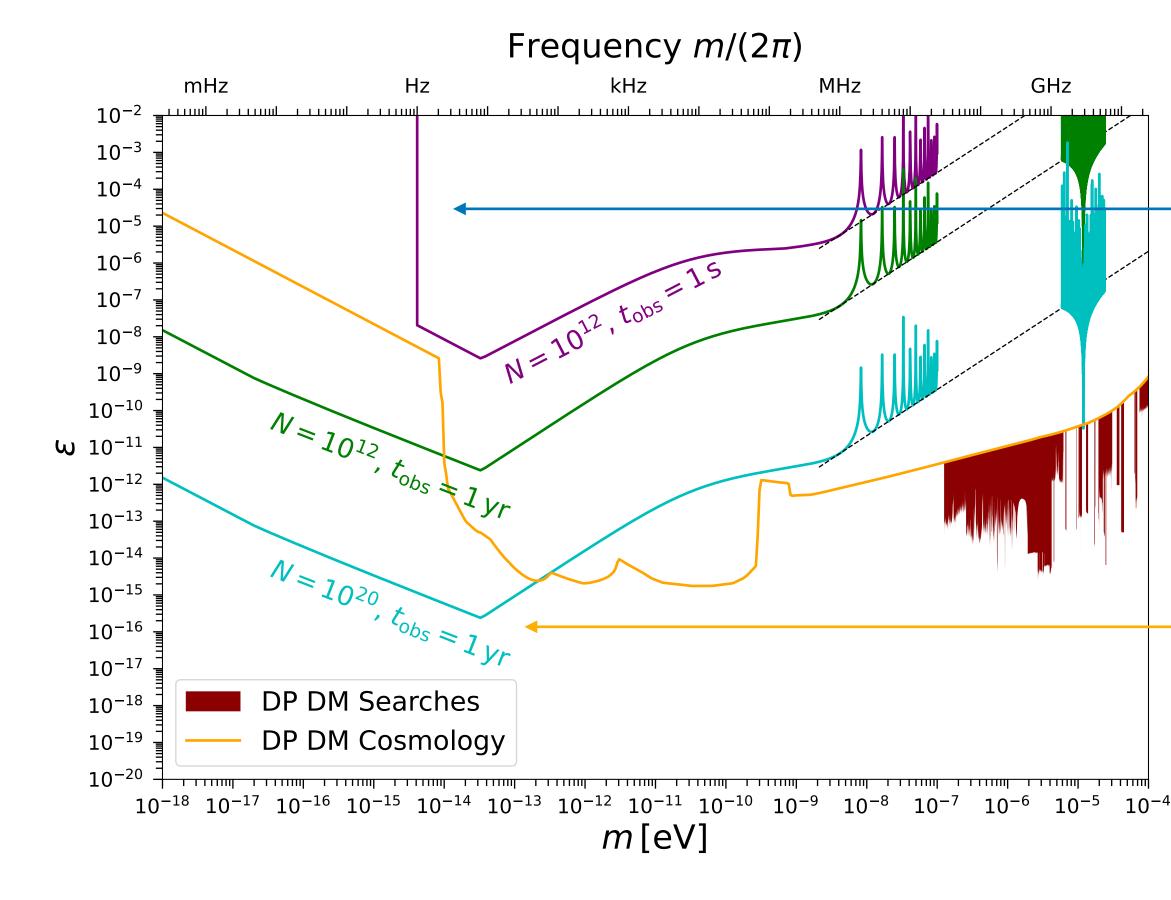


AC magnetometry

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Assumptions on magnetic shielding



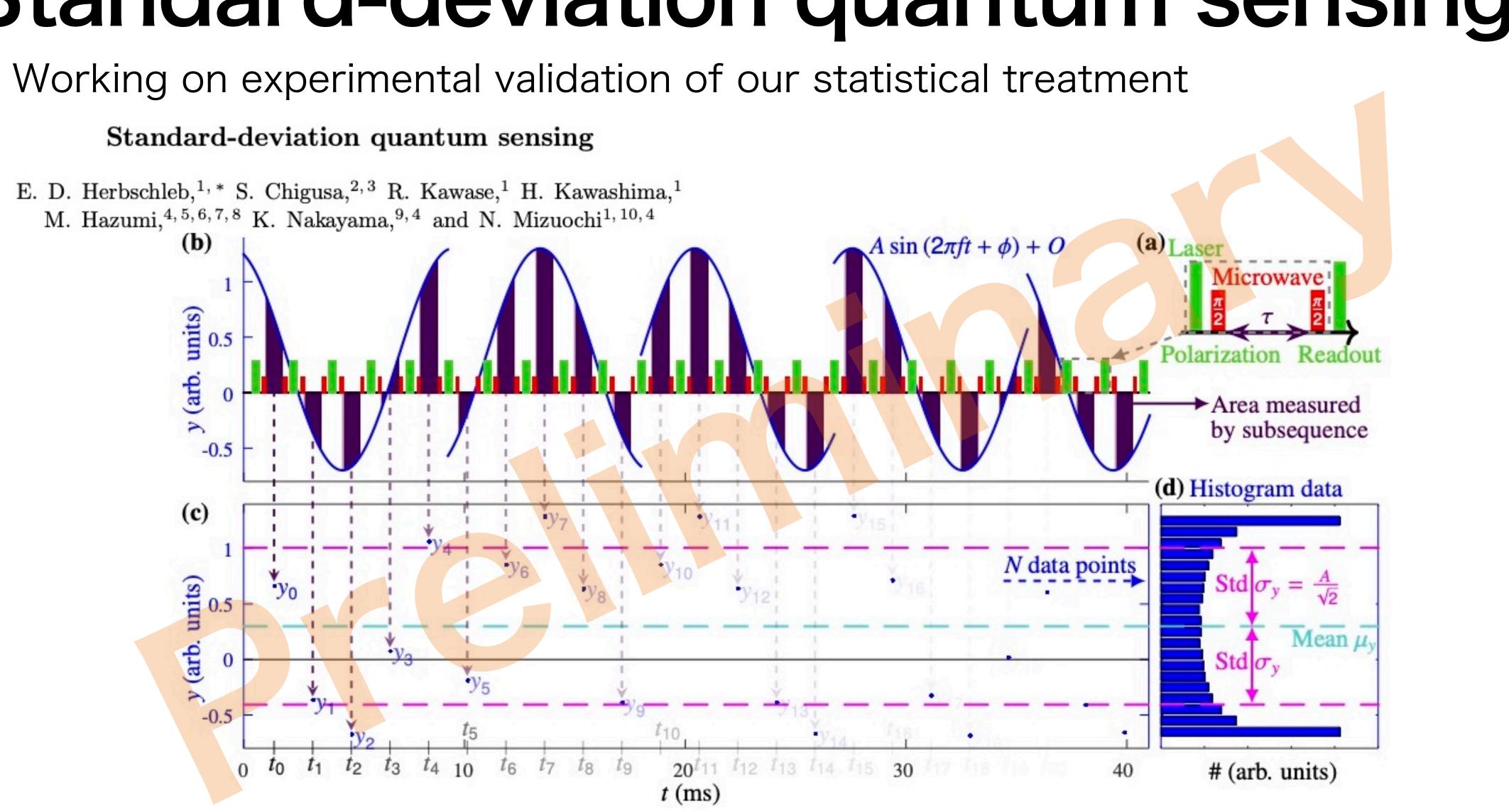
- We put shielding for $\frac{m}{2\pi} \lesssim \frac{1}{t_{obs}}$
 - Sensitivity is significantly suppressed
- Even without magnetic shielding, the inner core of the Earth/ionosphere are conducting and shielding fields
 - M. A. Fedderke+ [2106.00022]
 - Suppression factor $\propto mR_{\text{Earth}}$



Experimental status

Standard-deviation quantum sensing

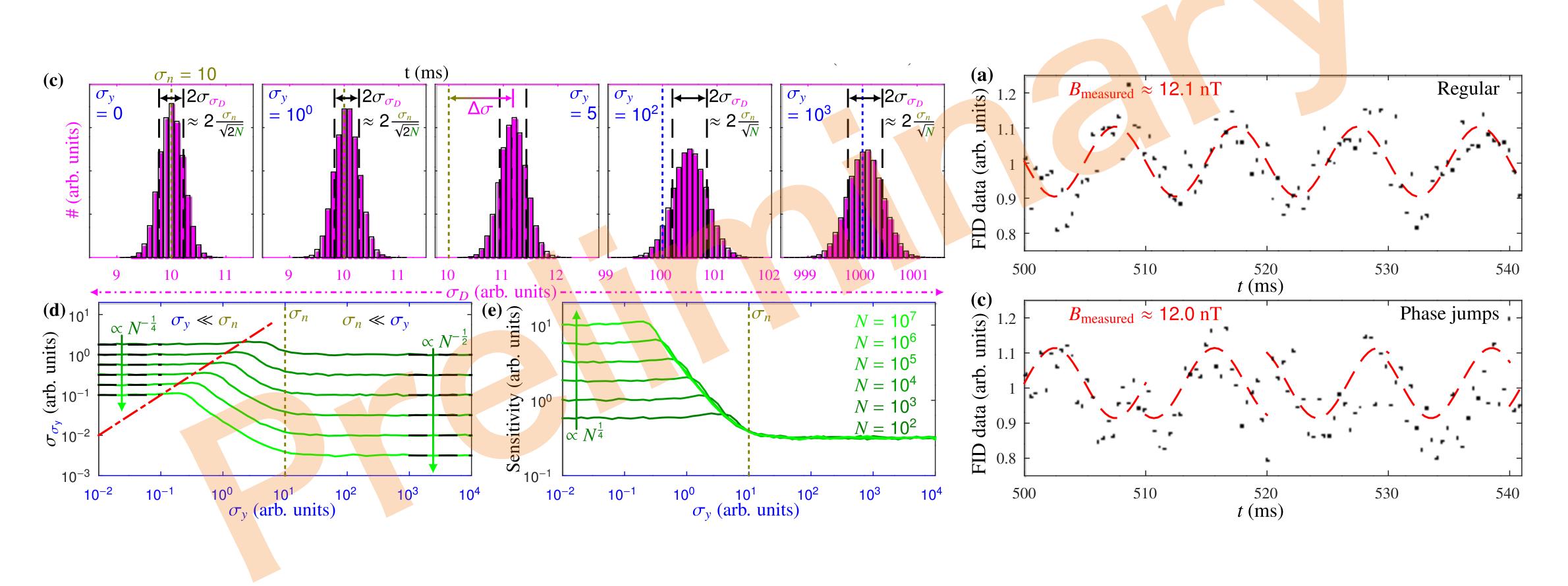
E. D. Herbschleb,^{1,*} S. Chigusa,^{2,3} R. Kawase,¹ H. Kawashima,¹ M. Hazumi,^{4, 5, 6, 7, 8} K. Nakayama,^{9, 4} and N. Mizuochi^{1, 10, 4}





Standard-deviation quantum sensing

- Obtained expected dependence on # of data points N
- Can estimate signal amplitude and frequency



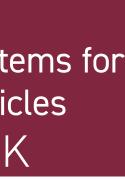
Discussions and conclusions

- We explored the potential of NV center magnetometry for DM search
- Benefits of this approach include:
 - Wide dynamic range = broad DM mass range is searched for
 - Not always need magnetic shielding
- - e.g.) Use of entanglement C. L. Degan+ "Quantum sensing" $|\psi\rangle = \bigotimes_c \frac{1}{\sqrt{2}} (|0\rangle_c + e^{i\varphi}|1\rangle_c) \rightarrow |\psi\rangle = \frac{1}{\sqrt{2}} (|000...\rangle + e^{iN\varphi}|111...\rangle)$
- Now setting up an experimental environment at QUP with NV + cryogenic

Some applications of advanced quantum sensing techniques can be considered



International Center for Quantum-field Measurement Systems for Studies of the Universe and Particles WPI research center at KEK





Backup slides

Sensitivity estimation The outcome of the spin-projection noise

$$|x\rangle \equiv \frac{1}{\sqrt{2}} (|0\rangle + |+\rangle)$$

$$\Delta S \equiv \frac{1}{2} \left[\langle x | \sigma_z^2 | x \rangle - (\langle x | \sigma_z | x \rangle)^2 \right]^{1/2} = \frac{1}{2}$$

Noise contribution is $\Delta S_{\rm sp} \sim \begin{cases} \frac{1}{2} \frac{1}{\sqrt{N(t_{\rm obs}/\tau)}} \\ \frac{1}{2} \frac{1}{\sqrt{N(\tau_a/\tau)}} \frac{1}{(t_{\rm obs}/\tau_a)} \end{cases}$

• Sensitivity curve is (SNR) $\equiv \frac{S}{\Delta S_{sp}} = 1$

$$(t_{\rm obs} < \tau_a)$$

$$\frac{1}{\left(t_{\rm obs}/\tau_a\right)^{1/4}} \quad (t_{\rm obs} > \tau_a)$$

So Chigusa @ PASCOS 2023 (6/27)



Sensitivity estimation

• The axion-induced effective magnetic field has an unknown velocity \mathbf{v}_{DM} and phase δ

$$\mathbf{B}_{\rm DM} \simeq \sqrt{2\rho_{\rm DM}} \frac{g_{aee}}{e} \mathbf{v}_{\rm DM} \sin(m_{\rm DM}t + \delta)$$

Random velocity v_{DM}

- The signal is proportional to $(v_{DM}^i)^2$ $(i = x, i)^2$ Random phase $\delta \in [0, 2\pi)$
- The signal is estimated as a function of
- compared with the noise

, y, z), which is averaged to
$$\sim \frac{1}{3}v_{\rm DM}^2$$

$$\delta : S(\delta) \propto \cos\left(\frac{m\tau}{2} + \delta\right)$$

• We obtain the average $\langle S \rangle_{\delta} = 0$ and the standard deviation $\sqrt{\langle S^2 \rangle} \neq 0$, which should be

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Technical noise mitigation

II. MAGNETOMETRY METHOD

In many high-sensitivity measurements, technical noise such as 1/f noise is mitigated by moving the sensing bandwidth away from dc via upmodulation. One method, common in NV-diamond magnetometry experiments, applies frequency [12,32,41,42] or phase modulation [19,43–45] to the MWs addressing a spin transition, which causes the magnetic-field information to be encoded in a band around the modulation frequency. Here we demonstrate a multiplexed [46–49] extension of this scheme, where information from multiple NV orientations is encoded in separate frequency bands and measured on a single optical detector. Lock-in demodulation and filtering then extracts the signal associated with each NV orientation, enabling concurrent measurement of all components of a dynamic magnetic field. J. M. Schloss+ '18

