

# Multiple/Degenerate 125 GeV Higgs Scenarios within the NMSSM Perspective

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# Outline

- 1 Preliminary Background
  - 125 GeV Higgs-like signal at the LHC and the Tevatron
  - NMSSM review
- 2 Single 125 GeV Higgs Scenarios
- 3 Degenerate 125 GeV Higgs Scenarios
  - $h_1 \sim 125 + h_2 \sim 125$  phenomenology
  - Diagnosing tools
- 4 Multiple 125 GeV Higgs Scenarios
  - $h_1 \sim 125 + h_2 \sim 136$  LHC–Tevatron scenario
  - $h_1 \sim 98 + h_2 \sim 125$  LEP–LHC scenario
  - Future test
- 5 Conclusion and Outlook

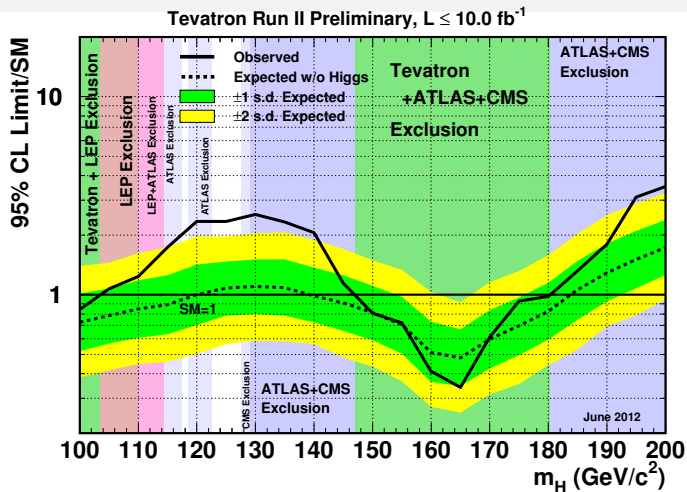
Higgs – God particle – has been attracting worldwide attention over the years.



A faded Chinese article published in the public daily newspaper in 1980s.

The title is "God, you **DO** really exist!"

## SM Higgs Search Overview (Prior to July 4th, 2012)



SM Higgs mass  $m_H$  excluded regions:

LEP:  $< 114 \text{ GeV}$  (1989-2000).

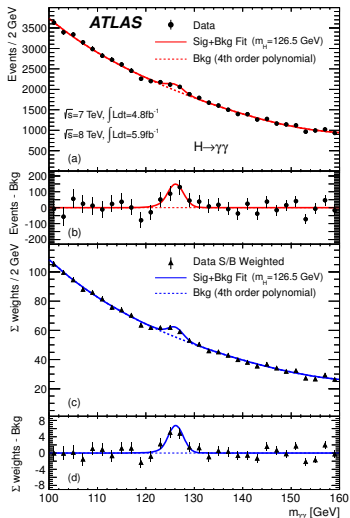
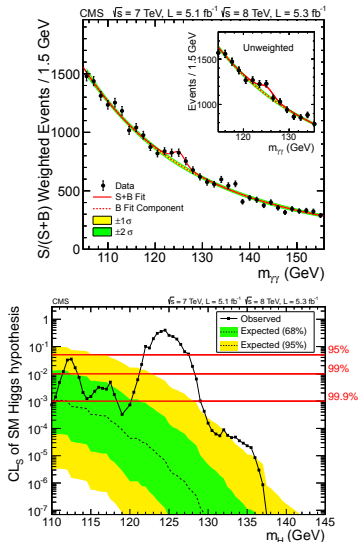
Tevatron:  $100 \text{ GeV} - 106 \text{ GeV}$  and  $147 \text{ GeV} - 179 \text{ GeV}$ .

CMS:  $127.5 \text{ GeV} - 600 \text{ GeV}$ .

ATLAS:  $110 \text{ GeV} - 117.5 \text{ GeV}$ ,  $118.5 \text{ GeV} - 122.5 \text{ GeV}$  and  $129 \text{ GeV} - 539 \text{ GeV}$ .

## 125 GeV Higgs-like signal at the LHC

After over **thirty** years of waiting, this summer, CMS and ATLAS both saw a “new boson” decaying into two photons, with a mass at **around 126 GeV**:



# A HISTORIC moment in science.

It is a privilege to witness the Higgs discovery, on July 4th, 2012.

## 国际新闻

INTERNATIONAL NEWS

新快报

A29

# 天哪！这真是“上帝粒子”吗？

欧洲核子中心激动宣布可能发现希格斯-玻色子：“我们对宇宙的理解，将要改变！”

新华社日内瓦4日电，欧洲核子中心(CERN)的科学家在日内瓦的欧洲核子研究中心，宣布他们可能发现了“上帝粒子”——一种被称为希格斯玻色子的基本粒子。这一发现被认为是物理学史上最重要的突破之一，因为它解释了为什么其他基本粒子具有质量。CERN的科学家们表示，他们通过大型强子对撞机(LHC)的实验，在2012年7月4日首次观测到了这种粒子的信号。这一发现是科学家们几十年来努力的结果，也是人类对自然界最深层奥秘的一次重大探索。希格斯玻色子的发现，将帮助我们更好地理解宇宙的基本构成和运行规律。



### 物理学家的发现

#### 物理学家们的发现值得期待

这一发现被认为是物理学史上最重要的突破之一，因为它解释了为什么其他基本粒子具有质量。CERN的科学家们表示，他们通过大型强子对撞机(LHC)的实验，在2012年7月4日首次观测到了这种粒子的信号。这一发现是科学家们几十年来努力的结果，也是人类对自然界最深层奥秘的一次重大探索。希格斯玻色子的发现，将帮助我们更好地理解宇宙的基本构成和运行规律。

### 物理学家的发现

#### 83岁希格斯：未想过有生之年能见到

希格斯玻色子的发现，是科学家们几十年来努力的结果。希格斯本人表示，他从未想过有生之年能见到这一发现。这一发现不仅验证了理论物理学家的预言，也为人类对宇宙的理解打开了新的窗口。



#### 还需时间确认 意大利理论物理学家

尽管初步数据令人鼓舞，但科学家们仍需更多的数据来确认这一发现。意大利理论物理学家表示，他们将继续进行实验，以进一步验证希格斯玻色子的存在。这一过程可能需要数月甚至数年的时间。

### "Higgs" Announcement webcast link for later tonight.

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64 评论 分享 保存 hide report

### 网友围观：好像假牛 但是不懂

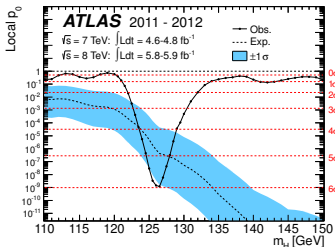
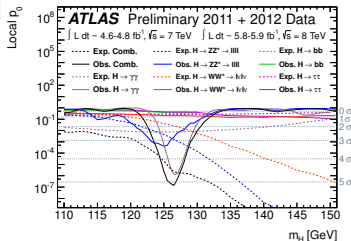
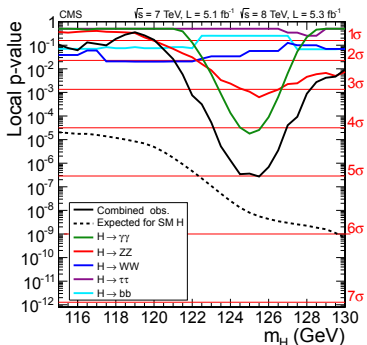
网友们对希格斯玻色子的发现表示好奇，但对其原理并不了解。有网友调侃说：“好像假牛，但是不懂”。这反映了普通大众对前沿科学知识的渴望和求知欲。

### “这该死的粒子”！

一些网友对希格斯玻色子的发现表示不满，称其为“该死的粒子”。他们认为这一发现过于复杂，难以理解。然而，科学家们表示，这一发现是人类智慧的结晶，是人类对自然规律的探索。

## 125 GeV Higgs-like signal at the LHC

Local p-values was updated on July 31st, 2012 for both collaborations.



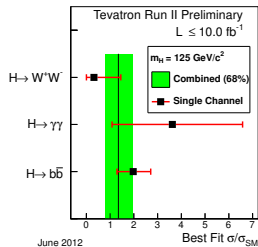
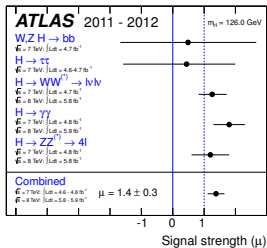
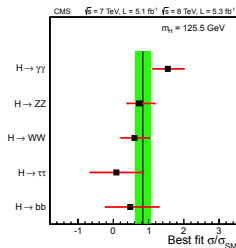
CMS and ATLAS provide an essentially  $5\sigma$  and  $6\sigma$  signal, respectively, for a Higgs-like resonance with mass of order 123–128 GeV.

With the new data, "Seeing is believing" !

# 125 GeV Higgs-like signal

Question: whether or not it *is* the SM Higgs? how to distinguish?

- reduced coupling
- signal strength as defined  $R_Y^h(X) \equiv \frac{\sigma(Y \rightarrow h_i) \text{BR}(h_i \rightarrow X)}{\sigma(Y \rightarrow h_{\text{SM}}) \text{BR}(h_{\text{SM}} \rightarrow X)}$



	gg fusion			VBF		VH
$m_H = 125$	$\gamma\gamma$	$ZZ^* \rightarrow 4\ell$	$WW^* \rightarrow 2\ell 2\nu$	$\gamma\gamma$	$\tau^+\tau^-$	$b\bar{b}$
ATLAS	$1.8 \pm 0.5$	$1.2 \pm 0.6$	$1.26 \pm 0.45$	$2.7 \pm 1.5$	$0.44^{+1.55}_{-2}$	$0.5 \pm 2$
CMS	$1.55^{+0.47}_{-0.45}$	$0.73^{+0.47}_{-0.36}$	$0.6^{+0.45}_{-0.4}$	$2.6 \pm 1.3$	$0.09^{+0.78}_{-0.76}$	$0.5^{+0.8}_{-0.7}$
Combined	$1.66 \pm 0.36$	$1.02 \pm 0.38$	x	$> 1$	very small	x
	high resolution			poor resolution		

Tevatron: the evidence for the Higgs boson is based principally on the  $W + H$  with  $H \rightarrow b\bar{b}$  decay mode, the observed enhancements relative to the SM rate by a factor of  $1.97^{+0.74}_{-0.69}$ .



# NMSSM=MSSM+Singlet Brief Review

I will focus on the discussion within the NMSSM perspective in which both a Higgs mass of order 125 GeV and significant  $\gamma\gamma$  mode enhancements are easily obtained.

$$W_{NMSSM} = W_{MSSM} + \cancel{H_u H_d} + \lambda S H_u H_d + \frac{\kappa}{3} S^3$$

with the additional  $m_S^2 |S|^2$ ,  $\lambda A_\lambda H_u H_d S$  and  $\frac{1}{3} \kappa A_\kappa S^3$  terms in the soft SUSY-breaking Lagrangian. The NMSSM is very attractive:

- 1 it solves the  $\mu$  problem of the MSSM:  $\mu_{\text{eff}} = \lambda \langle S \rangle \rightarrow M_{\text{SUSY}}$  ✓.
- 2 The three CP-even Higgs fields,  $H_u$ ,  $H_d$  and  $S$  mix and yield

### NMSSM Higgs Sector

- 3 CP-even neutral scalars:  $h_1, h_2, h_3$
- 2 CP-odd neutral pseudoscalar:  $a_1, a_2$
- 2 charged scalars:  $H^\pm$

- 3 The lightest CP-even Higgs mass

$$m_{h_1}^2 \underset{\text{tree level}}{\approx} \overbrace{M_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta} - \frac{\lambda^2}{\kappa^2} v^2 (\lambda - \kappa \sin 2\beta)^2 + \frac{3m_t^4}{4\pi^2 v^2} \left[ \ln \left( \frac{m_S^2}{m_t^2} \right) + \frac{A_t^2}{m_S^2} \left( 1 - \frac{A_t^2}{12m_S^2} \right) \right]$$

where  $m_S^2 \sim m_{Q_3}^2$

# Purpose

Study the semi-unified version of the NMSSM consistent with *at least* a fairly SM-like Higgs at 125 GeV and implications thereof.

- “Semi-unified” we mean a model which has universal  $m_0$ ,  $m_{1/2}$ , and  $A_0$  at the GUT scale with NUHM relaxation for  $m_{H_u}^2$ ,  $m_{H_d}^2$  and  $m_S^2$ , and general  $A_\lambda$  and  $A_\kappa$ , together with the parameters  $v$ ,  $\tan\beta$ ,  $\lambda$ .
- The constraints are imposed at the GUT scale and then low-scale parameters are obtained by RGE evolution.

# Basic Constraints

- Having a proper **RGE solution**, **no Landau pole** and **a neutralino LSP**.
- Higgs mass limits are from LEP, TEVATRON, and early LHC data; SUSY mass limits are essentially from LEP.
- **$B$ -physics constraints**

Observables	Constraints
$\Delta M_d$	$0.507 \pm 0.008$ ( $2\sigma$ )
$\Delta M_s$	$17.77 \pm 0.24$ ( $2\sigma$ )
$\text{BR}(B \rightarrow X_s \gamma)$	$3.55 \pm 0.51$ ( $2\sigma$ )
$\text{BR}(B^+ \rightarrow \tau^+ \nu)$	$(1.67 \pm 0.78) \times 10^{-4}$ ( $2\sigma$ )
$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$	$< 4.5 \times 10^{-9}$ (95% C.L.)

- Regarding dark matter constraints, we accept all points that have the relic density  $\Omega h^2 < 0.136$ , particularly,  $0.094 \leq \Omega h^2 \leq 0.136$  is the '**WMAP window**'.
- 2011 XENON100 bound on the spin-independent LSP–proton scattering cross section.<sup>1</sup>
- **the anomalous magnetic moment of the muon  $\delta a_\mu$**  (discussed shortly).

<sup>1</sup>For points with  $\Omega h^2 < 0.094$ , we rescale these bounds by a factor of  $0.11/\Omega h^2$ .

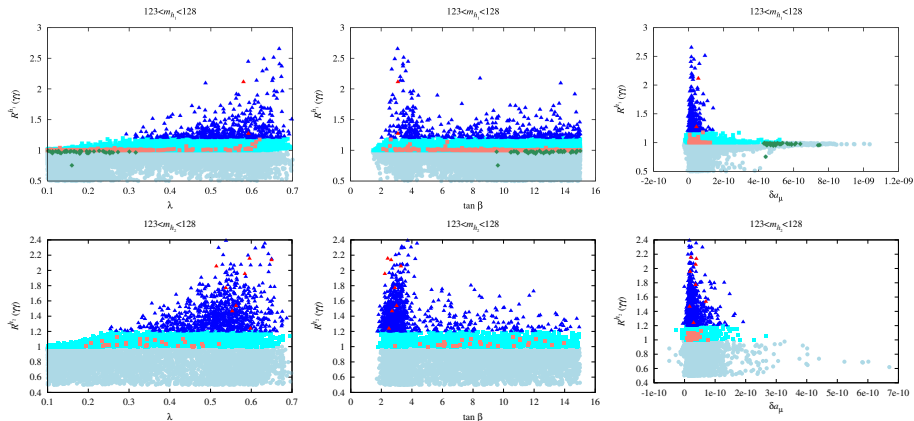
## Part I: Single Higgs Scenarios

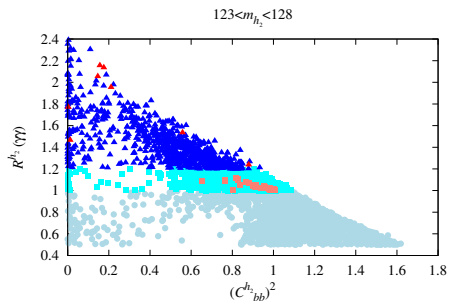
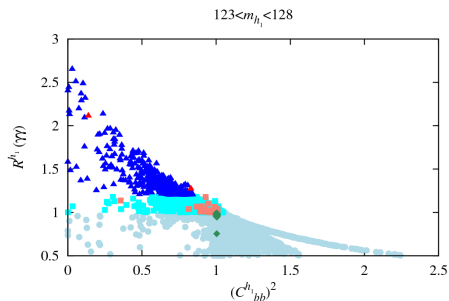
- $h_1$  or  $h_2$  either lies in the 123–128 GeV mass window.

$\gamma\gamma$  Enhancement Realization

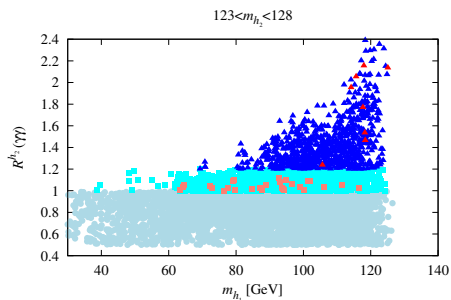
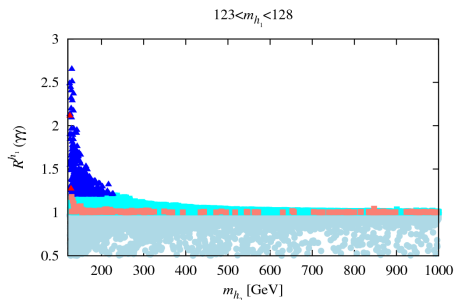
$\gamma\gamma$  enhancements appear to be possible only if

- the superpotential coupling  $\lambda$  is **large** (and  $\tan\beta$  is preferably **small**)
- the  $\delta a_\mu$  constraint is **greatly relaxed**



$\gamma\gamma$  Enhancement Mechanism

The enhancement happens when the partial width of a 125 GeV Higgs boson into  $b\bar{b}$  is strongly reduced.

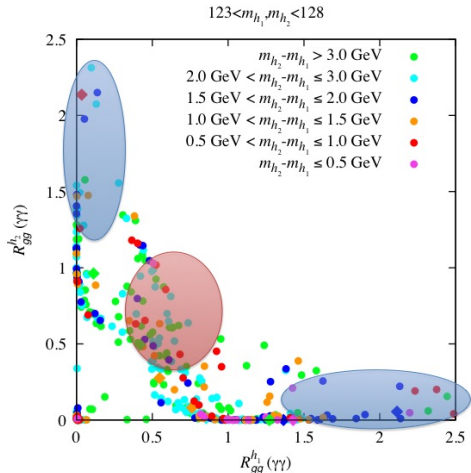
$\gamma\gamma$  Enhancement Dependence on the Higgs Mass

In addition to large  $\lambda$  (and preferably small  $\tan\beta$ ), enhanced  $\gamma\gamma$  rates are most natural when the  $h_1$  has mass similar to the second lightest CP-even Higgs,  $h_2$ , (with one of them being primarily the doublet-like  $H_u$  while the other has a large singlet  $S$  component) **with enhancement particularly likely if the  $h_1$  and  $h_2$  are degenerate.**

## Part II: Degenerate Higgs Scenarios

- $h_1$  and  $h_2$  both lie in the 123–128 GeV mass window.



Individual  $\gamma\gamma$  Rates


We combine  $h_1$  and  $h_2$  signals as follows in defining

- effective Higgs mass:

$$m_h^Y(X) \equiv \frac{R_Y^{h_1}(X)m_{h_1} + R_Y^{h_2}(X)m_{h_2}}{R_Y^{h_1}(X) + R_Y^{h_2}(X)}$$

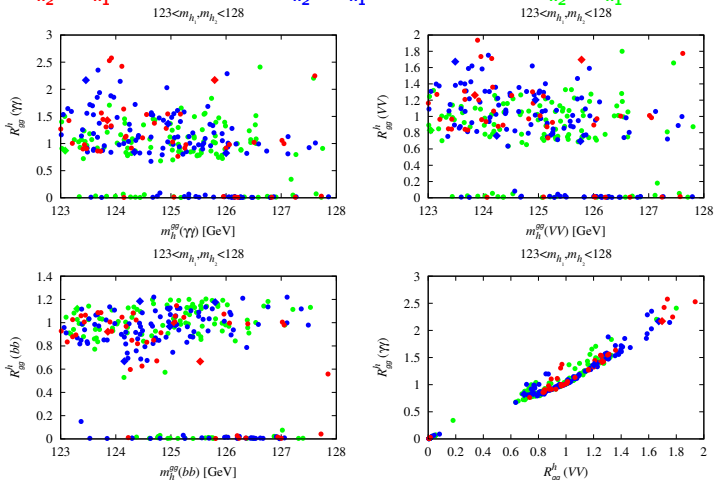
- effective Higgs signal:

$$R_Y^h(X) = R_Y^{h_1}(X) + R_Y^{h_2}(X)$$

The extent to which it is appropriate to combine the rates from the  $h_1$  and  $h_2$  depends upon the degree of degeneracy and the experimental resolution, estimated to be of order  $\sigma_{\text{res}} \sim 1.5$  GeV. The widths of the  $h_1$  and  $h_2$  are very much smaller than this resolution.

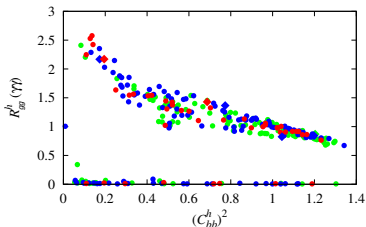
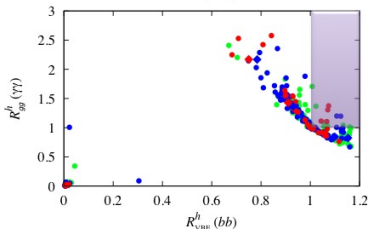
## Combined Signal Rates

•  $m_{h_2} - m_{h_1} \leq 1$  GeV; •  $1$  GeV  $< m_{h_2} - m_{h_1} \leq 2$  GeV; •  $2$  GeV  $< m_{h_2} - m_{h_1} \leq 3$  GeV



- Enhanced  $\gamma\gamma$  and  $VV$  rates from gluon fusion are very common.
- There is a very strong correlation between  $R_{gg}^h(\gamma\gamma)$  and  $R_{gg}^h(VV)$  described approximately by  $R_{gg}^h(\gamma\gamma) \sim 1.25 R_{gg}^h(VV)$ . In particular, if  $R_{gg}^h(\gamma\gamma) \sim 1.5$ , as suggested by current experimental results, then in this model  $R_{gg}^h(VV) \geq 1.2$ .

## Enhancement Mechanism – Degenerate Scenarios

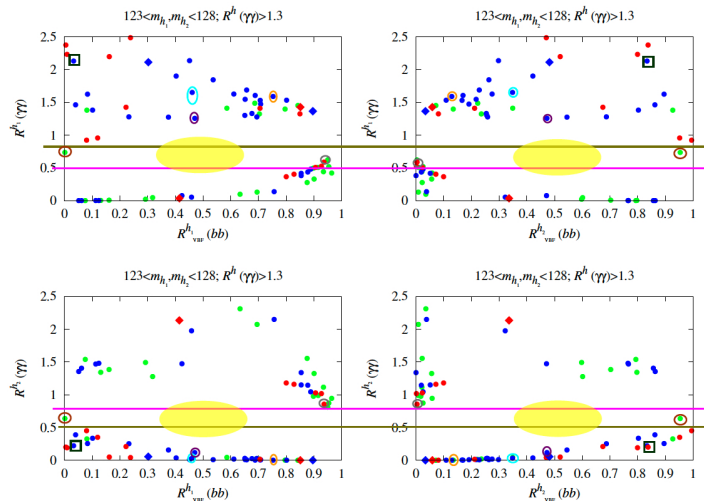
 $123 < m_{h_1}, m_{h_2} < 128$  $123 < m_{h_1}, m_{h_2} < 128$ 

## The primary mechanism:

- large net  $\gamma\gamma$  branching ratio is achieved by reducing the average total width by reducing the average  $b\bar{b}$  coupling strength.
- anti-correlation between  $R_{gg}^h(\gamma\gamma)$  and  $R_{W^* \rightarrow Wh}^h(b\bar{b}) = R_{VBF}^h(b\bar{b})$ .

- In general, the larger  $R_{gg}^h(\gamma\gamma)$  is, the smaller the value of  $R_{W^* \rightarrow Wh}^h(b\bar{b})$ .
- Enhancement of  $Wh$  production with  $h \rightarrow b\bar{b}$  is rather limited; indeed the maximal value of  $R_{VBF}^h(b\bar{b}) = R_{W^* \rightarrow Wh}^h(b\bar{b})$  is of order 1.2.
- There are parameter choices for which both the  $\gamma\gamma$  rate and the  $W^* \rightarrow Wh(\rightarrow b\bar{b})$  rate can be enhanced relative to the SM. This is a **unique** feature as a result of there being contributions to these rates from **both** the  $h_1$  and  $h_2$ .

# Various Types of Enhancement II

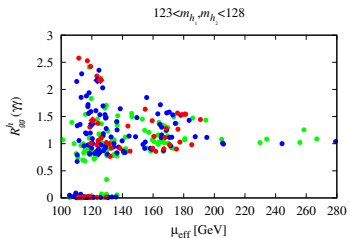
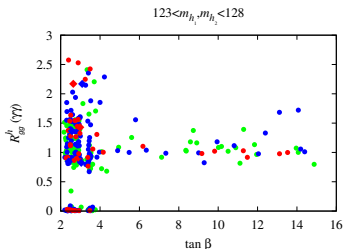
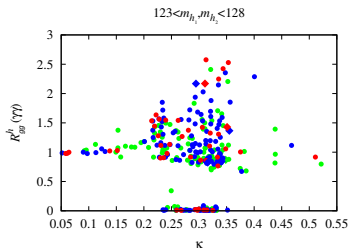
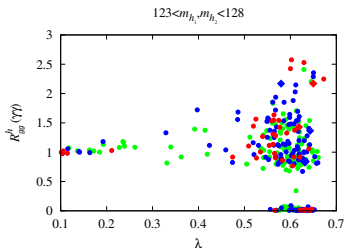


Two mechanisms to realize the enhancement:

- ① one Higgs “steals” (through mixing) some of the  $b\bar{b}$  width of the other Higgs to make the latter enhanced.
- ② either the  $\gamma\gamma$  or the  $b\bar{b}$  signal receives substantial contributions from both the  $h_1$  and the  $h_2$  for the  $\gamma\gamma$  final state) while the other final state is dominated by just one of the two Higgses.

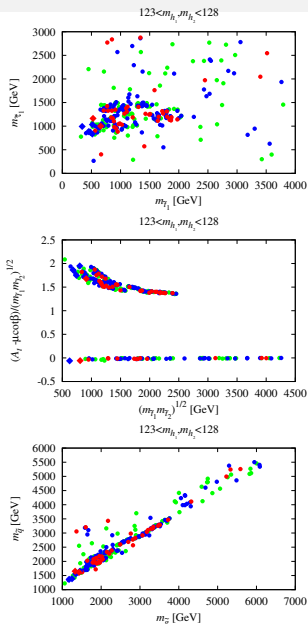
NO points where the  $\gamma\gamma$  and  $b\bar{b}$  rates both receive substantial contributions from both  $h_1$  and  $h_2$ .

## Dependence on NMSSM Parameters



The **largest**  $R_{gg}^h(\gamma\gamma)$  values arise at **large**  $\lambda$ , **moderate**  $\kappa$ , **small**  $\tan \beta < 5$  (but note that  $R_{gg}^h(\gamma\gamma) > 1.5$  is possible even for  $\tan \beta = 15$ ) and **small**  $\mu_{\text{eff}} < 150$  GeV. Such low values of  $\mu_{\text{eff}}$  are very favorable in point of fine tuning, in particular if stops are also light.

# Implication for SUSY Particles

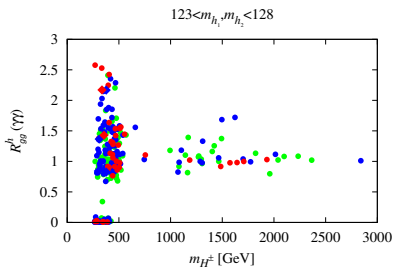
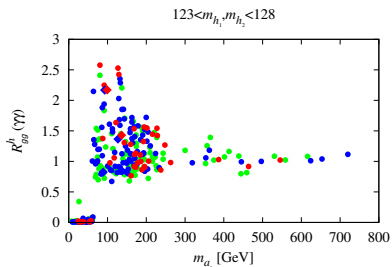


① Indeed, the few points which we found in the WMAP window always have  $m_{\tilde{\chi}_1^\pm} < 700$  GeV.

② A good fraction of our points with degenerate  $h_1, h_2$  and  $R(\gamma\gamma) > 1.3$  features light stops with  $M_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} \lesssim 1$  TeV. The stop mixing is typically large in these cases,  $(A_t - \mu_{\text{eff}} \cot \beta) / M_{\text{SUSY}} \approx 1.5$ – $2$ .

③ Squark and gluino masses are above about 1.25 TeV ranging up to as high as 6 TeV (where our scanning more or less ended). The WMAP-window points with large  $R_{gg}^h(\gamma\gamma)$  are located at low masses of  $m_{\tilde{g}} \sim 1.3$  TeV and  $m_{\tilde{q}} \sim 1.6$  TeV.

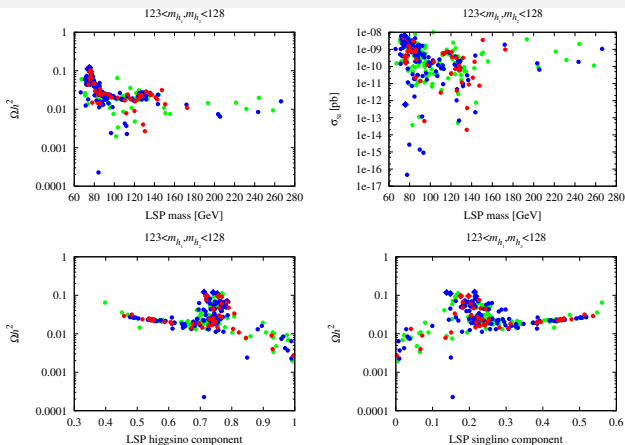
## Correlation on Other Higgs Bosons



(note that  $m_{a_2} \simeq m_{h_3} \simeq m_{H^\pm}$ )

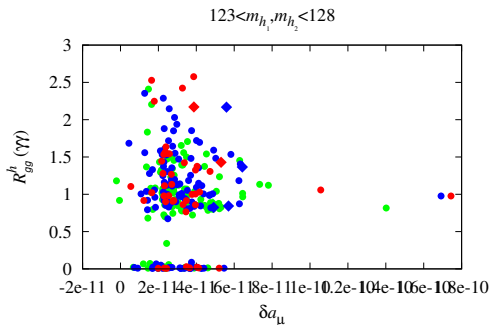
- The general trend is that the maximum  $R_{gg}^h(\gamma\gamma)$  possible decreases rapidly as  $m_{a_1}$  and  $m_{H^\pm}$  increase.
- Values above 1.7 are associated with masses for the  $a_2$ ,  $h_3$  and  $H^\pm$  of order  $\lesssim 500$  GeV and for the  $a_1$  of order  $70 \lesssim m_{a_1} \lesssim 150$  GeV, 250 GeV being the lowest allowed  $m_{H^\pm}$ .
- Although  $m_{a_1} \sim 125$  GeV is common for points with  $R_{gg}^h(\gamma\gamma) > 1$ , the contribution of the  $a_1$  to the  $\gamma\gamma$  signal is always small, typically  $R_{gg}^{a_1}(\gamma\gamma) \lesssim 0.01$  (due to large singlet component of the  $a_1$  for all  $R_{gg}^h(\gamma\gamma) > 1$  points).

## Dark Matter Properties



- WMAP-window points have a rather limited range of LSP masses, roughly  $m_{\tilde{\chi}_1^0} \in [60, 80]$  GeV.
- Corresponding  $\sigma_{SI}$  values range from *few*  $\times 10^{-9}$  pb to as low as *few*  $\times 10^{-11}$  pb.
- **large  $\Omega h^2$ :** mixed higgsino–singlino, with a singlino component of the order of 20%.  
**low  $\Omega h^2$ :** the LSP is **dominantly higgsino** (owing to small  $\mu_{\text{eff}}$ ).



What is the Status on  $\delta a_\mu$ ?

It is not possible to find scenarios of this degenerate/enhanced type while predicting a value of  $\delta a_\mu$  consistent with that needed to explain the current discrepancy. In particular, the very largest value of  $\delta a_\mu$  achieved is of order  $1.8 \times 10^{-10}$  ( $\delta a_\mu < 6 \times 10^{-11}$  for the WMAP-window).

Large  $\delta a_\mu$  only be possible if  $\lambda < 0.1$ , for which the Higgs signal in the  $\gamma\gamma$  and  $VV^*$  ( $V = W, Z$ ) final states for Higgs in the 123–128 GeV window is very SM-like.

**Interpretation:** implicitly assume that the observed discrepancy in  $a_\mu$  comes, at least in part, from a source other than the NMSSM.

## double ratio

Let us now take a look in more detail.

$$R_Y^{h_i}(X) = \frac{\sigma(Y \rightarrow h_i) \text{BR}(h_i \rightarrow X)}{\sigma(Y \rightarrow h_{\text{SM}}) \text{BR}(h_{\text{SM}} \rightarrow X)} = (C_Y^{h_i})^2 \frac{\Gamma(h_i \rightarrow X)}{\Gamma(h_{\text{SM}} \rightarrow X)} \frac{\Gamma_{\text{tot}}(h_{\text{SM}})}{\Gamma_{\text{tot}}(h_i)} = (C_Y^{h_i})^2 (C_X^{h_i})^2 \dots$$

where  $Y = gg$  for gluon fusion and  $Y = WW, ZZ$  for  $W, Z$  fusion and  $W, Z$  strahlung, this latter also implies  $R_{VBF}^{h_i}(X) = R_{V^* \rightarrow VH}^{h_i}(X)$  and  $C_\Gamma^{h_i}$  is the ratio of the  $h_i$  total width the SM Higgs total width.

### The diagnostic tools

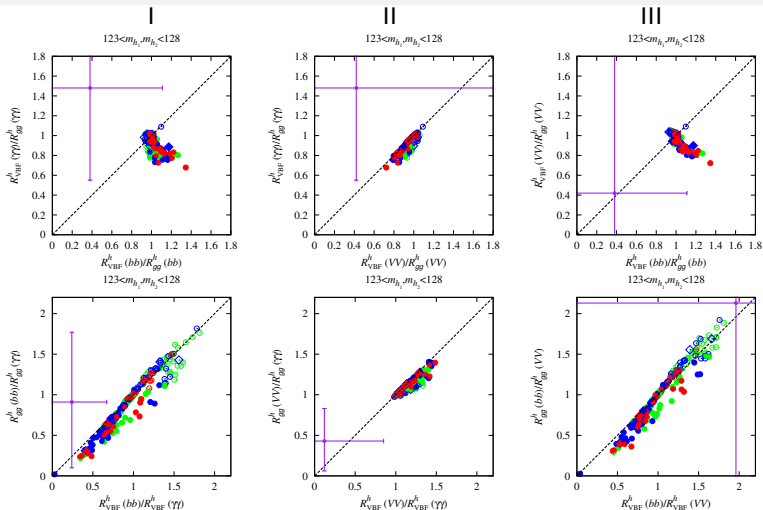
we propose the **double ratios** to reveal the existence of a second, quasi-degenerate (but non-interfering in the small width approximation) Higgs state are :

$$\text{I): } \frac{R_{VBF}^h(\gamma)/R_{gg}^h(\gamma)}{R_{VBF}^h(bb)/R_{gg}^h(bb)}, \quad \text{II): } \frac{R_{VBF}^h(\gamma)/R_{gg}^h(\gamma)}{R_{VBF}^h(WW)/R_{gg}^h(WW)}, \quad \text{III): } \frac{R_{VBF}^h(WW)/R_{gg}^h(WW)}{R_{VBF}^h(bb)/R_{gg}^h(bb)}$$

each of which should be unity if only a single Higgs boson is present but, due to the non-factorizing nature of the sum, are generally expected to deviate from 1 if two (or more) Higgs bosons are contributing to the net  $h$  signals.

Of course, the above three double ratios are not all independent. Which will be most useful depends upon the precision with which the  $R^h$ 's for different initial/final states can be measured.

## double ratio examination – NMSSM (i)



- III) is very like I) due to the correlation between the  $R_{gg}^h(\gamma\gamma)$  and  $R_{gg}^h(WW)$  values.
- Any one of these double ratios will often, but not always, deviate from unity.
- The probability of such deviation increases dramatically if we require  $R_{gg}^h(\gamma\gamma) > 1$ .

## What does current LHC data say about these various double ratios?

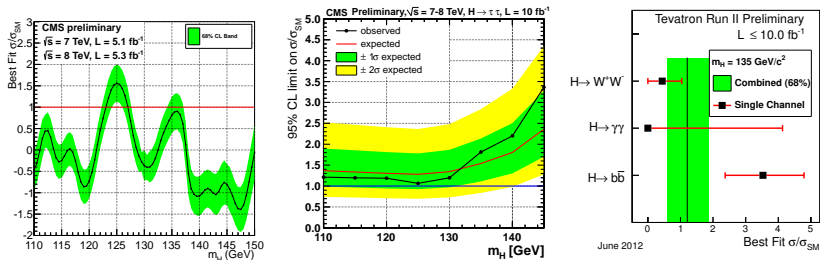
- Obviously, current statistics are inadequate to discriminate whether or not the double ratios deviate from unity.
- For a  $\sqrt{s} = 14$  TeV run with  $L = 100 \text{ fb}^{-1}$  ( $300 \text{ fb}^{-1}$ ), the number of Higgs events will be about a factor of 25 (77) larger than the number produced for  $L \sim 5 \text{ fb}^{-1}$  at  $\sqrt{s} = 7$  TeV plus  $L \sim 6 \text{ fb}^{-1}$  at  $\sqrt{s} = 8$  TeV. It means the error bars should be reduced by roughly a factor of 5 (9), levels that could indeed reveal a deviation from unity, or at least remove some model points if no deviation within that error is seen.
- Of course improvements in the experimental analyses may further increase the sensitivity.

We thus conclude that our diagnostic tools will ultimately prove viable and perhaps crucial for determining if the  $\sim 125$  GeV Higgs signal is really only due to a single Higgs-like resonance or if two resonances are contributing, the latter having significant probability in model contexts if enhanced  $\gamma\gamma$  rates are indeed confirmed at higher statistics.

## Part III: Multiple Higgs Scenarios

- $h_1 \sim 98$  GeV (LEP) and  $h_2 \sim 125$  GeV
- $h_1 \sim 125$  GeV and  $h_2 \sim 136$  GeV (Tevatron)

## 136 GeV Higgs-like signal



- 1 CMS observed an **additional excess** corresponding to  $M_H \sim 136$  GeV in the  $\gamma\gamma$  decay mode.
- 2 CMS observed **a deficit** for  $M_H \sim 125$  GeV for the *VBF*-tag class of events, but **an excess** for  $M_H \gtrsim 132$  GeV in the  $\tau^+\tau^-$  decay mode.
- 3 The best fit to the measurement of  $M_H$  at the Tevatron corresponds to  $M_H \sim 135$  GeV with an enhanced signal rate  $3.53^{+1.26}_{-1.16}$  in the  $b\bar{b}$  decay mode with *WH*-tag.

$m_h = 136$	gg fusion			VBF	VH
	$\gamma\gamma$	$ZZ^* \rightarrow 4\ell$	$WW^* \rightarrow \ell\nu\ell\nu$	$\tau^+\tau^-$	$b\bar{b}$
ATLAS	$0.0^{+0.4}$				
CMS	$0.9 \pm 0.4$				
Combined	$0.45 \pm 0.3$	$\leq 0.2$	$\times$	$< 1.81$	$\times$
	high resolution		poor resolution		

$WW^{(*)}$  channel (with *VH*-tag) is ignored due to its low mass resolution.

Effective  $b\bar{b}$  Signal

The central value 3.53 in the  $b\bar{b}$  is difficult to explain.

- the  $VH$  production cross section  $\propto (C_{VV}^{h_i})^2$  cannot be enhanced with respect to the SM.
- the SM Higgs branching fraction of  $\sim 40\%$  for  $m_{h_{\text{SM}}} = 135$  GeV can be enhanced at most by a factor of 2.5 in the unphysical limit  $C_{dd}^{h_i} \rightarrow \infty$ .

## 125+136 LHC–Tevatron scenario

- If  $R_{VBF}^{h_1}(\tau\tau)$  is as small as observed by CMS, the values for  $R_{VH}(b\bar{b})$  measured at the Tevatron should originate primarily from  $h_2$  with  $m_{h_2} \sim 135\text{--}136$  GeV.  $\checkmark$

This possibility is one of the main advantages.

- However, the contribution of  $h_1$  to the signal rate  $R_{VH}(b\bar{b})$  obtained assuming  $m_{h_{\text{SM}}} \sim 135$  GeV can still be sizable, since the production cross section of  $h_1$  is  $\sim 30\%$  larger.

$$R_{VH}^{\text{eff}}(b\bar{b}) \simeq R_{VH}^{h_2}(b\bar{b}) + 1.3 \times R_{VH}^{h_1}(b\bar{b})$$

- In addition, the contribution from  $h_2$  to the signal rate  $R_{VH}(b\bar{b})$  assuming  $m_{h_{\text{SM}}} \sim 125$  GeV should be as large as possible.

## What are the consequences on the reduced couplings?

	gg fusion			VBF		VH
$m_{h_1} = 125$	$\gamma\gamma$	$ZZ^* \rightarrow 4\ell$	$WW^* \rightarrow 2\ell 2\nu$	$\gamma\gamma$	$\tau^+\tau^-$	$b\bar{b}$
Combined	$1.66 \pm 0.36$	$1.02 \pm 0.38$	$\times$	$> 1$	very small	$1.97^{+0.74}_{-0.69}$
	high resolution			poor resolution		

① keeping  $C_{gg}^{h_1} \sim C_{uu}^{h_1}$  not be small,  $C_{dd}^{h_1}$  reduced  $\implies \Gamma(h_1 \rightarrow b\bar{b}) \implies \Gamma_{\text{tot}}(h_1)$

$$\implies \begin{cases} \text{enhanced } R_{gg}^{h_1}(\gamma\gamma) \quad \checkmark \\ \text{enhanced } R_{gg}^{h_1}(ZZ), \text{ together with a slightly reduced } C_{ZZ}^{h_1} \quad \checkmark \end{cases}$$

②  $C_{dd}^{h_1}$  reduced  $\implies C_{\tau\tau}^{h_1}$  reduced  $\implies$  small  $R_{VBF}^{h_1}(\tau\tau) \quad \checkmark \implies$  small  $R_{VH}^{h_1}(b\bar{b}) \quad \times$

	gg fusion			VBF	VH
$m_{h_2} = 136$	$\gamma\gamma$	$ZZ^* \rightarrow 4\ell$	$WW^* \rightarrow \ell\nu\ell\nu$	$\tau^+\tau^-$	$b\bar{b}$
Combined	$0.45 \pm 0.3$	$\leq 0.2$	$\times$	$< 1.81$	$3.53^{+1.26}_{-1.16}$
	high resolution			poor resolution	

③ then need  $C_{dd}^{h_2}$  enhanced  $\implies$  large  $R_{VH}^{\text{eff}}(b\bar{b}) \quad \checkmark$ , in addition that  $C_{VV}^{h_2}$  is not small.

④ small  $C_{gg}^{h_2} \sim C_{uu}^{h_2} \implies$  small  $R_{gg}^{h_2}(\gamma\gamma)$  and  $R_{gg}^{h_2}(ZZ) \quad \checkmark$



General NMSSM **DOES** Work—an exemplary point

NMSSM parameters (The dimensionful parameters are given in GeV):

$\lambda$	$\kappa$	$\tan \beta$	$\mu_{\text{eff}}$	$A_\lambda$	$A_\kappa$
0.617	0.253	1.77	143	164	337

general NMSSM specific:

$M_1$	$M_2$	$M_3$	$m_{\tilde{q}_{1,2}}, m_{\tilde{b}_R}$	$m_{\tilde{\ell}}, m_{\tilde{q}_3}$	$A_t = A_b$	$A_\tau$
220	400	1100	1500	1000	-2500	-1000

Higgs masses and component decompositions:

$m_{h_1}$	$m_{h_2}$	$m_{h_3}$	$m_{a_1}$	$m_{a_2}$	$m_{H^\pm}$
125	136	289	95	282	272

Higgs	$S_{i,d}$	$S_{i,u}$	$S_{i,s}$	$C_{dd}^{h_i}$	$C_{uu}^{h_i}$	$C_{VV}^{h_i}$	$C_{gg}^{h_i}$	$C_{\gamma\gamma}^{h_i}$
$h_1$	-0.24	-0.67	0.70	-0.48	-0.77	-0.70	0.77	0.85
$h_2$	0.54	0.51	0.67	1.09	0.58	0.71	0.54	0.66
$h_3$	0.81	-0.54	-0.24	1.64	-0.62	-0.07	0.65	0.28

Signal rates:

Higgs	$R_{gg}(\gamma\gamma)$	$R_{VBF}(\gamma\gamma)$	$R_{gg}(VV)$	$R_{VH}(VV)$	$R_{VH}(b\bar{b})$	$R_{gg}(\tau\tau)$
$h_1$	1.30	1.09	0.90	0.75	0.36	0.42
$h_2$	0.16	0.27	0.18	0.31	0.74	0.43
$h_3$	0.58	0.01	0.04	0.004	0.23	19.6

$R_{VH}^{\text{eff}}(b\bar{b}) \sim 1.20$ . Note that  $R_{VBF}(VV) = R_{VH}(VV)$  and  $R_{VBF}(\tau\tau) \sim R_{VH}(b\bar{b})$ .

## How about in the semi-constrained NMSSM?

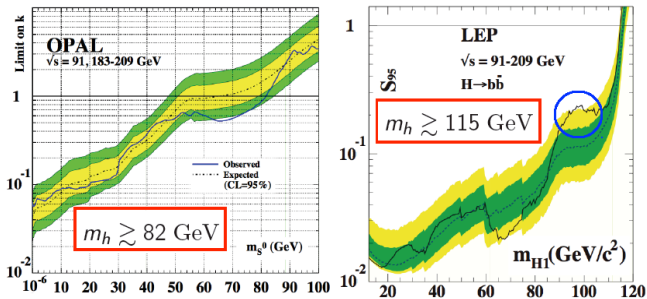
- Can find scenarios where  $h_1$  and  $h_2$  have masses of about 125 and 136 GeV, respectively.
- However, we did **NOT** find any points where the slightly enhanced  $R_{gg}^{h_1}(ZZ)$ , the substantially enhanced  $R_{gg}^{h_1}(\gamma\gamma)$ , the suppressed  $R_{gg}^{h_2}(\gamma\gamma)$  and very small  $R_{gg}^{h_2}(ZZ)$  are **all satisfied simultaneously**. at least one of the last three conditions has to be relaxed to find valid points.

	$R_{gg}^{h_1}(\gamma\gamma)$	$R_{gg}^{h_1}(ZZ)$	$R_{gg}^{h_2}(\gamma\gamma)$	$R_{gg}^{h_2}(ZZ)$	$R_{VBF}^{h_2}(\tau\tau)$	$R_{VH}^{h_2}(b\bar{b})$
case I	✓	✓	$\lesssim 0.06$	✓	✓	
case II	$\lesssim 1.3$	✓	✓	✓	✓	never large

*Maybe case I is a good thing in the end?*

## LEP Excess around 98 GeV

The LEP excess is clearly inconsistent with a SM-like Higgs boson around 98 GeV, being only about 10 – 20% of the rate predicted for the  $h_{SM}$ .



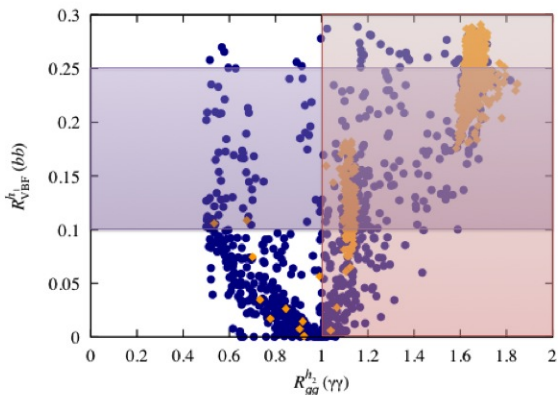
An interesting question:

whether the LHC signal and the small ( $\sim 2\sigma$ ) LEP excess in  $e^+e^- \rightarrow Zb\bar{b}$  in the vicinity of  $M_{b\bar{b}} \sim 98$  GeV using the  $h'$  with  $m_{h'} \sim 98$  GeV could be **simultaneously** explained?

## 98 + 125 GeV LEP-LHC scenarios

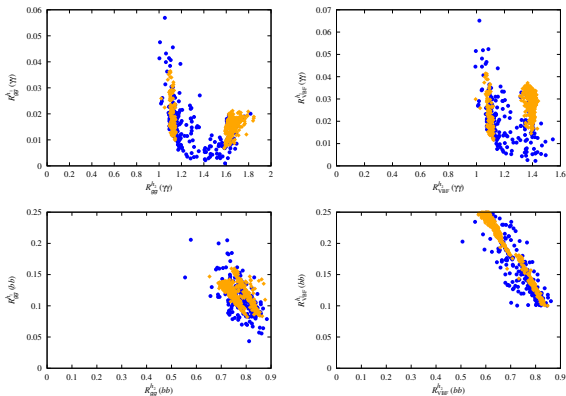
## 98 +125 LEP-LHC scenario

- Consistency with such a result for the  $h'$  is natural if the  $h'$  couples at a reduced level to  $ZZ$ , which, in turn, is automatic if the  $h$  has substantial  $ZZ$  coupling, as required by the observed LHC signals.
- To describe the LEP and LHC data under the NMSSM framework, the  $h_1$  and  $h_2$  must have  $m_{h_1} \sim 98$  GeV and  $m_{h_2} \sim 125$  GeV, respectively, with the  $h_1$  being largely singlet and the  $h_2$  being primarily doublet (mainly  $H_u$  for the scenarios we consider).
- A 125 GeV Higgs state  $h_2$  with enhanced  $\gamma\gamma$  signal rate is easily obtained for large  $\lambda$  and small  $\tan\beta$ .

LEP-LHC Fit ( $m_{h_1} \in [96, 100]$  GeV,  $m_{h_2} \in [123, 128]$  GeV)

- Those points with  $R_{VBF}^{h_1}(bb)$  between about 0.1 and 0.25 would provide the best fit to **the LEP excess**. ( $R_{VBF}^{h_1}(bb)$  is equivalent to  $R_{Vh_1}^{h_1}(bb)$  as relevant for LEP.)
- A large portion of such points have  $R_{gg}^{h_2}(\gamma) > 1$  as preferred by **LHC data**.

In all the remaining plots:  $R_{gg}^{h_2}(\gamma) > 1$  and  $0.1 \leq R_{VBF}^{h_1}(bb) \leq 0.25$ .

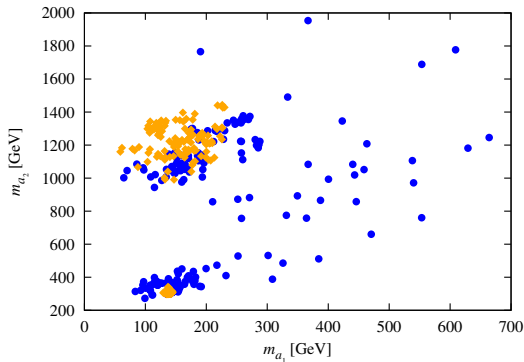
Enhanced  $\gamma\gamma$  Rates

- $h_2$  can easily have an enhanced  $\gamma\gamma$  signal for both  $gg$  and VBF production.
- The  $\gamma\gamma$  signal arising from the  $h_1$  for both production mechanisms is quite small and unlikely to be observable.
- $R_{gg}^{h_2}(bb)$  and  $R_{VBF}^{h_2}(bb)$  values that are associated with reduced  $b\bar{b}$  width (relative to the SM) are reduced to have enhanced  $R_{gg}^{h_2}(\gamma\gamma)$  and  $R_{VBF}^{h_2}(\gamma\gamma)$ .
- $R_{gg}^{h_1}(bb)$  and  $R_{VBF}^{h_1}(bb)$  values are such that the  $h_1$  could not yet have been seen at the Tevatron or LHC.
- For WMAP-window points the largest  $R_{VBF}^{h_1}(bb)$  values occur for the light- $m_{\tilde{\chi}_1^0}$  point group II.

	$R_{gg}^{h_2}(\gamma\gamma)$	$m_{\tilde{\chi}_1^0}$	$m_{\tilde{t}_1}$	
I	$\sim 1.1$	$> 93$ GeV	$> 1.8$ TeV	
II	$\sim 1.6$	$\sim 77$ GeV	197 GeV – 1 TeV	✓

If  $R_{gg}^{h_2}(\gamma\gamma)$  ends up converging to a large value, then masses for all strongly interacting SUSY particles would be close to current limits if the present 98 + 125 GeV LEP-LHC Higgs scenario applies.

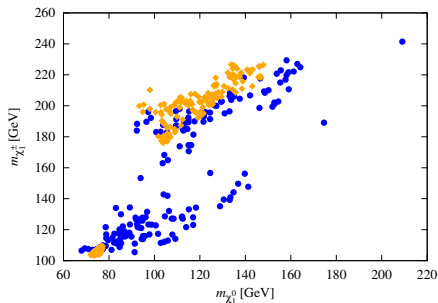
## Expectation for Other NMSSM Higgses



$m_{h_3} \simeq m_{H^\pm} \simeq m_{a_2}$  for the scenarios considered.

Small  $m_{a_1}$  is typical of the WMAP-window points.

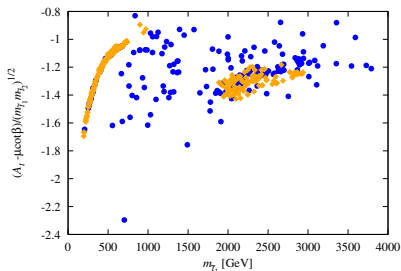
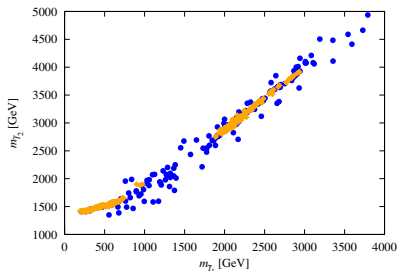
## Prediction for SUSY Particles



- Low values of  $m_{\tilde{\chi}_1^0}$  are typical for the scan points, but more particular to this model are the rather low values of  $m_{\tilde{\chi}_1^\pm}$ .
- ATLAS and CMS are currently performing analyses, but the difficulty in isolating the leptons or jets associated with  $\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 + X$  decays would arise when  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0}$  is small.
- For the WMAP-window points  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0}$  is typically quite substantial, at least 35 GeV for the low- $m_{\tilde{\chi}_1^0}$  points.

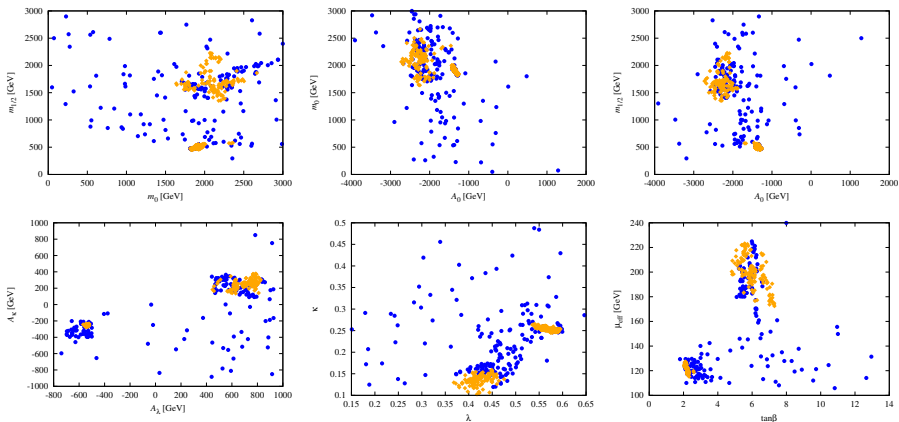


## Prediction for SUSY Particles



- Of particular interest is the very large range of  $m_{\tilde{t}_1}$ . We observed the possibility of  $m_{\tilde{t}_1}$  as small as 197 GeV and mostly modest values of the mixing parameter  $(A_t - \mu \cot \beta) / \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$ .
- For lighter values of  $m_{\tilde{t}_1}$  (typical of the low- $m_{\tilde{\chi}_1^0}$  WMAP points), the  $\tilde{t}_1$  always decays via  $\tilde{t}_1 \rightarrow \tilde{\chi}_1^+ b$  or  $\tilde{t}_1 \rightarrow \tilde{\chi}_1^0 t$ , the latter being absent when  $m_{\tilde{t}_1} < m_{\tilde{\chi}_1^0} + m_t$ .
- At high  $m_{\tilde{t}_1}$ , these same channels are present but also  $\tilde{t}_1 \rightarrow \tilde{\chi}_{2,3,4,5}^0 t$  can be important, which channels being present depending upon whether  $m_{\tilde{t}_1} - m_{\tilde{\chi}_{2,3,4,5}^0} - m_t > 0$  or not.

## NMSSM Parameters



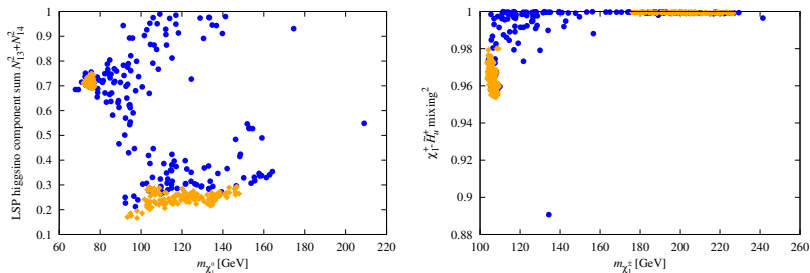
No particular regions of these parameters appear to be singled out aside from some preference for negative values of  $A_0$ .

Low  $\mu_{\text{eff}} \implies$  not much fine-tuning.

- Large  $m_{\tilde{\chi}_1^0}$  group (II):  $\tan\beta \in [5, 7]$  and  $\lambda \in [0.37, 0.48]$ .
- Low  $m_{\tilde{\chi}_1^0}$  group (I): large Higgsino component with  $\tan\beta \in [2, 2.6]$  and  $\lambda \in [0.53, 0.6]$ .

## LSP Composition

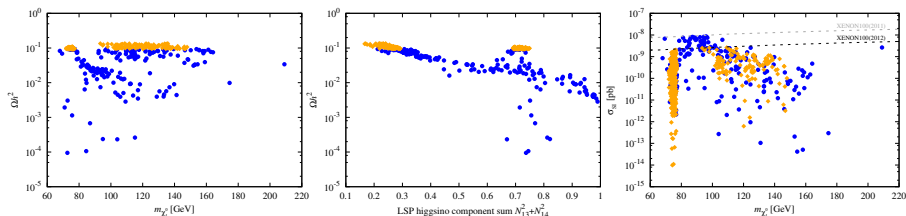
The composition of the  $\tilde{\chi}_1^0$  and the  $\tilde{\chi}_1^\pm$  are crucial when it comes to the relic density of the  $\tilde{\chi}_1^0$  (LSP dark matter candidate).



in the WMAP window

- Low  $m_{\tilde{\chi}_1^0}$  group (I), the  $\tilde{\chi}_1^0$  can have a large Higgsino fraction since the  $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow W^+ W^-$  annihilation mode (mainly via  $t$ -channel exchange of the light Higgsino-like chargino) is below threshold.
- $m_{\tilde{\chi}_1^0} > 93$  GeV group (II), the points can lie in the WMAP window only if the  $\tilde{\chi}_1^0$  does **NOT** have a large Higgsino fraction, in the other words, the LSP is **dominantly singlino** (under the approximation that **the singlino fraction = 1 - Higgsino fraction**).

## Dark Matter Properties



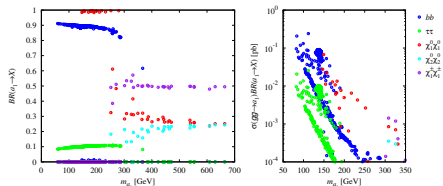
Why the relic density  $\Omega h^2$  too small?

- The main mechanism is rapid  $\tilde{\chi}_1^0 \tilde{\chi}_1^0$  annihilation to  $W^+ W^-$  due to a substantial Higgsino component of the  $\tilde{\chi}_1^0$ . Indeed, the relic density of a Higgsino LSP is typically of order  $\Omega h^2 \approx 10^{-3} - 10^{-2}$ .
- $\Omega h^2$  increases as the Higgsino component declines.

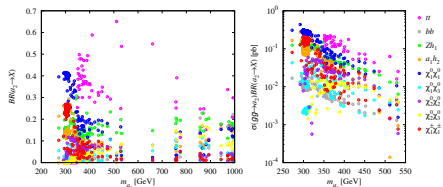
Incidentally, the 2012 XENON100 limits on the spin-independent cross section  $\sigma_{SI}$  are obeyed by all the WMAP points. Experiments will probe some of the  $\sigma_{SI}$  values that survive 2012 limits relatively soon, especially the  $m_{\tilde{\chi}_1^0} > 93$  GeV points that are in the WMAP window. However, it is also noteworthy that the  $m_{\tilde{\chi}_1^0} \sim 75$  GeV WMAP-window points can have very small  $\sigma_{SI}$ .

Future Test: Confirm or Rule out?

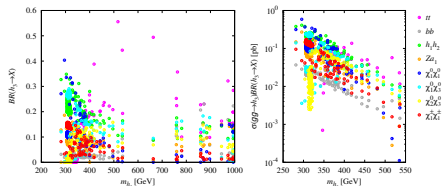
## LHC Test – Direct Higgs Production and Decay

a<sub>1</sub>

- singlet dominant, so rather small production rates
- The largest  $\sigma\text{BR}(X)$ :  $X = b\bar{b}$  final state (have huge backgrounds)
- $\sigma\text{BR}(X)$  for  $X = \tilde{\chi}_1^0\tilde{\chi}_1^0$  can be significant (when allowed)

a<sub>2</sub>

- doublet dominant, so better discovery prospects
- $m_{a_2} > 2m_t$ ,  $t\bar{t}$  final state:  $\sigma\text{BR} > 0.01$  pb for  $m_{a_2} < 550$  GeV
- $m_{a_2} < 2m_t$ ,  $a_1h_2$  final state with both  $a_1$  and  $h_2$  decaying to  $b\bar{b}$  might be visible.

h<sub>3</sub>

- the possibly visible final states are  $t\bar{t}$  for  $m_{h_3} > 2m_t$  and  $h_1h_2$  for  $m_{h_3} < 2m_t$

For both the  $a_2$  and  $h_3$ ,  $\sigma\text{BR}(X)$  is substantial for  $X = \tilde{\chi}_1^0\tilde{\chi}_1^0$ , but to isolate this invisible final state would require an additional photon or jet tag.

# Conclusions

We have assessed the extent to which various semi-constrained NMSSM (scNMSSM) scenarios with *at least* a  $\sim 125$  GeV Higgs  $h_1$  are able to describe this LHC signal.

- we proposed a novel idea — “degenerate Higgs” and examined scNMSSM scenarios in which both  $h_1$  and  $h_2$  have mass near 125 GeV. As what we expected, very substantially enhanced  $\gamma\gamma$  and other signals are possible.
- we developed diagnostic tools that would provide incontrovertible evidence for the presence of more than one Higgs near 125 GeV in the LHC data.
- In addition, we studied interesting multiple Higgs scenarios: (i) 125+136 LHC–Tevatron scenario, the best fit to the Tevatron results in the  $b\bar{b}$  channel and the mild excesses at CMS in the  $\gamma\gamma$  channel at 136 GeV and in the  $\tau\tau$  channel above 132 GeV can be explained by  $h_2$  in this mass range, together with  $h_1$  at 125 GeV discovered at the LHC. (ii) 98+125 LEP–LHC scenario,  $h_1$  is consistent with the small LEP excess at 98 GeV and  $h_2$  has the primary features of the LHC Higgs-like signals at 125 GeV, including an enhanced  $\gamma\gamma$  rate.

- The phenomenological NMSSM is a natural extension.

*Instead of being the end of story, the recent discovery of the 125 GeV Higgs-like signal has brought particle physics research into the start of a new era. We are in the midst of an exciting debate on the nature of the 125 GeV state.*

*We are currently waiting to see if the future LHC data supports the various multi-Higgs proposals outlined earlier, or, alternatively, suggests that alternative theories are Nature's choice.*



# Thank you for your attention!

Thanks to Profs. Gunion and Kraml (and etc.) for their patient guidance and help, and also for their strong recommendations for my 2013 LHC-TI Fellowship application.

Happy Thanksgiving!

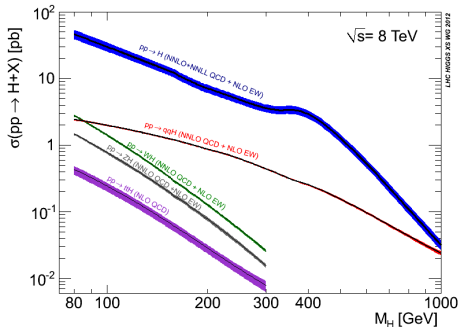
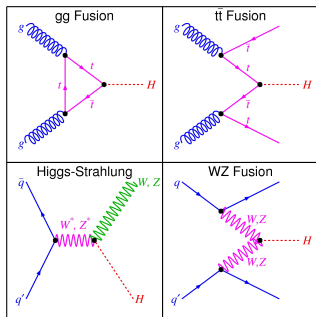
Back Up

# SM Higgs Boson Production

LHC:

- gluon fusion ( $ggF$ ) and vector boson fusion (VBF) – main production.
- $V^* \rightarrow VH$  ( $V = W$  or  $Z$ ) with  $H \rightarrow b\bar{b}$  and  $VBF \rightarrow H$  with  $H \rightarrow \tau^+\tau^-$ .

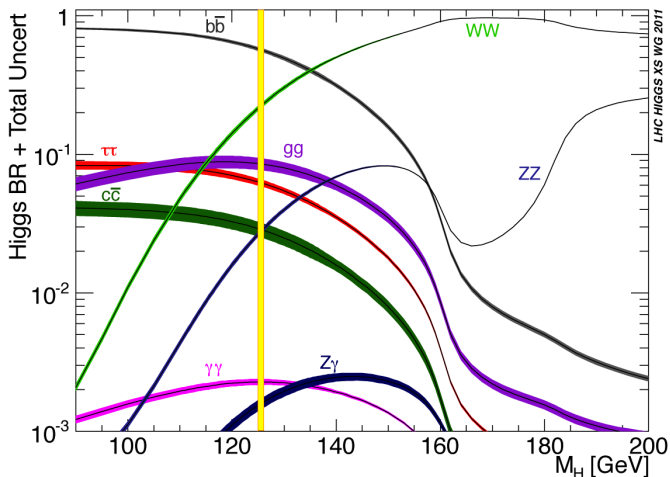
Tevatron:  $V^* \rightarrow VH$  ( $V = W$  or  $Z$ ) with  $H \rightarrow b\bar{b}$  only.



$gg \rightarrow H$ : Good for  $WW$ ,  $ZZ$ ,  $\gamma\gamma$  final states; Bad for  $H \rightarrow b\bar{b}$  (overwhelming QCD backgrounds!)

$qq \rightarrow V^* \rightarrow VH$ : Good for  $H \rightarrow b\bar{b}$  final states in the Leptons + Jets search ( $W/Z$  boson decays to leptons which are straightforward to select).

# SM Higgs Boson Decay

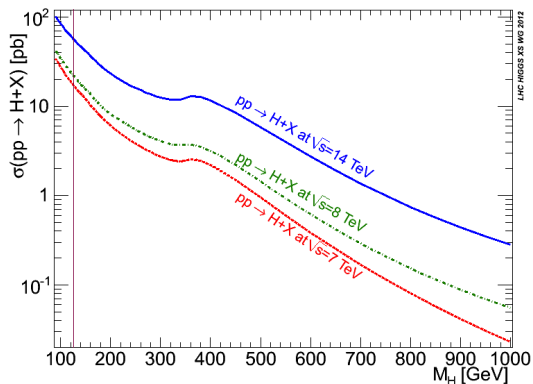


$ZZ$ ,  $\gamma\gamma$ ,  $Z\gamma$ : small branching ratio but clean signatures and NO missing energy.

$WW$ : more sizable branching ratio; two leptons + missing energy.

$b\bar{b}$ : largest branching ratio; quark hadronization into jets.

# Higgs Production Comparison



the expected increase in the number of Higgs events at  $\sqrt{s} = 14$  TeV for integrated luminosity of  $100 \text{ fb}^{-1}$  will be a factor of roughly

$$100 \times 49.85 / (5 \times 15.32 + 6 \times 19.52) = 25$$

larger than for the  $(5 \text{ fb}^{-1}, 7 \text{ TeV} + 6 \text{ fb}^{-1}, 8 \text{ TeV})$  results.

This implies a decrease in the error bars by a factor of order  $1/\sqrt{25} \sim 1/5$  based just on

## More on Rs in the doublets+singlets Models

$$R_Y^{h_i}(X) = \frac{\sigma(Y \rightarrow h_i) \text{BR}(h_i \rightarrow X)}{\sigma(Y \rightarrow h_{\text{SM}}) \text{BR}(h_{\text{SM}} \rightarrow X)} = (C_Y^{h_i})^2 \frac{\Gamma(h_i \rightarrow X)}{\Gamma(h_{\text{SM}} \rightarrow X)} \frac{\Gamma_{\text{tot}}(h_{\text{SM}})}{\Gamma_{\text{tot}}(h_i)} = (C_Y^{h_i})^2 (C_X^{h_i})^2 \dots$$

where  $Y = gg$  for gluon fusion and  $Y = WW, ZZ$  for  $W, Z$  fusion and  $W, Z$  strahlung, this latter also implies  $R_{VBF}^{h_i}(X) = R_{V^* \rightarrow VH}^{h_i}(X)$ .

### Assumptions:

- 1 in SUSY models,  $C_{\tau\tau} \approx C_{dd}$  (up to radiative corrections relevant for very large  $\tan\beta$ ).
- 2  $C_{WW}^{h_i} = C_{ZZ}^{h_i} \leq 1$  (the custodial symmetry stored in any doublets+singlets models).
- 3 The radiatively induced coupling to gluons originates mostly from top quark loops,  $C_{gg}^{h_i} \approx C_{tt}^{h_i}$ .
- 4 The radiatively induced coupling to photons originates mostly from the  $W$  loop,  $C_{\gamma\gamma}^{h_i} \approx C_{WW}^{h_i}$ , neglecting the much smaller top loop contribution and loops involving non-SM particles.

NOT all the  $R^{h_i}$ 's in the context of any doublets plus singlets model are independent.

A complete set:  $R_{gg}^h(WW), R_{gg}^h(bb), R_{gg}^h(\gamma\gamma), R_{VBF}^h(WW), R_{VBF}^h(bb), R_{VBF}^h(\gamma\gamma)$

# The Scale-invariant NMSSM

NMSSM solves  $\mu$ -problem by adding one singlet  $S$ , at the cost of adding 3 more particles

$$\mathcal{L}_{\text{NMSSM}} = \mathcal{L}_{\text{kinetic}} + \mathcal{L}_{\text{int}} + \mathcal{L}_{\text{soft}}^{\text{NMSSM}}$$

The interactions are generated by the superpotential

$$W_{\text{NMSSM}} = \bar{u} Y_u Q H_u - \bar{d} Y_d Q H_d - \bar{e} Y_e L H_d + \lambda S H_u H_d + \frac{\kappa}{3} S^3$$

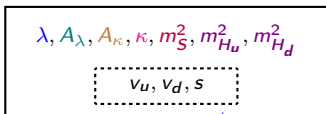
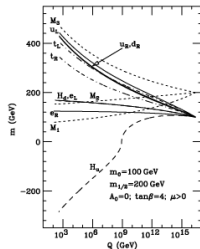
and the soft SUSY breaking terms are

$$\mathcal{L}_{\text{soft}} \left\{ \begin{array}{l} \mathcal{L}_{\text{gaugino}} = -\frac{1}{2} \left( M_3 \tilde{G}^a \tilde{G}_a + M_2 \tilde{W}^\alpha \tilde{W}_\alpha + M_1 \tilde{B} \tilde{B} \right) + \text{h.c.} \\ \mathcal{L}_{\text{sfermions}} = -\tilde{Q}_L^* m_Q^2 \tilde{Q}_L - \tilde{L}_L^* m_L^2 \tilde{L}_L - \tilde{u}_R^* m_U^2 \tilde{u}_R - \tilde{d}_R^* m_D^2 \tilde{d}_R - \tilde{e}_R^* m_E^2 \tilde{e}_R \\ \mathcal{L}_{\text{Higgs}} = -m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - m_S^2 S^* S \\ \mathcal{L}_{\text{trilinear}} = -\left( \tilde{u}_R A_u \tilde{Q}_L H_u - \tilde{d}_R A_d \tilde{Q}_L H_d - \tilde{e}_R A_e \tilde{L}_L H_d + \lambda A_\lambda H_u H_d S + \frac{1}{3} \kappa A_\kappa S^3 \right) \\ \quad + \text{h.c.} \end{array} \right.$$

$\mathbb{Z}_3$ -symmetry: a multiplication of all components of chiral superfields by a phase  $e^{2\pi i/3}$ .

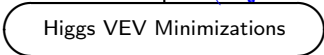
# NMSSM Parameters

- GUT scale parameters (assume unification)
  - Gaugino masses:  $m_{1/2} \rightarrow M_1, M_2, M_3$
  - Squark masses:  $m_0 \rightarrow m_{\tilde{Q}}^2, m_{\tilde{L}}^2, m_{\tilde{U}}^2, m_{\tilde{D}}^2, m_{\tilde{E}}^2$
  - Trilinear couplings:  $A_0 \rightarrow A_u, A_d, A_e$
- SUSY scale parameters

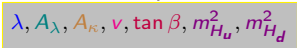


$$v_u \left( m_{H_u}^2 + \mu_{\text{eff}}^2 + \lambda^2 v_d^2 + \frac{g_1^2 + g_2^2}{4} (v_u^2 - v_d^2) \right) - v_d \mu_{\text{eff}} (A_\lambda + \kappa s) = 0$$

$$v_d \left( m_{H_d}^2 + \mu_{\text{eff}}^2 + \lambda^2 v_u^2 - \frac{g_1^2 + g_2^2}{4} (v_u^2 - v_d^2) \right) - v_u \mu_{\text{eff}} (A_\lambda + \kappa s) = 0$$



$$s \left( m_S^2 + \kappa A_\kappa s + 2\kappa^2 s^2 + \lambda^2 (v_u^2 + v_d^2) - 2\lambda \kappa v_u v_d \right) - \lambda v_u v_d A_\lambda = 0$$



Various choices for different scenarios



# NMSSM Higgs sector

$$D^a |_{\text{Higgs}} = -g \left[ (H_u^*)^\alpha (\tau^a)_\alpha^\beta (H_u)_\beta + (H_d^*)^\alpha (\tau^a)_\alpha^\beta (H_d)_\beta + |S|^2 \right],$$

$$D' |_{\text{Higgs}} = -\frac{g'}{2} \left( |H_u^+|^2 + |H_u^0|^2 - |H_d^0|^2 - |H_d^-|^2 \right)$$

$$\begin{aligned} V = & (|\mu + \lambda S|^2 + m_{H_u}^2) \left( |H_u^+|^2 + |H_u^0|^2 \right) + (|\mu + \lambda S|^2 + m_{H_d}^2) \left( |H_d^-|^2 + |H_d^0|^2 \right) \\ & + \frac{g^2}{2} |H_u^+ H_d^{0*} + H_u^0 H_d^{-*}|^2 + \frac{g^2 + g'^2}{8} \left( |H_u^+|^2 + |H_u^0|^2 - |H_d^0|^2 - |H_d^-|^2 \right)^2 \\ & + m_S^2 |S|^2 + |\kappa S^2 + \lambda (H_u^+ H_d^- - H_u^0 H_d^0)|^2 + \left[ (b + \lambda A_\lambda S) (H_u^+ H_d^- - H_u^0 H_d^0) + \frac{1}{3} \kappa A_\kappa S^3 + h.c. \right] \end{aligned}$$

Expanding the Higgs fields around the VEVs

$$H_u = \begin{pmatrix} H_u^+ \\ H_u^0 \end{pmatrix} \rightarrow \begin{pmatrix} 0 \\ v_u / \sqrt{2} \end{pmatrix} + \begin{pmatrix} \text{Re} H_u^+ \\ \text{Im} H_u^+ \end{pmatrix}$$

$$H_d = \begin{pmatrix} H_d^0 \\ H_d^- \end{pmatrix} \rightarrow \begin{pmatrix} v_u / \sqrt{2} \\ 0 \end{pmatrix} + \begin{pmatrix} \text{Re} H_d^0 \\ \text{Im} H_d^0 \end{pmatrix}$$

$$S \rightarrow v_u / \sqrt{2} + \text{Re} S + i \text{Im} S$$

Higgs mass eigenstates

$$\begin{pmatrix} \text{Re} H_d^0 \\ \text{Re} H_u^0 \\ \text{Re} S \end{pmatrix} \xrightarrow{\alpha} \begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix}, \quad \begin{pmatrix} \text{Im} H_d^0 \\ \text{Im} H_u^0 \\ \text{Im} S \end{pmatrix} \xrightarrow{\beta} \begin{pmatrix} a_1 \\ \text{N.G.B.} \\ a_2 \end{pmatrix}, \quad \begin{pmatrix} H_u^+ \\ H_d^{-*} = H_d^+ \end{pmatrix} \xrightarrow{\beta} \begin{pmatrix} \text{N.G.B.} \\ H^+ \end{pmatrix}.$$

# NMSSM Higgs Sector

The Higgs sector of the NMSSM is described by the six parameters

$$\lambda, \kappa, A_\lambda, A_\kappa, \tan\beta = v_u/v_d, \mu_{\text{eff}}.$$

The couplings of the Higgs states depend on their decompositions into the CP-even weak eigenstates  $H_d$ ,  $H_u$  and  $S$ , which are given by

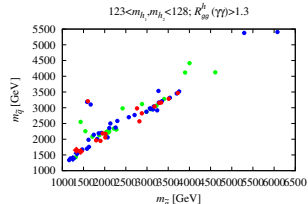
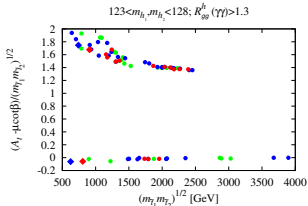
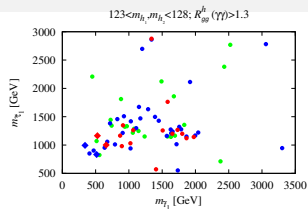
$$\begin{aligned} h_1 &= S_{1,d} H_d + S_{1,u} H_u + S_{1,s} S, \\ h_2 &= S_{2,d} H_d + S_{2,u} H_u + S_{2,s} S. \end{aligned}$$

Then, the reduced couplings of  $h_i$  are

$$C_{dd}^{h_i} = \frac{S_{i,d}}{\cos\beta}, \quad C_{uu}^{h_i} = \frac{S_{i,u}}{\sin\beta}, \quad C_{VV}^{h_i} = \cos\beta S_{i,d} + \sin\beta S_{i,u}.$$

The loop-induced reduced couplings  $C_{gg}^{h_i}$  and  $C_{\gamma\gamma}^{h_i}$  have to be computed including contributions from SUSY particles in the loops, including scalar  $\tau$ -leptons, charginos and more.

# Implication for SUSY Particles

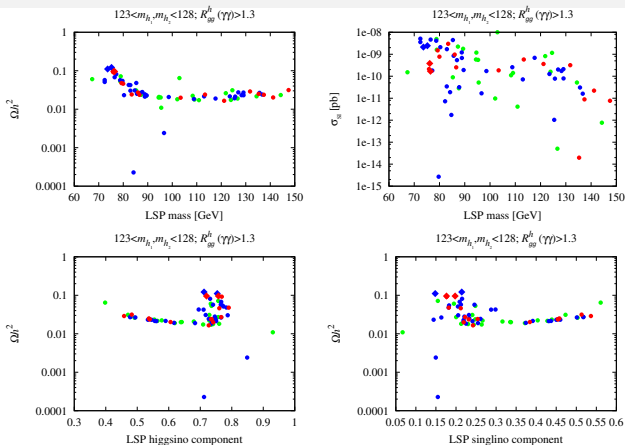


- Indeed, the few points which we found in the WMAP window always have  $m_{\tilde{t}_1} < 700$  GeV.

- A good fraction of our points with degenerate  $h_1, h_2$  and  $R(\gamma\gamma) > 1.3$  features light stops with  $M_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} \lesssim 1$  TeV. The stop mixing is typically large in these cases,  $(A_t - \mu_{\text{eff}} \cot \beta) / M_{\text{SUSY}} \approx 1.5$ – $2$ .

- Squark and gluino masses are above about 1.25 TeV ranging up to as high as 6 TeV (where our scanning more or less ended). The WMAP-window points with large  $R_{gg}^h(\gamma\gamma)$  are located at low masses of  $m_{\tilde{g}} \sim 1.3$  TeV and  $m_{\tilde{q}} \sim 1.6$  TeV.

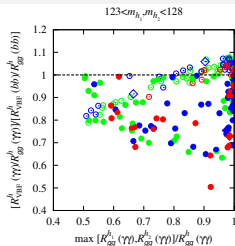
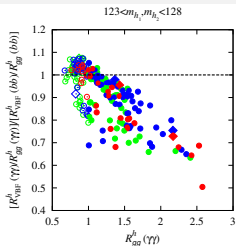
# Dark Matter Properties



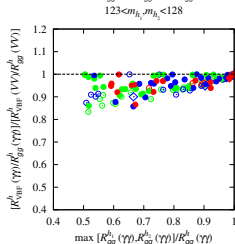
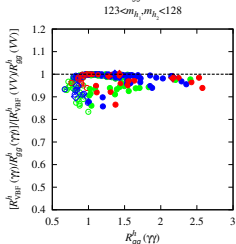
- WMAP-window points have a rather limited range of LSP masses, roughly  $m_{\tilde{\chi}_1^0} \in [60, 80]$  GeV.
- Corresponding  $\sigma_{\text{SI}}$  values range from *few*  $\times 10^{-9}$  pb to as low as *few*  $\times 10^{-11}$  pb.
- **large  $\Omega h^2$ :** mixed higgsino–singlino, with a singlino component of the order of 20%.
- **low  $\Omega h^2$ :** the LSP is **dominantly higgsino** (owing to small  $\mu_{\text{eff}}$ ).

# double ratio examination – NMSSM (ii)

I



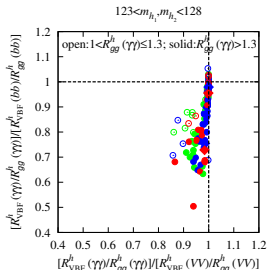
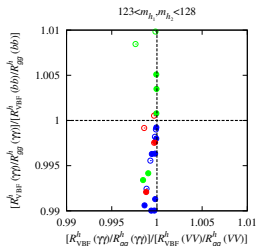
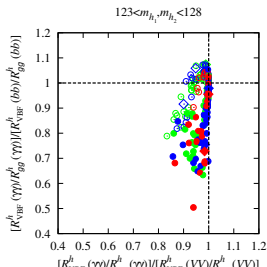
II



- For the NMSSM, it seems that **double ratio I**) provides the greatest discrimination for the majority of the degenerate scenarios that have **enhanced  $\gamma\gamma$  rates**.
- It particularly, being sensitive to the  $b\bar{b}$  final state, singles out degenerate Higgs scenarios even when one or the other of  $h_1$  or  $h_2$  dominates the  $gg \rightarrow \gamma\gamma$  rate.
- In comparison, **double ratio II**) is most useful for scenarios with  $R_{gg}^h(\gamma\gamma) \sim 1$ .

## double ratio examination – NMSSM (iii)

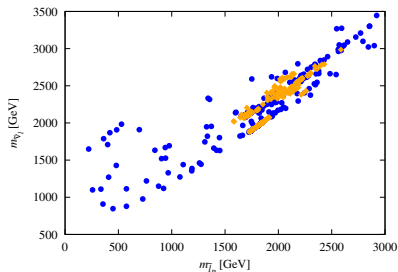
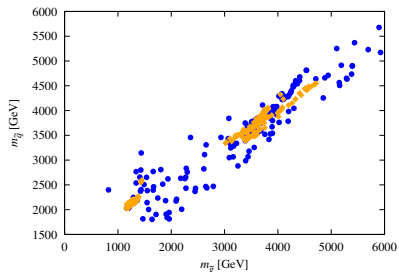
As illustrated below, the greatest discriminating power is clearly obtained by measuring both double ratios.



In fact, a close examination reveals that there are no points for which *both* double ratios are exactly 1!

*Of course, experimental errors may lead to a region containing a certain number of points in which both double ratios are merely consistent with 1 within the errors.*

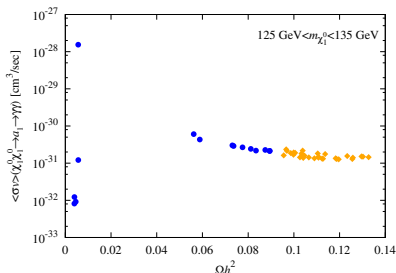
# Prediction for SUSY Particles



- The predicted  $m_{\tilde{q}}$  and  $m_{\tilde{g}}$  are beyond current experimental limits, although the lowest values (as found in particular in the low- $m_{\tilde{\chi}_1^0}$  WMAP-window scenarios) may soon be probed.
- Note that  $m_{\tilde{g}}$  can be below  $m_{\tilde{\ell}_R}$  (as common in constrained models when  $m_0$  is large) for some points, including the low- $m_{\tilde{\chi}_1^0}$  points in the WMAP window.

## Able to describe the Fermi-LAT data?

Whether or not any of the points are such as to describe the monochromatic signal  $\langle\sigma v\rangle(\tilde{\chi}_1^0\tilde{\chi}_1^0\rightarrow\gamma\gamma)\sim 10^{-27}\text{cm}^3/\text{sec}$  at 130 GeV observed in the Fermi-LAT data?



For points with  $m_{\tilde{\chi}_1^0} \in [125, 135]$  GeV, it is the **s-channel  $a_1$  diagram** that can give a large  $\langle\sigma v\rangle$ .

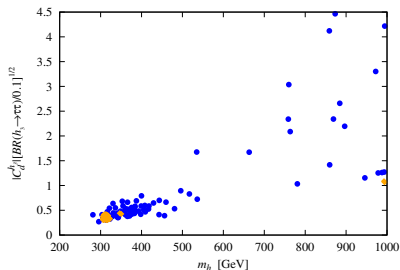
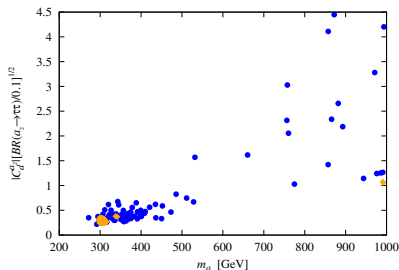
- In the WMAP window, they have values of  $\langle\sigma v\rangle$  four orders of magnitude below that required to explain the excess.
- Those points with the largest  $\langle\sigma v\rangle$  always have quite small  $\Omega h^2$  and hence  $\rho_{DM}$ .
- Incidentally, all the points in our plots are fully consistent with the current bounds from the continuum  $\gamma$  spectrum as measured by Fermi-LAT.

If the 130 GeV gamma ray line is confirmed, then a fully general NMSSM model (no GUT scale unifications) might be simultaneously consistent with all experiment measurements.



# LHC Test – Direct Higgs Production and Decay

A final possible detection mode is  $gg \rightarrow a_2, h_3 \rightarrow \tau^+ \tau^-$ .



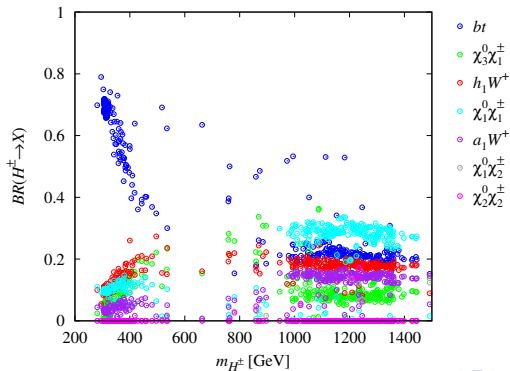
in the effective down-quark coupling, 0.1 is a reference value of  $BR(H, A \rightarrow \tau^+ \tau^-)$  implicit in the MSSM limit plots discussed below.

- Noting that  $m_{a_2} \simeq m_{h_3}$  and the fact that the two plots are nearly identical shows that we may sum the  $a_2$  and  $h_3$  signals together in the same manner as the  $H$  and  $A$  signals are summed together in the case of the analogous plot of  $\tan \beta$  vs.  $m_A \simeq m_H$  in the case of the MSSM.
- Limits from CMS  $4.6 \text{ fb}^{-1}$  data are of order  $C_d^{a_2, h_3}(\text{eff}) \lesssim 7 - 8$  for  $m_{a_2} \simeq m_{h_3} \in [150, 220] \text{ GeV}$  rising rapidly to reach  $\sim 50$  at degenerate mass of order 500 GeV.

## LHC Test – Charged Higgs Decay

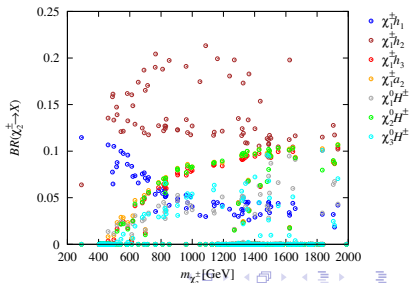
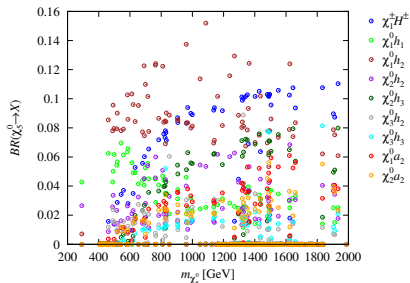
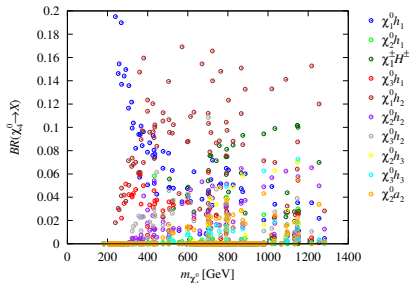
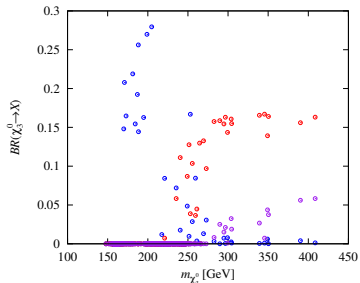
Regarding the  $H^\pm$ , prospects for its discovery at masses for which  $H^+H^-$  production has substantial cross section appear to be promising in the  $bt$  final state provided reconstruction of the  $bt$  mass is possible with good efficiency and one or more  $b$  tags are sufficient to reject SM background.

Also very interesting would be detection of  $H^\pm \rightarrow h_1 W^\pm$  in the  $h_1 \rightarrow b\bar{b}$  final state using mass reconstruction for the  $b\bar{b}$  and a leptonic trigger from the  $W^\pm$  to reject backgrounds.



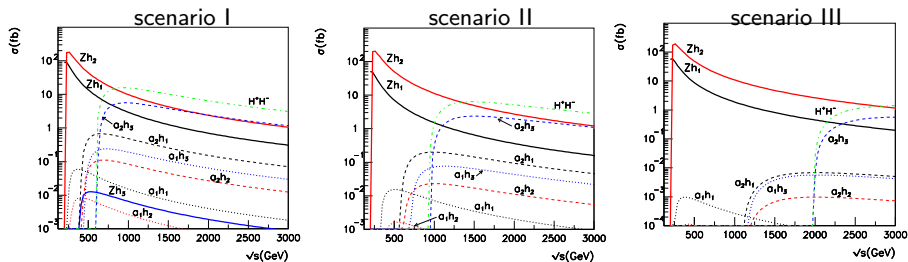
# LHC Test – Higgses from Neutralino Decays

Whether some of the Higgs bosons can be detected via ino-pair production?



# Linear Collider Test

An  $e^+e^-$  collider would be the ideal machine to produce the additional Higgs states and resolve the scenario.



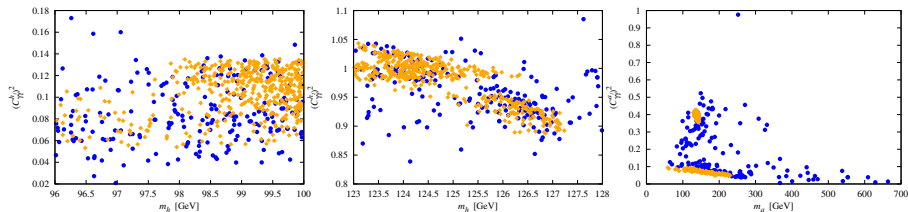
Scen.	$m_{h_1}$	$m_{h_2}$	$m_{h_3}$	$m_{a_1}$	$m_{a_2}$	$m_{H^\pm}$	$m_{\tilde{\chi}_1^0}$	$\Omega h^2$	LSP-S	LSP- $\tilde{H}$	$R_{\tilde{g}\tilde{g}}^{h_2}(\gamma\gamma)$
I	99	124	311	140	302	295	76	0.099	18%	75%	1.62
II	97	124	481	217	473	466	92	0.026	20%	74%	1.53
III	99	126	993	147	991	989	115	0.099	75%	25%	1.14

With an integrated luminosity of  $1000 \text{ fb}^{-1}$ , substantial event rates for many  $Z$ +Higgs and Higgs pair final states are predicted. Of course,  $Zh_1$  and  $Zh_2$  production have the largest cross sections and lowest thresholds. The next lowest thresholds are for  $a_1 h_1$  production, but the cross sections are quite small.

In the  $e^+e^-$  collider case, it would be easy to isolate signals in many final states.

# Photon Collider Test

In the  $\gamma\gamma$  collider, the  $\gamma$ 's are obtained by backscattering of laser photons off the energetic  $e$ 's. A huge range of energies is possible for such a  $\gamma\gamma$  collider, ranging from low to high center of mass energies depending upon the center of mass energy of the underlying electron collider.



- The fairly SM-like  $h_2$  at  $\sim 125$  GeV can be studied easily at such a collider since its  $\gamma\gamma$  coupling is close to SM strength.
- Even though the  $h_1$  and  $a_1$  are largely singlet, both have  $\gamma\gamma$  couplings-squared that are often of order  $0.1 \times \text{SM}$  and above (at the same mass). Indeed, this coupling becomes stronger as  $\lambda$  is increased.

# Muon Collider Test

A muon-collider with  $\sqrt{s}$  close to the Higgs mass in question would be a particularly ideal machine to study any Higgs boson with  $\mu^+\mu^-$  coupling that is not too different from that of a SM Higgs boson of similar mass.

